

Dirac Equation in Cosmological Inertial Frame

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ABSTRACT

Dirac equation is a one order-wave equation. Wave function uses as a probability amplitude in quantum mechanics. We make Dirac Equation from wave function, Type A in cosmological inertial frame. The Dirac equation satisfy Klein-Gordon equation in cosmological inertial frame.

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1. Introduction

Dirac equation is in special relativity theory,

$$(i\hbar\gamma^\mu\partial_\mu - mcI)\psi = 0 \quad ,$$

I is 4×4 unit matrix ,

$$\gamma^0 = \begin{pmatrix} 0 & I' \\ I' & 0 \end{pmatrix}, \quad \gamma^i = \begin{pmatrix} 0 & -\sigma^i \\ \sigma^i & 0 \end{pmatrix}, \quad \sigma^1 = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}, \quad \sigma^2 = \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix}, \quad \sigma^3 = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}$$

I' is 2×2 unit matrix, σ^i is Pauli's matrix. (1)

2. Dirac Equation from Wave Function-Type A in Cosmological Inertial Frame

Dirac equation is the wave equation. Therefore, Dirac equation is in cosmological inertial frame,[2]

Wave function Type A:

$$r \rightarrow r\sqrt{\Omega(t_0)} \quad , \quad t \rightarrow \frac{t}{\sqrt{\Omega(t_0)}} \quad ,$$

t_0 is the cosmological time. $\Omega(t_0)$ is the expanding ratio of universe in the cosmological time t_0 .

$$(i\hbar\sqrt{\Omega(t_0)}\gamma^0\partial_0 + i\hbar\frac{1}{\sqrt{\Omega(t_0)}}\gamma^i\partial_i - mcI)\phi = 0 \quad (2)$$

If $\bar{\partial}_\mu$ is

$$\bar{\partial}_\mu = (\sqrt{\Omega(t_0)}\partial_0, \frac{1}{\sqrt{\Omega(t_0)}}\partial_i) \quad (3)$$

Dirac equation is in cosmological inertial frame,

$$(i\hbar\gamma^\mu\bar{\partial}_\mu - mcI)\phi = 0 \quad (4)$$

Eq(4) multiply $i\hbar\gamma^\nu\bar{\partial}_\nu$, hence

$$(-\hbar^2(\gamma^\mu\bar{\partial}_\mu)(\gamma^\nu\bar{\partial}_\nu) - i\hbar(\gamma^\nu\bar{\partial}_\nu)mcI)\phi = 0 \quad (5)$$

In this time,

$$i\hbar\gamma^\nu\bar{\partial}_\nu\phi = mcI\phi \quad (6)$$

Hence, Eq(5) is

$$(-\hbar^2\gamma^\mu\gamma^\nu\bar{\partial}_\mu\bar{\partial}_\nu - m^2c^2I)\phi = 0 \quad (7)$$

In this time, matrix γ^μ is

$$\frac{1}{2}(\gamma^\mu \gamma^\nu + \gamma^\nu \gamma^\mu) = \frac{1}{2}\{\gamma^\mu, \gamma^\nu\} = \eta^{\mu\nu} I \quad (8)$$

Therefore,[1],[3]

$$\begin{aligned} & \frac{1}{2}(\gamma^\mu \gamma^\nu + \gamma^\nu \gamma^\mu) \bar{\partial}_\mu \bar{\partial}_\nu \phi + \frac{m^2 c^2}{\hbar^2} I \phi \\ & = (\eta^{\mu\nu} \bar{\partial}_\mu \bar{\partial}_\nu + \frac{m^2 c^2}{\hbar^2}) I \phi = 0 \end{aligned} \quad (9)$$

Rq(9) is the matrix equation of Klein-Gordon.

Dirac spinor ϕ is $\phi = (\phi_1, \phi_2, \phi_3, \phi_4)$. ϕ 's hermitian conjugate $\phi^+ = (\phi_1^*, \phi_2^*, \phi_3^*, \phi_4^*)$.

Hence, ϕ 's adjoint spinor $\bar{\phi}$ is

$$\bar{\phi} = \phi^+ \gamma^0, \quad \bar{\phi} (i\gamma^\mu \bar{\partial}_\mu + mcI) = 0 \quad (10)$$

Hence, positive probability density j^0 is

$$j^0 = \bar{\phi} \gamma^0 \phi = \phi^+ \phi = |\phi_1|^2 + |\phi_2|^2 + |\phi_3|^2 + |\phi_4|^2 \quad (11)$$

3. Conclusion

We found Dirac equation from Wave Function-Type A in cosmological special theory of relativity. The wave function uses as a probability amplitude.

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