# Aerosol Transport by Turbulent Continua 

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#### Abstract

The stochastic transport equations, derived rigorously under the condition of continnum fluctuations in the framework of an ensemble theory, both in differential and integral form, are then verified by establishing an unambiguous connection between this stochastics and the associated deterministics.


Keywords - transport equations, turbulence, stochastics and deterministics, natural causality

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## Chapter 1

## Introduction

Aerosol transport through turbulent continua is characterized by the fact that aerosols can only follow movements of fluid elements if they neither fall below nor exceed a certain size and weight. This is in contrast to molecular diffusion through matter, which is often successfully accomplished by known diffusion equations. This physical process is fundamentally different from turbulent aerosol transport. In the former case, the diffusing molecules have an intrinsic motion between two interactions, while in the latter case, only predetermined paths are followed.

These predetermined paths must correspond to a continuum system in which fluid elements follow the collective motions of many individual molecules moving locally apparently independently. I.e. fluid elements, which in their totality represent a fluctuating continuum, and their paths are abstract quantities and not points of matter.

First, a fluid and turbulent fluid continuum is defined. According to this, a purely stochastic aerosol transport is excluded and a stochastic ensemble consideration is developed.

The stochastic transport equations, derived rigorously under the condition of continnum fluctuations in the framework of an ensemble theory, both in differential and integral form, are then verified by establishing an unambiguous connection between this stochastics and the associated deterministics.

## Chapter 2

## Definition of a Moved Fluid

### 2.1 Definition of a fluid element

At every time, space points $(\overrightarrow{\boldsymbol{x}})$ are assigned to fluid elements in a unique correspondence. As this applies to every space point $(\overrightarrow{\mathbf{x}})$ of the fluid field, the set of fluid elements is seen as a continuum. A Continuum of fluid element points (simply called fluid elements) is considered, where a fluid environment of non infinitesimal size is uniquely allocated to every fluid element point. Two infinitesimally neighboring fluid elements differ apart from their distance by their velocities and not quite identical material distributions of their neighborhoods. The neighborhoods of two nearby fluid elements overlap. A fluid element is shifted moving the material of its neighborhood. Though the material of such a fluid element may have changed marginally after an infinitesimal time interval $t_{\varepsilon}$, it can be identified principally by its prior material status. As every molecule possesses its own identity, there has to be at least an infinitesimally greater difference of material distribution to the neighborhoods of other fluid elements.
The neighborhoods exchange material with neighborhoods of adjacent fluid elements and vary their thermodynamic state (a local thermodynamic state does not necessarily exist). Their size is not infinitesimal, because a local thermodynamic state (if physically existent) has to be detectable at least in thought experiment. The open neighborhoods have equally sized spherical shapes, generally. Near a solid border they are descibed by parts of spheres. Infinitesimally adjacent fluid elements possess overlapping neighborhoods. In an $\varepsilon$-surrounding they move in parallel. So one obtains a fluid, which is assumed to be a dense fluctuating point set, though there is no continuous matter distribution in Space-Time. That means it is possible to follow theoretically the history of every fluid element, though it has exchanged a lot of its initial material altering its local thermodynamic state.

The fluid is an abstract, dense set of fluctuating fluid elements, which do not generally correspond to material points. A continuum of moved fluid elements is considered each uniquely assigned to a neighborhood and a velocity.

$$
\begin{equation*}
\overrightarrow{\mathbf{v}}_{t_{\epsilon}}=\frac{\overrightarrow{\mathbf{x}}_{2}-\overrightarrow{\mathbf{x}}_{1}}{t_{\epsilon}} \tag{2.1}
\end{equation*}
$$

The fluidelement first determined in space point $\overrightarrow{\mathbf{x}}_{1}$ and $t_{\epsilon}$-time later detected at $\overrightarrow{\mathbf{x}}_{2}$ is identified having at time $t_{0}+t_{\epsilon}$ in $\overrightarrow{\mathbf{x}}_{2}$ in comparison to all other points $\overrightarrow{\mathbf{x}}$ the most similar material to that of $\overrightarrow{\mathbf{x}}_{1}$ in $t_{0}$. In this connection it is remarqued, that parts of the individual aerosols or molecules may be identified, too. The accuracies of the considered motion quantities are determined by $t_{\varepsilon}$-measurement processes. $t_{\varepsilon}$ characterising the accuracy. According to a process $\lim t_{\epsilon} \rightarrow 0$, the fluid elements move along trajectories that can a sufficient number of times be continuously differentiated forming a continuum as a whole. This continuum has a velocity vector field with $\boldsymbol{\operatorname { r o t }}(\overrightarrow{\mathbf{v}}) \neq \mathbf{0}$ generally. ${ }^{1}$ Though $\boldsymbol{\operatorname { r o t }}(\overrightarrow{\mathbf{v}})$ has dimension $[\mathbf{1} / \mathbf{s e c}]$ in the laminar case it does not refer to a rotation.

### 2.2 The Orthogonality of $\operatorname{rot}(\overrightarrow{\mathrm{v}})$ and $\vec{v}$ is a Consequence of the Moved Fluid Continuum

A fluid continuum is characterized by

[^0]1. continuously differentiable velocities
2. parallel velocities in an $\epsilon$ - surrounding of a space point $\overrightarrow{\mathbf{x}}$

Considering without loss of generality a fluid movement of velocity $\overrightarrow{\mathbf{v}}\left(\overrightarrow{\mathbf{x}}_{0}\right)=\left(v_{x}, 0,0\right)$ in a space point $\overrightarrow{\mathbf{x}}_{0}$ in cartesian coordinates, the velocity is described in an $\epsilon$-neighborhood and parallel to the x-coordinate as follows:

$$
\overrightarrow{\mathbf{v}}(\overrightarrow{\mathbf{x}})=\left(\begin{array}{c}
\mathbf{v}_{x}\left(\overrightarrow{\mathbf{x}_{\mathbf{0}}}\right) \\
\mathbf{0} \\
\mathbf{0}
\end{array}\right)+\left(\begin{array}{c}
\left.\frac{\partial \mathbf{v}_{x}}{\partial x}\right|_{\overrightarrow{\vec{x}}_{0}}+\left.\frac{\partial \mathbf{v}_{x}}{\partial y}\right|_{\overrightarrow{\mathbf{x}}_{0}}+\left.\frac{\partial \mathbf{v}_{x}}{\partial z}\right|_{\overrightarrow{\mathbf{x}}_{0}} \\
\left.\frac{\partial \mathbf{v}_{y}}{\partial x}\right|_{\overrightarrow{\mathbf{x}}_{0}}+\left.\frac{\partial \mathbf{v}_{y}}{\partial y}\right|_{\overrightarrow{\mathbf{x}}_{0}}+\left.\frac{\partial \mathbf{v}_{y}}{\partial z}\right|_{\overrightarrow{\mathbf{x}}_{0}} \\
\left.\frac{\partial \mathbf{v}_{z}}{\partial x}\right|_{\overrightarrow{\mathbf{x}}_{0}}+\left.\frac{\partial \mathbf{v}_{z}}{\partial y}\right|_{\overrightarrow{\mathbf{x}}_{0}}+\left.\frac{\partial \mathbf{v}_{z}}{\partial z}\right|_{\overrightarrow{\mathbf{x}}_{0}}
\end{array}\right)\left(\begin{array}{c}
\Delta x \\
\Delta y \\
\Delta z
\end{array}\right) \ldots
$$

The velocity components $\mathbf{v}_{y}(\overrightarrow{\mathbf{x}})$ and $\mathbf{v}_{z}(\overrightarrow{\mathbf{x}})$ osculate at the velocity $\overrightarrow{\mathbf{v}}\left(\overrightarrow{\mathbf{x}}_{0}\right)=\left(v_{x}, 0,0\right)$ spatially approaching (constant time $t_{0}$ ),

$$
\begin{aligned}
& \mathbf{v}_{y}\left(x_{0}, y, z_{0}\right) \longrightarrow \mathbf{v}_{y}\left(x_{0}, y_{0}, z_{0}\right)=\mathbf{0} \\
& \mathbf{v}_{z}\left(x_{0}, y_{0}, z\right) \longrightarrow \mathbf{v}_{z}\left(x_{0}, y_{0}, z_{0}\right)=\mathbf{0} .
\end{aligned}
$$

That means especially, that all the partial derivations by $\mathbf{y}$ - or $\mathbf{z}$-coordinate of 1 . order of $\mathbf{v}_{y}(\overrightarrow{\mathbf{x}})$ and $\mathbf{v}_{z}(\overrightarrow{\mathbf{x}})$ disappear in the point $\left(x_{0}, y_{0}, z_{0}\right)$.

$$
\begin{align*}
\left.\lim _{z \rightarrow \boldsymbol{z}_{0}} \frac{\Delta \mathbf{v}_{y}}{\Delta z}\right|_{\overrightarrow{\mathbf{x}}_{0}} & =\left.\lim _{\boldsymbol{y} \rightarrow \boldsymbol{y}_{0}} \frac{\Delta \mathbf{v}_{z}}{\Delta y}\right|_{\overrightarrow{\mathbf{x}}_{0}}=\mathbf{0} .  \tag{2.2}\\
\overrightarrow{\mathbf{x}}_{0} & =\left(x_{0}, y_{0}, z_{0}\right)
\end{align*}
$$

Applying the differential quotients in the $\overrightarrow{\boldsymbol{\nabla}} \times$-operator expresssed in cartesian coordinates gives for the fluid velocity

$$
(\overrightarrow{\boldsymbol{\nabla}} \times \overrightarrow{\mathbf{v}})_{\left.\right|_{\overrightarrow{\mathrm{x}}_{0}}}=\left(\begin{array}{c}
0  \tag{2.3}\\
\frac{\partial v_{x}}{\partial z_{y}}-\frac{\partial v_{z}}{\partial x} \\
\frac{\partial v_{y}}{\partial x}-\frac{\partial v_{x}}{\partial y}
\end{array}\right)_{\left.\right|_{\overrightarrow{\mathbf{x}}_{0}}}, \overrightarrow{\mathbf{v}}\left(\overrightarrow{\mathbf{x}}_{0}\right)=\left(v_{x}, 0,0\right)
$$

## $\Longrightarrow$

The orthogonality of $\vec{\nabla} \times \overrightarrow{\mathbf{v}} \perp \overrightarrow{\mathbf{v}}$ is a fundamental quality ${ }^{23}$ and a necessary condition for continuous fluid flow.

In this orthogonality velocity vector fields differ from deformation vector fields.

[^1]
## Chapter 3

## Definition of a Turbulent Fluid

Trying to identify the state of movement of a fluid element in turbulent fluids by a velocity $\vec{v}_{t_{\epsilon}}$ it should be recognized, that the state of movement is not yet determined, as the path in every space point (except in turning points) is uniquely adapted by an infinitesimal circle segment. In the infinitesimal neighborhood of a path point the velocity is identified by an instantaneous axis of rotation $\overrightarrow{\boldsymbol{\omega}}_{t_{\epsilon}}$ and a radius vector $\overrightarrow{\mathbf{r}}_{t_{\epsilon}} \cdot{ }^{1}$

$$
\begin{equation*}
\overrightarrow{\mathbf{v}}_{t_{\epsilon}}=\vec{\omega}_{t_{\epsilon}} \times \overrightarrow{\mathbf{r}}_{t_{\epsilon}} \tag{3.1}
\end{equation*}
$$

In a turbulently moved fluid the fluid elements move on curved trajectories in some space time points having turning points with $\overrightarrow{\boldsymbol{\omega}}_{\boldsymbol{t}_{\epsilon}}=\mathbf{0}$ and a curvature vector $\overrightarrow{\mathbf{b}}_{\boldsymbol{t}_{\epsilon}}=\mathbf{0}$. The considered vectorial motion quantities $\overrightarrow{\boldsymbol{\omega}}_{\boldsymbol{t}_{\epsilon}}$ and $\overrightarrow{\mathbf{r}}_{\boldsymbol{t}_{\epsilon}}$ are determined by $t_{\varepsilon}$-measurement processes, which are calculated later on by a limes process $\lim \boldsymbol{t}_{\boldsymbol{\epsilon}} \rightarrow 0$. A fluid element originating from the point $\overrightarrow{\mathbf{x}}_{0}$ crossing $\overrightarrow{\mathbf{x}}_{1}$ after the time $t_{\varepsilon}$ reaches $\overrightarrow{\mathbf{x}}_{2}$ after a further time $t_{\varepsilon}$.

$$
\overrightarrow{\mathbf{x}}_{0} \xrightarrow{\mathbf{t}_{\epsilon}} \overrightarrow{\mathbf{x}}_{1} \xrightarrow{\mathbf{t}_{\epsilon}} \overrightarrow{\mathbf{x}}_{2}
$$

A segment of a circle is clearly drawn through these 3 points with radius vector $\overrightarrow{\mathbf{r}}_{t_{\varepsilon}}$ and velocity of rotation $\overrightarrow{\boldsymbol{\omega}}_{t_{\varepsilon}}$ in $\overrightarrow{\mathbf{x}}_{1}$, unless a turning point is passed through. The local state of motion can not be described by velocity only, neither statistically nor deterministically. ${ }^{2}$
Thus the fluid element in the space-time-point $(\vec{x}, t)$ is identified principally by the contents of the matter of its neighborhood and state of movement expressed by $\overrightarrow{\boldsymbol{\omega}}_{t_{\epsilon}}$ and $\overrightarrow{\mathbf{r}}_{t_{\epsilon}}$. In that way defined fluid elements move on sufficiently often continuously differentiable trajectories. At each instant they lead to a new continuum of fluctuating fluid elements with several times continuously differentiable velocity field. The continuum of moving fluid elements represents the turbulent collective motion of a discontinuously spaced Matter. This is the result of the connection between deterministics and stochastics in the sense of an ensemble theory, which is presented in the following. The field of turbulence is described by the two vector fields $\overrightarrow{\boldsymbol{\omega}}_{t_{\epsilon}}$ and $\overrightarrow{\mathbf{b}}_{t_{\epsilon}}$,

$$
\begin{equation*}
\overrightarrow{\mathbf{b}}_{t_{\epsilon}}=\overrightarrow{\mathbf{r}}_{t_{\epsilon}} / r_{t_{\epsilon}}^{2} \quad \text {-curvature vector field. } \tag{3.2}
\end{equation*}
$$

In addition, the results show that

$$
\begin{equation*}
\overrightarrow{\boldsymbol{\omega}}_{t_{\epsilon}}=\frac{1}{2} \operatorname{rot}\left(\overrightarrow{\mathbf{v}}_{t_{\epsilon}}\right) \tag{3.3}
\end{equation*}
$$

$\operatorname{rot}(\overrightarrow{\mathbf{v}})$ has the meaning of a local rotation in the frame of turbulence. An infinitesimal disturbance of stationary pipe flow leads to an change of the significance $\operatorname{of} \operatorname{rot}(\overrightarrow{\mathbf{v}})$, where $\operatorname{rot}(\overrightarrow{\mathbf{v}})$ does not correspond to a rotation initially. Whether starting motions of turbulence are suppressed, depends on an existent viscosity. These decelerations are generally weak. The beginning of turbulent movements avoid Newtonian friction as well as pressure gradients by means of hereto orthogonal motions.

Vortex fields in turbulence (local rotation fields will be identified with vortex fields) and radius fields may have turning points along the paths of the fluid elements, which means $\overrightarrow{\boldsymbol{\omega}}=\mathbf{0}$ and $\overrightarrow{\boldsymbol{r}}=\infty$. ${ }^{3}$ In this case the velocities are to be calculated by interpolation or extrapolation from the neighborhood. The fluid elements are accompanied by a moving frame of $\overrightarrow{\boldsymbol{\omega}}, \overrightarrow{\boldsymbol{b}}$ and $\overrightarrow{\boldsymbol{v}}$ along their paths.

[^2]Fluid elements, at a time are infinitesimally adjacent, have later moved away from each other and represent with new neighbors a new continuum. However, since also their material environments have changed, their past and future stay is to be determined only from the knowledge of the perfect spatiotemporal movement field. To calculate these fields, a system of equations is needed that couples other independent fields, such as the acceleration field.
Independent Lagrangian turbulence calculations are not possible.

## Chapter 4

## Definition of Markov Processes with Natural Causality

The probabilistic theory is related to random distributions of velocities $\overrightarrow{\boldsymbol{\pi}}$ moving from $(\overrightarrow{\mathbf{x}}, t)$ to $\left(\overrightarrow{\mathbf{x}}+\overrightarrow{\boldsymbol{\pi}} t_{\epsilon}, t+t_{\epsilon}\right)$. These velocity distributions may get together of vortex and curvature vector fields

$$
\overrightarrow{\boldsymbol{\pi}}=\overrightarrow{\boldsymbol{\omega}} \times \frac{\overrightarrow{\boldsymbol{b}}}{b^{2}}
$$

The transport from $\left(\overrightarrow{\mathbf{x}}-t_{\varepsilon} \overrightarrow{\boldsymbol{\pi}}^{\prime}, \boldsymbol{t}-\boldsymbol{t}_{\boldsymbol{\epsilon}}\right)$ to $(\overrightarrow{\mathbf{x}}, t)$ is addionally controlled by transition probabilities

$$
W_{\boldsymbol{t}_{\epsilon}}=W_{\boldsymbol{t}_{\epsilon}}\left(\overrightarrow{\mathrm{x}}, t, \overrightarrow{\boldsymbol{\pi}}, \overrightarrow{\boldsymbol{\pi}}^{\prime}\right)
$$

resulting in

$$
f_{t_{\epsilon}}(\overrightarrow{\mathbf{x}}, t, \vec{\pi})=\int_{\vec{\pi}^{\prime}} W_{\boldsymbol{t}_{\epsilon}}\left(\overrightarrow{\mathrm{x}}, t, \overrightarrow{\boldsymbol{\pi}}, \overrightarrow{\boldsymbol{\pi}}^{\prime}\right) f_{\boldsymbol{t}_{\epsilon}}\left(\overrightarrow{\mathrm{x}}-t_{\varepsilon} \overrightarrow{\boldsymbol{\pi}}^{\prime}, t-t_{\boldsymbol{\epsilon}}, \vec{\pi}^{\prime}\right) d \overrightarrow{\boldsymbol{\pi}}^{\prime}
$$

Such a relation we call a Markov Process of natural causality. According to Sen [6] there is a so called Newtonian causality in nonrelativistic physics implying the possibility of unlimited velocities. However Newtonian causality is restricted to Newtonian mechanics and stochastic processes of physics ending with diffusion equations when applied practically. ${ }^{1}$ This applies not for formulations of the general or linear Boltzmann Equation. In electrodynamics the velocity of light is the limiting velocity. The Newtonian causality proves to be a limiting case of non relativistic classical physics. Subsequently a causal Markov Process is continuously used or derived. Overarching master equations can not exist, physically. The transition probabilities $W_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}$ depend on a time quantity $\boldsymbol{t}_{\boldsymbol{\epsilon}}$ related to continuum fluctuations of measurement accuracy according to vectorial motion quantities. For $\boldsymbol{t}_{\boldsymbol{\epsilon}} \rightarrow \boldsymbol{0}$ (exact motion quantities) the transition probabillity $W_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}$ degenerates to a $\delta$-function.

Simultaneous details of space and momentum are not possible in the context of quantum mechanics. The Schrödinger Equation for free "quantum particles"

$$
\begin{equation*}
i \hbar \frac{\partial \psi(\overrightarrow{\mathbf{x}}, t)}{\partial t}=-\frac{\hbar^{2}}{2 \mu} \vec{\nabla}^{2} \psi(\overrightarrow{\mathbf{x}}, t) \tag{4.1}
\end{equation*}
$$

can be transformed into a linear homogenuous integral eqution [3] [4]

$$
\begin{equation*}
\psi(\overrightarrow{\mathbf{x}}, t)=i \int G\left(\overrightarrow{\mathbf{x}}, t ; \overrightarrow{\mathbf{x}^{\prime}}, t^{\prime}\right) \psi\left(\overrightarrow{\mathbf{x}^{\prime}}, t^{\prime}\right) d \overrightarrow{\mathbf{x}^{\prime}} \tag{4.2}
\end{equation*}
$$

The Green function

$$
\begin{equation*}
G\left(\overrightarrow{\mathbf{x}}, t ; \overrightarrow{\mathbf{x}}^{\prime}, t^{\prime}\right)=\langle\overrightarrow{\mathbf{x}}| \exp \left(-\frac{\mathbf{i}}{\hbar}\left(t-t^{\prime}\right) \mathbf{H}\right)\left|\overrightarrow{\mathbf{x}}^{\prime}\right\rangle \tag{4.3}
\end{equation*}
$$

[^3]is called Feynman kernel, too.
In the case of the diffusion equation
\[

$$
\begin{equation*}
\frac{\partial \rho(\overrightarrow{\mathbf{x}}, t)}{\partial t}=D \vec{\nabla}^{2} \rho(\overrightarrow{\mathbf{x}}, t) \tag{4.4}
\end{equation*}
$$

\]

an equivalent integral equation the Green function understood as transition probabillity from ( $\overrightarrow{\mathbf{x}}^{\prime}, t^{\prime}$ ) to ( $\overrightarrow{\mathbf{x}}, t$ ) exists with

$$
\begin{equation*}
\rho(\overrightarrow{\mathbf{x}}, t)=\int_{V^{\prime}} G\left(\overrightarrow{\mathbf{x}}, t ; \overrightarrow{\mathbf{x}}^{\prime}, t^{\prime}\right) \rho\left(\overrightarrow{\mathbf{x}}^{\prime}, t^{\prime}\right) d \overrightarrow{\mathbf{x}}^{\prime} \tag{4.5}
\end{equation*}
$$

and the Green function

$$
\begin{equation*}
G\left(\overrightarrow{\mathbf{x}}, t ; \overrightarrow{\mathbf{x}}^{\prime}, t^{\prime}\right)=\left(\frac{1}{4 \pi D\left(t-t^{\prime}\right)}\right)^{\frac{3}{2}} e^{-\frac{\left(\vec{x}-\vec{x}^{\prime}\right)^{2}}{4 \pi D\left(t-t^{\prime}\right)}} \tag{4.6}
\end{equation*}
$$

Equations based on a 'heat-kernel"-structure are not exact in classical physics (as well as the Newtonian mechanics). They are usually referred to as Markov processes.

In quantum mechanics and quantum field theory natural causality is not possible because of the uncertainty principle. In Relativity there is the maximal possible velocity, the velocity of light.

## Chapter 5

## Stochastic Transport of Aerosols by turbulent Continuum-Fluctuations

### 5.1 Introduction

The motion of passive aerosols by turbulent continuum fluctuations is examined. The aerosoles are moved not affecting this field. Their trajectories correspond in every $\varepsilon$-neiborhood of a point to a circle segment passed with the velocity

$$
\begin{equation*}
\overrightarrow{\mathrm{v}}_{\boldsymbol{t}_{\epsilon}}=\vec{\omega}_{\boldsymbol{t}_{\epsilon}} \times \overrightarrow{\boldsymbol{r}}_{\boldsymbol{t}_{\epsilon}}, \quad \vec{\omega}_{\boldsymbol{t}_{\epsilon}} \perp \overrightarrow{\boldsymbol{r}}_{\boldsymbol{t}_{\epsilon}} \tag{5.1}
\end{equation*}
$$

The considered motion quantities $\overrightarrow{\boldsymbol{\omega}}_{\boldsymbol{t}_{\epsilon}}$ and $\overrightarrow{\boldsymbol{r}}_{\boldsymbol{t}_{\epsilon}}$ are determined in the thought experiment by finding the successive positions of a single aerosol moving from a point $\overrightarrow{\boldsymbol{x}}_{0}$ after a time $t_{\varepsilon}$ to $\overrightarrow{\boldsymbol{x}}_{1}$ and another time $t_{\varepsilon}$ to $\overrightarrow{\boldsymbol{x}}_{2}$. By these 3 points a circle segment is uniquely defined for the point $\overrightarrow{\boldsymbol{x}}_{1}$ with radius vector $\overrightarrow{\boldsymbol{r}}_{t_{\varepsilon}}$ and a rotation speed $\overrightarrow{\boldsymbol{\omega}}_{t_{\varepsilon}}$.

$$
\begin{gather*}
\overrightarrow{\boldsymbol{r}}_{t_{\varepsilon}}=\boldsymbol{r}_{t_{\varepsilon}} \cdot \overrightarrow{\boldsymbol{\Theta}}_{t_{\varepsilon}} \\
\overrightarrow{\boldsymbol{\omega}}_{t_{\varepsilon}}=\boldsymbol{\omega}_{t_{\varepsilon}} \cdot \overrightarrow{\boldsymbol{\Omega}}_{t_{\varepsilon}} \tag{5.2}
\end{gather*}
$$

In the special case $\overrightarrow{\boldsymbol{\omega}}_{t_{\varepsilon}} \rightarrow \mathbf{0}$ and $\overrightarrow{\boldsymbol{r}} \rightarrow+\infty$ the velocity $\overrightarrow{\mathbf{v}}_{\boldsymbol{t}}$ is revealed out of its neighborhood. ${ }^{1}$ The aerosol density distributions are received in a thought experiment by an unlimited number of deterministic ensemble-systems. In every point ( $\overrightarrow{\boldsymbol{x}}, t$ ) a continuously differentiable aerosol density distribution of the motion quantities $\overrightarrow{\boldsymbol{\omega}}_{t_{\epsilon}}$ and $\overrightarrow{\boldsymbol{r}}_{t_{\epsilon}}$ is assigned in accordance with

$$
\begin{equation*}
f_{\boldsymbol{t}_{\epsilon}}=f_{\boldsymbol{t}_{\epsilon}}(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\omega}}, \overrightarrow{\boldsymbol{r}}) \tag{5.3}
\end{equation*}
$$

The with $\boldsymbol{t}_{\boldsymbol{\epsilon}}$ indexed functions are automatically assumed to contain motion quantities of corresponding measurement accuracies. The indexing of the motion quantities can be omitted if the functions are indexed. After execution of a limiting process

$$
\begin{equation*}
\lim _{\boldsymbol{t}_{\epsilon} \rightarrow 0} f_{\boldsymbol{t}_{\epsilon}}(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\omega}}, \overrightarrow{\boldsymbol{r}})=f(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\omega}}, \overrightarrow{\boldsymbol{r}}) \tag{5.4}
\end{equation*}
$$

f and $(\overrightarrow{\boldsymbol{\omega}}, \overrightarrow{\boldsymbol{r}})$ are understood according to an exact measuring process. Integrating the aerosol density distribution over the motion quantities one obtains expectation values of a aerosol density not conforming with the actual aerosol density $\rho$.

$$
\begin{equation*}
<\rho_{t_{\varepsilon}}(\vec{x}, t)>=\int_{2 \pi} \int_{4 \pi} \int_{0}^{\infty} \int_{0}^{\infty} f_{\boldsymbol{t}_{\epsilon}}(\overrightarrow{\boldsymbol{x}}, t, \boldsymbol{\omega} \cdot \overrightarrow{\boldsymbol{\Omega}}, \boldsymbol{r} \cdot \overrightarrow{\boldsymbol{\Theta}}) d \boldsymbol{\omega} d \boldsymbol{r} d \overrightarrow{\boldsymbol{\Omega}} d \overrightarrow{\boldsymbol{\Theta}} \neq \rho_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}(\overrightarrow{\boldsymbol{x}}, t) \tag{5.5}
\end{equation*}
$$

A rigorously derived partial differential equation is obtained, which can be used to calculate the evolution of the spatiotemporal aerosol density distributions. The initially unbounded number of unknown coefficients is attributed to local time scaling. The abstractly formulated transition probabilities get concrete functional dependencies.

[^4]
### 5.2 The Transport as Markov Process with Natural Causality

A aerosol at location $\overrightarrow{\boldsymbol{x}}$ and time t changing its velocity from $\overrightarrow{\boldsymbol{v}}^{\prime}=\left(\overrightarrow{\boldsymbol{\omega}}^{\prime} \times \overrightarrow{\boldsymbol{r}}^{\prime}\right)$ to $\overrightarrow{\boldsymbol{v}}=(\overrightarrow{\boldsymbol{\omega}} \times \overrightarrow{\boldsymbol{r}})$ is given by the transition probability

$$
\begin{equation*}
W_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}=W_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}\left(\overrightarrow{\boldsymbol{x}}, t ; \overrightarrow{\boldsymbol{\omega}}, \overrightarrow{\boldsymbol{r}} ; \overrightarrow{\boldsymbol{\omega}}^{\prime}, \overrightarrow{\boldsymbol{r}}^{\prime}\right) \tag{5.6}
\end{equation*}
$$

with

$$
\begin{equation*}
\int_{0}^{\infty} \int_{0}^{\infty} \int_{4 \pi} \int_{2 \pi} W_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}\left(\overrightarrow{\boldsymbol{x}}, t ; \overrightarrow{\boldsymbol{\omega}}, \overrightarrow{\boldsymbol{r}} ; \overrightarrow{\boldsymbol{\omega}}^{\prime}, \overrightarrow{\boldsymbol{r}}^{\prime}\right) d \boldsymbol{\omega}^{\prime} d r^{\prime} d \boldsymbol{\Omega}^{\prime} d \boldsymbol{\Theta}^{\prime}=1 . \tag{5.7}
\end{equation*}
$$

$\Longrightarrow$

$$
\begin{equation*}
f_{t_{\epsilon}}(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\omega}}, \overrightarrow{\boldsymbol{r}})=\int_{0}^{\infty} \int_{0}^{\infty} \int_{4 \pi} \int_{2 \pi} W_{\boldsymbol{t}_{\epsilon}}\left(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\omega}}, \overrightarrow{\boldsymbol{r}}, \overrightarrow{\boldsymbol{\omega}}^{\prime}, \overrightarrow{\boldsymbol{r}}^{\prime}\right) f_{\boldsymbol{t}_{\epsilon}}\left(\overrightarrow{\boldsymbol{x}}-\overrightarrow{\boldsymbol{\omega}}^{\prime} \times \overrightarrow{\boldsymbol{r}}^{\prime} \cdot t_{\epsilon}, \overrightarrow{\boldsymbol{\omega}}^{\prime}, \overrightarrow{\boldsymbol{r}}^{\prime}, t-t_{\epsilon}\right) d \boldsymbol{\omega}^{\prime} d r^{\prime} d \boldsymbol{\Omega}^{\prime} d \Theta^{\prime} \tag{5.8}
\end{equation*}
$$

Continuity is required respectively of all variables of the transition probability $W_{t_{\epsilon}}$. The sequence of velocities $\overrightarrow{\boldsymbol{v}}_{t_{\epsilon}}^{\prime}, \overrightarrow{\boldsymbol{v}}_{t_{\epsilon}}$ means a motion from

$$
\begin{equation*}
\left(\overrightarrow{\boldsymbol{x}}-\overrightarrow{\boldsymbol{\omega}}_{t_{\epsilon}}^{\prime} \times \overrightarrow{\boldsymbol{r}}_{t_{\epsilon}}^{\prime} \cdot t_{\boldsymbol{\epsilon}}, t-t_{\epsilon}, \overrightarrow{\boldsymbol{\omega}}_{t_{\epsilon}}^{\prime} \times \overrightarrow{\boldsymbol{r}}_{t_{\epsilon}}^{\prime}\right) \quad \text { to } \quad\left(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\omega}}_{t_{\epsilon}} \times \overrightarrow{\boldsymbol{r}}_{t_{\epsilon}}\right) . \tag{5.9}
\end{equation*}
$$

For the limiting process $t_{\varepsilon} \rightarrow 0$ the transition probabilities $W_{t_{\epsilon}}$ prove to be physical realizations of test functions of the distribution theory.

$$
\begin{equation*}
\lim _{\boldsymbol{t}_{\epsilon} \rightarrow 0} W_{\boldsymbol{t}_{\epsilon}}=\delta\left(\overrightarrow{\boldsymbol{\omega}}, \overrightarrow{\boldsymbol{r}} ; \overrightarrow{\boldsymbol{\omega}}^{\prime}, \overrightarrow{\boldsymbol{r}}^{\prime}\right) . \tag{5.10}
\end{equation*}
$$

The passive scalar aerosols precisely reproduce the motions of the fluctuation field. For the aerosol density distribution $f_{\boldsymbol{t}_{\epsilon}}(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\omega}}, \overrightarrow{\boldsymbol{r}})$ the following separation aproach is used without loss of generality:

$$
\begin{equation*}
f_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}\left(\overrightarrow{\boldsymbol{x}}-\overrightarrow{\boldsymbol{\omega}}^{\prime} \times \overrightarrow{\boldsymbol{r}}^{\prime} \cdot t_{\boldsymbol{\epsilon}}, t, \overrightarrow{\boldsymbol{\omega}}^{\prime}, \overrightarrow{\boldsymbol{r}}^{\prime}\right)=G_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}\left(\overrightarrow{\boldsymbol{x}}-\overrightarrow{\boldsymbol{\omega}}^{\prime} \times \overrightarrow{\boldsymbol{r}}^{\prime} \cdot t_{\boldsymbol{\epsilon}}, t, \overrightarrow{\boldsymbol{\omega}}^{\prime}, \overrightarrow{\boldsymbol{r}}^{\prime}\right) \bar{f}_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}\left(\overrightarrow{\boldsymbol{x}}-\overline{\boldsymbol{\Omega}} \overrightarrow{\boldsymbol{\Omega}}^{\prime} \times \overrightarrow{\boldsymbol{\Theta}}^{\prime} \cdot t_{\boldsymbol{\epsilon}}, t, \overrightarrow{\boldsymbol{\Omega}}^{\prime}, \overrightarrow{\boldsymbol{\Theta}}^{\prime}\right) \tag{5.11}
\end{equation*}
$$

with

$$
\begin{align*}
& \int_{0}^{\infty} \int_{0}^{\infty} G_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}(\overrightarrow{\boldsymbol{x}}, t, \omega \overrightarrow{\boldsymbol{\Omega}}, r \overrightarrow{\boldsymbol{\Theta}}) d \omega d r=1 \\
& \int_{0}^{\infty} \int_{0}^{\infty} G_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}(\overrightarrow{\boldsymbol{x}}, t, \omega \overrightarrow{\boldsymbol{\Omega}}, r \overrightarrow{\boldsymbol{\Theta}}) \omega r d \omega d r=\bar{v}(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\Omega}}, \overrightarrow{\boldsymbol{\Theta}})  \tag{5.12}\\
& \bar{v}(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\Omega}}, \overrightarrow{\boldsymbol{\Theta}})=\bar{\omega}(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\Omega}}, \overrightarrow{\boldsymbol{\Theta}}) \cdot \bar{r}(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\Omega}}, \overrightarrow{\boldsymbol{\Theta}})
\end{align*}
$$

$\Longrightarrow$

$$
\begin{equation*}
\bar{f}_{\boldsymbol{t}_{\epsilon}}\left(\overrightarrow{\boldsymbol{x}}-\bar{v} \overrightarrow{\boldsymbol{\Omega}} \times \overrightarrow{\boldsymbol{\Theta}} \cdot t_{\boldsymbol{\epsilon}}, t, \overrightarrow{\boldsymbol{\Omega}}, \overrightarrow{\boldsymbol{\Theta}}\right)=\int_{0}^{\infty} \int_{0}^{\infty} f_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}\left(\overrightarrow{\boldsymbol{x}}-\overrightarrow{\boldsymbol{\omega}} \times \overrightarrow{\boldsymbol{r}} \cdot t_{\boldsymbol{\epsilon}}, t, \boldsymbol{\omega} \cdot \overrightarrow{\boldsymbol{\Omega}}, \boldsymbol{r} \cdot \overrightarrow{\boldsymbol{\Theta}}\right) d \omega d r \tag{5.13}
\end{equation*}
$$

One obtains a transition probability $\bar{W}_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}$ only depending on the directions by integrating $W_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}$ over the amounts $\omega^{\prime}, r^{\prime}, \omega, r$

$$
\begin{equation*}
\bar{W}_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}\left(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\Omega}}, \overrightarrow{\boldsymbol{\Theta}}, \overrightarrow{\boldsymbol{\Omega}}^{\prime}, \overrightarrow{\boldsymbol{\Theta}}^{\prime}\right)=\int_{0}^{\infty} \int_{0}^{\infty} \int_{0}^{\infty} \int_{0}^{\infty} W_{\boldsymbol{t}_{\epsilon}} G_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}\left(\overrightarrow{\boldsymbol{x}}-\overrightarrow{\boldsymbol{\omega}}^{\prime} \times \overrightarrow{\boldsymbol{r}}^{\prime} \cdot t_{\boldsymbol{\epsilon}}, t-t_{\boldsymbol{\epsilon}}, \overrightarrow{\boldsymbol{\omega}}^{\prime}, \overrightarrow{\boldsymbol{r}}^{\prime}\right) d \omega^{\prime} d r^{\prime} d \omega d r \tag{5.14}
\end{equation*}
$$

The integration

$$
\begin{equation*}
\int_{0}^{\infty} \int_{0}^{\infty}(5.8) d \omega d r \tag{5.15}
\end{equation*}
$$

gives

$$
\begin{gather*}
\bar{f}_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\Omega}}, \overrightarrow{\boldsymbol{\Theta}})=\int_{0}^{\infty} \int_{0}^{\infty} \int_{0}^{\infty} \int_{0}^{\infty} \int_{4 \pi} \int_{2 \pi} W_{\boldsymbol{t}_{\boldsymbol{\epsilon}}} f_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}\left(\overrightarrow{\boldsymbol{x}}-\overrightarrow{\boldsymbol{\omega}}^{\prime} \times \overrightarrow{\boldsymbol{r}}^{\prime} \cdot t_{\epsilon}, t-t_{\boldsymbol{\epsilon}}, \overrightarrow{\boldsymbol{\omega}}^{\prime}, \overrightarrow{\boldsymbol{r}}^{\prime}\right) d \omega^{\prime} d r^{\prime} d \omega d r d \overrightarrow{\boldsymbol{\Omega}}^{\prime} d \overrightarrow{\boldsymbol{\Theta}}^{\prime}  \tag{5.16}\\
\Longrightarrow \bar{f}_{\boldsymbol{t}_{\epsilon}}(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\Omega}}, \overrightarrow{\boldsymbol{\Theta}})=\int_{4 \pi} \int_{2 \pi} \bar{W}_{\boldsymbol{t}_{\epsilon}} \bar{f}_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}\left(\overrightarrow{\boldsymbol{x}}-\bar{v}^{\prime} \overrightarrow{\boldsymbol{\Omega}}^{\prime} \times \overrightarrow{\boldsymbol{\Theta}}^{\prime} \cdot t_{\epsilon}, t-t_{\epsilon}, \overrightarrow{\boldsymbol{\Omega}}^{\prime}, \overrightarrow{\boldsymbol{\Theta}}^{\prime}\right) d \overrightarrow{\boldsymbol{\Omega}}^{\prime} d \overrightarrow{\boldsymbol{\Theta}}^{\prime} \tag{5.17}
\end{gather*}
$$

In the integrand $\bar{f}_{\boldsymbol{t}_{\epsilon}}$ is developed around $\overrightarrow{\boldsymbol{x}}$ and t :

$$
\begin{equation*}
\bar{f}_{\boldsymbol{t}_{\epsilon}}\left(\overrightarrow{\boldsymbol{x}}-\Delta \overrightarrow{\boldsymbol{x}}, t-t_{\epsilon}, \overrightarrow{\boldsymbol{\Omega}}^{\prime}, \overrightarrow{\boldsymbol{\Theta}}^{\prime}\right)=\bar{f}_{\boldsymbol{t}_{\epsilon}}\left(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\Omega}}^{\prime}, \overrightarrow{\boldsymbol{\Theta}}^{\prime}\right)-\tau_{E} \cdot \epsilon \cdot\left[\frac{\partial \bar{f}_{\boldsymbol{t}_{\epsilon}}^{\prime}}{\partial t}+\bar{v}^{\prime} \overrightarrow{\boldsymbol{\Omega}}^{\prime} \times \overrightarrow{\boldsymbol{\Theta}}^{\prime} \cdot \nabla \bar{f}_{\boldsymbol{t}_{\epsilon}}^{\prime}+O\left(\epsilon^{2}\right)\right] \tag{5.18}
\end{equation*}
$$

This leads to

$$
\begin{align*}
& \frac{\int_{4 \pi} \int_{2 \pi} \bar{W}_{\boldsymbol{t}_{\boldsymbol{\epsilon}}} \bar{f}_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}^{\prime} d \overrightarrow{\boldsymbol{\Omega}}^{\prime} d \overrightarrow{\boldsymbol{\Theta}}^{\prime}-\bar{f}_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}}{\epsilon}= \\
& \int_{4 \pi} \int_{2 \pi} \bar{W}_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}\left(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\Omega}}^{\prime}, \overrightarrow{\boldsymbol{\Theta}}^{\prime}, \overrightarrow{\boldsymbol{\Omega}}, \overrightarrow{\boldsymbol{\Theta}}\right) \cdot \tau_{E}\left[\cdot \frac{\partial \bar{f}_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}^{\prime}}{\partial t}+\bar{v}^{\prime} \overrightarrow{\boldsymbol{\Omega}}^{\prime} \times \overrightarrow{\boldsymbol{\Theta}}^{\prime} \cdot t_{\boldsymbol{\epsilon}} \cdot \nabla \bar{f}_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}^{\prime}+O\left(\epsilon^{2}\right)\right] d \overrightarrow{\boldsymbol{\Omega}}^{\prime} d \overrightarrow{\boldsymbol{\Theta}}^{\prime} \tag{5.19}
\end{align*}
$$

As

$$
\begin{equation*}
\lim _{t_{\epsilon} \rightarrow 0} \bar{W}_{t_{\epsilon}}=\delta\left(\overrightarrow{\boldsymbol{\Omega}}, \overrightarrow{\boldsymbol{\Theta}} ; \overrightarrow{\boldsymbol{\Omega}}^{\prime}, \overrightarrow{\boldsymbol{\Theta}}^{\prime}\right) \tag{5.20}
\end{equation*}
$$

$\Longrightarrow$

$$
\begin{equation*}
\lim _{\epsilon \rightarrow 0} \frac{\int_{4 \pi} \int_{2 \pi} \bar{W}_{\boldsymbol{t}_{\epsilon}} \bar{f}_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}^{\prime} d \overrightarrow{\boldsymbol{\Omega}}^{\prime} d \overrightarrow{\boldsymbol{\Theta}}^{\prime}-\bar{f}_{\boldsymbol{t}_{\epsilon}}}{\epsilon \cdot \tau_{E}}=\frac{\partial \bar{f}}{\partial t}+\bar{v} \overrightarrow{\boldsymbol{\Omega}} \times \overrightarrow{\boldsymbol{\Theta}} \cdot \nabla \bar{f} \tag{5.21}
\end{equation*}
$$

Furtheron

$$
\begin{equation*}
\lim _{\epsilon \rightarrow 0} \frac{\int_{4 \pi} \int_{2 \pi} \bar{W}_{\boldsymbol{t}_{\epsilon}} \bar{f}_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}^{\prime} d \overrightarrow{\boldsymbol{\Omega}}^{\prime} d \overrightarrow{\boldsymbol{\Theta}}^{\prime}-\bar{f}_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}}{\epsilon \cdot \tau_{E}} \tag{5.22}
\end{equation*}
$$

is called exchange-term.

### 5.3 Calculation of the Exchange-Term

Exchange term dependencies of scalar products $\overrightarrow{\boldsymbol{\Omega}} \cdot \overrightarrow{\boldsymbol{\Omega}}^{\prime}$ and $\overrightarrow{\boldsymbol{\Theta}} \cdot \overrightarrow{\boldsymbol{\Theta}}^{\prime}$ are taken into account istead of individually depending directions $\overrightarrow{\boldsymbol{\Omega}}, \overrightarrow{\boldsymbol{\Omega}}^{\prime}$ and $\overrightarrow{\boldsymbol{\Theta}}, \overrightarrow{\boldsymbol{\Theta}}$ demanding the following relation

$$
\begin{equation*}
\lim _{\epsilon \rightarrow 0} \frac{\int_{2 \pi} \int_{4 \pi} \bar{W}_{t_{\epsilon}} \bar{f}_{t_{\epsilon}}^{\prime} d \overrightarrow{\boldsymbol{\Omega}}^{\prime} d \overrightarrow{\boldsymbol{\Theta}}^{\prime}-\bar{f}_{\boldsymbol{t}_{\epsilon}}}{\epsilon \cdot \tau_{E}}=\lim _{\epsilon \rightarrow 0} \frac{\int_{2 \pi} \int_{4 \pi} \widetilde{W}_{t_{\epsilon}}\left(\overrightarrow{\boldsymbol{\Omega}} \cdot \overrightarrow{\boldsymbol{\Omega}}^{\prime}, \overrightarrow{\boldsymbol{\Theta}} \cdot \overrightarrow{\boldsymbol{\Theta}}^{\prime}\right) \bar{f}_{\boldsymbol{t}_{\epsilon}}^{\prime} d \overrightarrow{\boldsymbol{\Omega}}^{\prime} d \overrightarrow{\boldsymbol{\Theta}}^{\prime}-\bar{f}_{\boldsymbol{t}_{\epsilon}}}{\epsilon \cdot t_{E}} . \tag{5.23}
\end{equation*}
$$

The following transitions

$$
\begin{array}{rll}
\tau_{E}=\text { const } & \longrightarrow & t_{E}=t_{E}(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\Omega}}, \overrightarrow{\boldsymbol{\Theta}}) \\
\bar{W}_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}\left(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\Omega}}, \overrightarrow{\boldsymbol{\Theta}} ; \overrightarrow{\boldsymbol{\Omega}}^{\prime}, \overrightarrow{\boldsymbol{\Theta}}^{\prime}\right) & \longrightarrow & \widetilde{W}_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}\left(\overrightarrow{\boldsymbol{\Omega}} \cdot \overrightarrow{\boldsymbol{\Omega}}^{\prime}, \overrightarrow{\boldsymbol{\Theta}} \cdot \overrightarrow{\boldsymbol{\Theta}}^{\prime}\right) \tag{5.24}
\end{array}
$$

are regarded. Moreover, a separation of $\overrightarrow{\boldsymbol{\Omega}} \cdot \overrightarrow{\boldsymbol{\Omega}}^{\prime}$ and $\overrightarrow{\boldsymbol{\Theta}} \cdot \overrightarrow{\boldsymbol{\Theta}}^{\prime}$ is asumed:

$$
\begin{equation*}
\widetilde{W}_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}\left(\overrightarrow{\boldsymbol{\Omega}} \cdot \overrightarrow{\boldsymbol{\Omega}}^{\prime}, \overrightarrow{\boldsymbol{\Theta}} \cdot \overrightarrow{\boldsymbol{\Theta}}^{\prime}\right)=V_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}\left(\overrightarrow{\boldsymbol{\Omega}} \cdot \overrightarrow{\boldsymbol{\Omega}}^{\prime}\right) \cdot M_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}\left(\overrightarrow{\boldsymbol{\Theta}} \cdot \overrightarrow{\boldsymbol{\Theta}}^{\prime}\right) \tag{5.25}
\end{equation*}
$$

Functions of the unit vectors $\overrightarrow{\boldsymbol{\Omega}}$ and $\overrightarrow{\boldsymbol{\Theta}}$ are presented by a complete orthogonal function system representing an extension of the spherical harmonics called turbulence functions.

$$
\begin{align*}
& \bar{f}_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\Omega}}, \overrightarrow{\boldsymbol{\Theta}})=\sum_{l=0}^{+\infty} \sum_{m=-l}^{+l} \sum_{k=-\infty}^{+\infty} \bar{f}_{\boldsymbol{t}_{\boldsymbol{\epsilon}} l m k}(\overrightarrow{\boldsymbol{x}}, t) Q_{l m k}(\overrightarrow{\boldsymbol{\Omega}}, \overrightarrow{\boldsymbol{\Theta}}) \\
&=\sum_{l=0}^{+\infty} \sum_{m=-l}^{+l} \sum_{k=-\infty}^{+\infty} \bar{f}_{\boldsymbol{t}_{\boldsymbol{\epsilon}} l m k}(\overrightarrow{\boldsymbol{x}}, t) Q_{l m k}^{*}(\overrightarrow{\boldsymbol{\Omega}}, \overrightarrow{\boldsymbol{\Theta}})  \tag{5.26}\\
& \int_{2 \pi} \int_{4 \pi} Q_{l m k}(\overrightarrow{\boldsymbol{\Omega}}, \overrightarrow{\boldsymbol{\Theta}}) Q_{l m k}^{*}\left(\overrightarrow{\boldsymbol{\Omega}}^{\prime}, \overrightarrow{\boldsymbol{\Theta}}^{\prime}\right) d \overrightarrow{\boldsymbol{\Omega}}^{\prime} d \overrightarrow{\boldsymbol{\Theta}}^{\prime}= \begin{cases}\frac{8 \pi^{2}}{2 l+1} & \text { for } \mathrm{l}=\mathrm{l}^{\prime} \text { and } \mathrm{m}=\mathrm{m} \\
0 & \text { else }\end{cases} \tag{5.27}
\end{align*}
$$

with

$$
\begin{align*}
Q_{l m k}(\overrightarrow{\boldsymbol{\Omega}}, \overrightarrow{\boldsymbol{\Theta}}) & =P_{l m}(\overrightarrow{\boldsymbol{\Omega}}) H_{k}(\overrightarrow{\boldsymbol{\Theta}}) \\
\int_{2 \pi} H_{k^{\prime}}(\overrightarrow{\boldsymbol{\Theta}}) H_{k}^{*}(\overrightarrow{\boldsymbol{\Theta}}) d \overrightarrow{\boldsymbol{\Theta}} & = \begin{cases}2 \pi & \text { for k' }{ }^{\prime}=\mathrm{k} \\
0 & \text { else }\end{cases}  \tag{5.28}\\
H_{k}(\overrightarrow{\boldsymbol{\Theta}}) & =e^{i k \theta}
\end{align*}
$$

The product $\overrightarrow{\boldsymbol{\Omega}} \cdot \overrightarrow{\boldsymbol{\Omega}}^{\prime}$ in the separated exchange function $V_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}$ is developed by spherical harmonics.

$$
\begin{gather*}
V_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}\left(\overrightarrow{\boldsymbol{\Omega}}^{\prime} \cdot \overrightarrow{\boldsymbol{\Omega}}\right)=\sum_{l=0}^{+\infty} V_{\boldsymbol{t}_{\boldsymbol{\epsilon}} l} P_{l}(\cos (\alpha))=\sum_{l=0}^{+\infty} V_{\boldsymbol{t}_{\boldsymbol{\epsilon}} l} \sum_{m=-l}^{m=+l} P_{l m}\left(\overrightarrow{\boldsymbol{\Omega}}^{\prime}\right) P_{l m}^{*}(\overrightarrow{\boldsymbol{\Omega}}) \\
\text { with }  \tag{5.29}\\
\lim _{t \boldsymbol{\epsilon} \rightarrow 0} V_{\boldsymbol{t}_{\epsilon}}\left(\overrightarrow{\boldsymbol{\Omega}}^{\prime} \cdot \overrightarrow{\boldsymbol{\Omega}}\right)=\delta\left(\overrightarrow{\boldsymbol{\Omega}}, \overrightarrow{\boldsymbol{\Omega}}^{\prime}\right)
\end{gather*}
$$

The product $\overrightarrow{\boldsymbol{\Theta}} \cdot \overrightarrow{\boldsymbol{\Theta}}^{\prime}$ in the separated exchange function $M_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}$ is developed by functions $H_{k}$.

$$
\begin{equation*}
M_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}\left(\overrightarrow{\boldsymbol{\Theta}}^{\prime} \cdot \overrightarrow{\boldsymbol{\Theta}}\right)=\sum_{k=0}^{+\infty} M_{\boldsymbol{t}_{\epsilon} k} \cos (k \beta)=\frac{1}{2} \sum_{k=0}^{k=+\infty} M_{\boldsymbol{t}_{\boldsymbol{\epsilon}} k}\left[H_{k}\left(\overrightarrow{\boldsymbol{\Theta}}^{\prime}\right) H_{k}^{*}(\overrightarrow{\boldsymbol{\Theta}})+H_{-k}\left(\overrightarrow{\boldsymbol{\Theta}}^{\prime}\right) H_{-k}^{*}(\overrightarrow{\boldsymbol{\Theta}})\right] \tag{5.30}
\end{equation*}
$$

with

$$
\begin{align*}
\cos (k \beta)= & \frac{1}{2}\left[H_{k}\left(\overrightarrow{\boldsymbol{\Theta}}^{\prime}\right) H_{k}^{*}(\overrightarrow{\boldsymbol{\Theta}})+H_{-k}\left(\overrightarrow{\boldsymbol{\Theta}}^{\prime}\right) H_{-k}^{*}(\overrightarrow{\boldsymbol{\Theta}})\right]=\frac{1}{2}\left[e^{i k\left(\theta^{\prime}-\theta\right)}+e^{-i k\left(\theta^{\prime}-\theta\right)}\right] \\
\overrightarrow{\boldsymbol{\Theta}}^{\prime} \cdot \overrightarrow{\boldsymbol{\Theta}}=\cos (\beta)=\cos \left(\theta^{\prime}-\theta\right)= & \frac{1}{2}\left[H_{1}\left(\overrightarrow{\boldsymbol{\Theta}}^{\prime}\right) H_{1}^{*}(\overrightarrow{\boldsymbol{\Theta}})+H_{-1}\left(\overrightarrow{\boldsymbol{\Theta}}^{\prime}\right) H_{-1}^{*}(\overrightarrow{\boldsymbol{\Theta}})\right]=\frac{1}{2}\left[e^{i\left(\theta^{\prime}-\theta\right)}+e^{-i\left(\theta^{\prime}-\theta\right)}\right]  \tag{5.31}\\
& \lim _{t_{\boldsymbol{\epsilon}} \rightarrow 0} M_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}\left(\overrightarrow{\boldsymbol{\Theta}}^{\prime} \cdot \overrightarrow{\boldsymbol{\Theta}}\right)=\delta\left(\overrightarrow{\boldsymbol{\Theta}}, \overrightarrow{\boldsymbol{\Theta}}^{\prime}\right)
\end{align*}
$$

$\Longrightarrow$

$$
\begin{align*}
& \int_{4 \pi} \int_{2 \pi} \widetilde{W}_{\boldsymbol{t}_{\boldsymbol{\epsilon}}} \bar{f}_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}^{\prime} d \overrightarrow{\boldsymbol{\Omega}}^{\prime} d \overrightarrow{\boldsymbol{\Theta}}=\int_{4 \pi} \int_{2 \pi} V_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}\left(\overrightarrow{\boldsymbol{\Omega}}^{\prime} \cdot \overrightarrow{\boldsymbol{\Omega}}\right) \cdot M_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}\left(\overrightarrow{\boldsymbol{\Theta}}^{\prime} \cdot \overrightarrow{\boldsymbol{\Theta}}\right) \bar{f}_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}^{\prime} d \overrightarrow{\boldsymbol{\Omega}}^{\prime} d \overrightarrow{\boldsymbol{\Theta}}^{\prime} \\
& =\int_{4 \pi} \int_{2 \pi}\left[\left\{\sum_{l=0}^{+\infty} V_{\boldsymbol{t}_{\boldsymbol{\epsilon}} l} \sum_{m=-l}^{m=+l} P_{l m}\left(\overrightarrow{\boldsymbol{\Omega}}^{\prime}\right) P_{l m}^{*}(\overrightarrow{\boldsymbol{\Omega}}) \cdot \frac{1}{2=\sum_{k=0}^{+\infty}} M_{\boldsymbol{t}_{\boldsymbol{\epsilon}} k}\left[H_{k}\left(\overrightarrow{\boldsymbol{\Theta}}^{\prime}\right) H_{k}^{*}(\overrightarrow{\boldsymbol{\Theta}})+H_{-k}\left(\overrightarrow{\boldsymbol{\Theta}}^{\prime}\right) H_{-k}^{*}(\overrightarrow{\boldsymbol{\Theta}})\right]\right\}\right. \\
& \left.\cdot \sum_{l=0}^{+\infty} \sum_{m=-l}^{+l} \sum_{k=-\infty}^{+\infty} \bar{f}_{\boldsymbol{t}_{\boldsymbol{\epsilon}} l m k}(\overrightarrow{\boldsymbol{x}}, t) P_{l m}^{*}\left(\overrightarrow{\boldsymbol{\Omega}}^{\prime}\right) H_{k}^{*}\left(\overrightarrow{\boldsymbol{\Theta}}^{\prime}\right)\right] d \overrightarrow{\boldsymbol{\Omega}}^{\prime} d \overrightarrow{\boldsymbol{\Theta}}^{\prime}  \tag{5.32}\\
& =\sum_{l=0}^{+\infty} V_{\boldsymbol{t}_{\boldsymbol{\epsilon}} l} \frac{4 \pi}{2 l+1} \sum_{m=-l}^{+l} P_{l m}^{*}(\overrightarrow{\boldsymbol{\Omega}}) \sum_{k=0}^{+\infty} M_{\boldsymbol{t}_{\boldsymbol{\epsilon}} k} 2 \pi \bar{f}_{\boldsymbol{t}_{\boldsymbol{\epsilon}} l m k}(\overrightarrow{\boldsymbol{x}}, t) H_{k}^{*}(\overrightarrow{\boldsymbol{\Theta}}) .
\end{align*}
$$

Finally the exchange-term results in

$$
\begin{align*}
& \lim _{\boldsymbol{\epsilon} \rightarrow 0} \frac{\int_{4 \pi} \int_{2 \pi} \widetilde{W}_{\boldsymbol{t}_{\boldsymbol{\epsilon}}} \bar{f}_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}^{\prime} d \overrightarrow{\boldsymbol{\Omega}}^{\prime} d \overrightarrow{\boldsymbol{\Theta}}^{\prime}-\bar{f}_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}}{\epsilon} \\
& =\lim _{\epsilon \rightarrow 0} \sum_{l=0}^{+\infty} \sum_{m=-l}^{+l} \sum_{k=0}^{+\infty} \frac{\left(V_{\boldsymbol{t}_{\boldsymbol{\epsilon}} l} \frac{4 \pi}{2 l+1} M_{\boldsymbol{t}_{\boldsymbol{\epsilon}} k} 2 \pi-1\right)}{\boldsymbol{\epsilon}} \bar{f}_{\boldsymbol{t}_{\boldsymbol{\epsilon}} l m k}(\overrightarrow{\boldsymbol{x}}, t) P_{l m}^{*}(\overrightarrow{\boldsymbol{\Omega}}) H_{k}^{*}(\overrightarrow{\boldsymbol{\Theta}})  \tag{5.33}\\
& =\sum_{l=0}^{+\infty} \sum_{m=-l}^{+l} \sum_{k=0}^{+\infty} \Upsilon_{l k} \bar{f}_{l m k}(\overrightarrow{\boldsymbol{x}}, t) P_{l m}(\overrightarrow{\boldsymbol{\Omega}}) H_{k}(\overrightarrow{\boldsymbol{\Theta}}) .
\end{align*}
$$

With the exchange coefficients

$$
\begin{equation*}
\Upsilon_{l k}=\lim _{\epsilon \rightarrow 0} \frac{\left(V_{\boldsymbol{t}_{\boldsymbol{\epsilon}} l} \frac{4 \pi}{2 l+1} M_{\boldsymbol{t}_{\boldsymbol{\epsilon}} k} 2 \pi-1\right)}{\boldsymbol{\epsilon}} \tag{5.34}
\end{equation*}
$$

the transport equation

$$
\begin{equation*}
\frac{\partial \bar{f}}{\partial t}+\bar{v} \overrightarrow{\boldsymbol{\Omega}} \times \overrightarrow{\boldsymbol{\Theta}} \cdot \nabla \bar{f}=\frac{1}{t_{E}} \sum_{l=0}^{+\infty} \sum_{m=-l}^{+l} \sum_{k=-\infty}^{+\infty} \Upsilon_{l k} \bar{f}_{l m k}(\overrightarrow{\boldsymbol{x}}, t) P_{l m}(\overrightarrow{\boldsymbol{\Omega}}) H_{k}(\overrightarrow{\boldsymbol{\Theta}}) \tag{5.35}
\end{equation*}
$$

is achieved. Further on it is shown that in $\Upsilon_{l k}$ the index k may be skipped.

### 5.4 Calculation of the Exchange-Coefficients $\Upsilon_{l}$

Considering an overall closed volume range V the aerosol number in the entire volume remains constant if no absorbtion is assumed.

$$
\begin{gather*}
\text { total number of aerosols }=\int_{\mathbb{V}} \int_{4 \pi} \int_{2 \pi} \bar{f} d \overrightarrow{\boldsymbol{\Omega}} d \overrightarrow{\boldsymbol{\Theta}} d \mathbb{V}=\text { const. }  \tag{5.36}\\
\Longrightarrow  \tag{5.37}\\
\frac{d}{d t} \int_{\mathbb{V}} \int_{4 \pi} \int_{2 \pi} \bar{f} d \overrightarrow{\boldsymbol{\Omega}} d \overrightarrow{\boldsymbol{\Theta}} d \mathbb{V}=\int_{\mathbb{V}} \int_{4 \pi} \int_{2 \pi}\left[\frac{\partial \bar{f}}{\partial t}+\bar{v} \overrightarrow{\boldsymbol{\Omega}} \times \overrightarrow{\boldsymbol{\Theta}} \cdot \nabla \bar{f}\right] d \overrightarrow{\boldsymbol{\Omega}} d \overrightarrow{\boldsymbol{\Theta}} d \mathbb{V}=\Upsilon_{0,0} \cdot \boldsymbol{V}=\mathbf{0}
\end{gather*}
$$

and thus

$$
\begin{equation*}
\Upsilon_{0,0}=0 \tag{5.38}
\end{equation*}
$$

Getting an overview over the exchange function $M_{t_{\epsilon}}$ the essential relations are presented again with the following equations:

$$
\begin{aligned}
M_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}\left(\overrightarrow{\boldsymbol{\Theta}}^{\prime} \cdot \overrightarrow{\boldsymbol{\Theta}}\right)=\sum_{k=0}^{+\infty} M_{\boldsymbol{t} \boldsymbol{\epsilon} k} \cos (k \beta)= & \frac{1}{2} \sum_{k=0}^{k=+\infty} M_{\boldsymbol{t}_{\boldsymbol{\epsilon}} k}\left[H_{k}\left(\overrightarrow{\boldsymbol{\Theta}}^{\prime}\right) H_{k}^{*}(\overrightarrow{\boldsymbol{\Theta}})+H_{-k}\left(\overrightarrow{\boldsymbol{\Theta}}^{\prime}\right) H_{-k}^{*}(\overrightarrow{\boldsymbol{\Theta}})\right] \\
\cos (k \beta)= & \frac{1}{2}\left[H_{k}\left(\overrightarrow{\boldsymbol{\Theta}}^{\prime}\right) H_{k}^{*}(\overrightarrow{\boldsymbol{\Theta}})+H_{-k}\left(\overrightarrow{\boldsymbol{\Theta}}^{\prime}\right) H_{-k}^{*}(\overrightarrow{\boldsymbol{\Theta}})\right]=\frac{1}{2}\left[e^{i k\left(\theta^{\prime}-\theta\right)}+e^{-i k\left(\theta^{\prime}-\theta\right)}\right] \\
\overrightarrow{\boldsymbol{\Theta}}^{\prime} \cdot \overrightarrow{\boldsymbol{\Theta}}=\cos (\beta)=\cos \left(\theta^{\prime}-\theta\right)= & \frac{1}{2}\left[H_{1}\left(\overrightarrow{\boldsymbol{\Theta}}^{\prime}\right) H_{1}^{*}(\overrightarrow{\boldsymbol{\Theta}})+H_{-1}\left(\overrightarrow{\boldsymbol{\Theta}}^{\prime}\right) H_{-1}^{*}(\overrightarrow{\boldsymbol{\Theta}})\right]=\frac{1}{2}\left[e^{i\left(\theta^{\prime}-\theta\right)}+e^{-i\left(\theta^{\prime}-\theta\right)}\right] \\
& \lim _{\epsilon \rightarrow 0} M_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}\left(\overrightarrow{\boldsymbol{\Theta}}^{\prime} \cdot \overrightarrow{\boldsymbol{\Theta}}\right)=\delta\left(\overrightarrow{\boldsymbol{\Theta}}, \overrightarrow{\boldsymbol{\Theta}}^{\prime}\right) \\
& \int_{2 \pi} H_{k^{\prime}}(\overrightarrow{\boldsymbol{\Theta}}) H_{k}^{*}(\overrightarrow{\boldsymbol{\Theta}}) d \overrightarrow{\boldsymbol{\Theta}}= \begin{cases}2 \pi & \text { für k} \\
0 & \text { else }\end{cases}
\end{aligned}
$$

$M_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}\left(\overrightarrow{\boldsymbol{\Theta}}^{\prime} \cdot \overrightarrow{\boldsymbol{\Theta}}\right)=\sum_{k=0}^{+\infty} M_{\boldsymbol{t}_{\boldsymbol{\epsilon}} k} \cos (k \beta)$ only takes values essentially different from 0 in an $\varepsilon$-neighborhood of $\beta=0$, such that $\overrightarrow{\boldsymbol{\Theta}}^{\prime} \cdot \overrightarrow{\boldsymbol{\Theta}}=\cos (\beta)=1-O\left(\varepsilon^{2}\right)$ is sufficient.$\Longrightarrow$

$$
\begin{align*}
2 \pi \cdot M_{\boldsymbol{t}_{\boldsymbol{\epsilon}} k} & = \\
\int_{-\pi}^{+\pi} M_{\boldsymbol{t}_{\boldsymbol{\epsilon}}} \cos (k \beta) d \beta & =\int_{-\pi}^{+\pi} M_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}(1-O(\epsilon)) d \beta=2 \pi \cdot M_{\boldsymbol{t}_{\boldsymbol{\epsilon}} 0}-O\left(\epsilon^{2}\right) . \tag{5.39}
\end{align*}
$$

On the other hand

$$
\begin{equation*}
\int_{2 \pi} M_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}(\overrightarrow{\boldsymbol{\Theta}} \cdot \overrightarrow{\boldsymbol{\Theta}})^{\prime} d \overrightarrow{\boldsymbol{\Theta}}^{\prime}=\frac{1}{2} \int_{2 \pi} \sum_{k=0}^{k=+\infty} M_{\boldsymbol{t}_{\boldsymbol{\epsilon}} k}\left[H_{k}\left(\overrightarrow{\boldsymbol{\Theta}}^{\prime}\right) H_{k}^{*}(\overrightarrow{\boldsymbol{\Theta}})+H_{-k}\left(\overrightarrow{\boldsymbol{\Theta}}^{\prime}\right) H_{-k}^{*}(\overrightarrow{\boldsymbol{\Theta}})\right] d \overrightarrow{\boldsymbol{\Theta}}^{\prime}=2 \pi \cdot M_{\boldsymbol{t}_{\boldsymbol{\epsilon}} 0}=1 \tag{5.40}
\end{equation*}
$$

is valid. $\Longrightarrow$

$$
\begin{equation*}
\lim _{t_{\epsilon} \rightarrow 0} M_{t_{\epsilon} k}=M_{t_{\epsilon} 0}=\frac{1}{2 \pi} . \tag{5.41}
\end{equation*}
$$

The calculation of the exchange coefficients is not influenced by $M_{t_{\epsilon}}$. The $\Upsilon$-values are given by

The transition probability is outlined by Legendre-polynomials respectively spherical harmonics:

$$
\begin{align*}
V_{\boldsymbol{t} \boldsymbol{\epsilon}}\left(\overrightarrow{\boldsymbol{\Omega}} \cdot \overrightarrow{\boldsymbol{\Omega}}^{\prime}\right)=\sum_{l=0}^{+\infty} V_{\boldsymbol{t}_{\boldsymbol{\epsilon}} l} P_{l}(\cos (\vartheta)) & =\sum_{l=0}^{+\infty} V_{\boldsymbol{t}_{\boldsymbol{\epsilon}} l} \sum_{m=-l}^{m=+l} P_{l m}(\overrightarrow{\boldsymbol{\Omega}}) P_{l m}^{*}\left(\overrightarrow{\boldsymbol{\Omega}}^{\prime}\right)  \tag{5.43}\\
\cos (\vartheta) & =\overrightarrow{\boldsymbol{\Omega}} \cdot \overrightarrow{\boldsymbol{\Omega}}^{\prime}=\mu
\end{align*}
$$

On the other hand is

$$
\begin{gather*}
\lim _{\boldsymbol{t}_{\boldsymbol{\epsilon}} \rightarrow 0} V_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}\left(\overrightarrow{\boldsymbol{\Omega}} \cdot \overrightarrow{\boldsymbol{\Omega}}^{\prime}\right)=\delta\left(\overrightarrow{\boldsymbol{\Omega}} \cdot \overrightarrow{\boldsymbol{\Omega}}^{\prime}\right) \\
\delta\left(\overrightarrow{\boldsymbol{\Omega}} \cdot \overrightarrow{\boldsymbol{\Omega}}^{\prime}\right)=\sum_{l=0}^{+\infty} \frac{2 l+1}{4 \pi} \sum_{m=-l}^{m=+l} P_{l m}(\overrightarrow{\boldsymbol{\Omega}}) P_{l m}^{*}\left(\overrightarrow{\boldsymbol{\Omega}}^{\prime}\right)=\sum_{l=0}^{+\infty} \frac{2 l+1}{4 \pi} \boldsymbol{P}_{l} \quad \text { see(8.17). } \tag{5.44}
\end{gather*}
$$

$V_{t_{\epsilon}}(\mu) \geqslant 0$ is only in the range $\mu \in[1-\varepsilon, 1]$ essentially different from 0 . So the Legendre polynomials are approximated by

$$
\begin{align*}
& P_{l}(\mu)=1-\left.\frac{d P_{l}}{d \mu}\right|_{1} \cdot \varepsilon+O\left(\varepsilon^{2}\right) \quad \varepsilon=1-\mu \\
& \left.\frac{d P_{l}}{d \mu}\right|_{1}=\frac{l(l+1)}{2} \text { see (8.1) } \boldsymbol{P}_{\mathbf{0}}=1, \boldsymbol{P}_{\mathbf{1}}=\mu  \tag{5.45}\\
& \Longrightarrow \\
& P_{l}(\mu)=P_{0}-\left(P_{0}-P_{1}\right) \frac{l(l+1)}{2}+O\left(\varepsilon^{2}\right) .
\end{align*}
$$

Using

$$
\begin{equation*}
\int_{-1}^{+1} P_{l} P_{l^{\prime}} d \mu=\delta_{l l^{\prime}} \frac{2}{2 l+1} \tag{5.46}
\end{equation*}
$$

follows

$$
\begin{equation*}
\int_{-1}^{+1} V_{\boldsymbol{t}_{\boldsymbol{\epsilon}}} P_{l} d \mu=2 V_{\boldsymbol{t}_{\epsilon} 0}-l(l+1) V_{\boldsymbol{t}_{\boldsymbol{\epsilon}} 0}+\frac{l(l+1)}{3} V_{\boldsymbol{t}_{\boldsymbol{\epsilon}} 1}=\frac{2}{2 l+1} V_{\boldsymbol{t}_{\boldsymbol{\epsilon}} l} . \tag{5.47}
\end{equation*}
$$

Furthermore is

$$
\begin{align*}
\int_{4 \pi} V_{t_{\boldsymbol{\epsilon}}}\left(\overrightarrow{\boldsymbol{\Omega}} \cdot \overrightarrow{\boldsymbol{\Omega}}^{\prime}\right) d \overrightarrow{\boldsymbol{\Omega}}^{\prime} & =\int_{4 \pi} V_{\boldsymbol{t}_{\epsilon} 0} d \overrightarrow{\boldsymbol{\Omega}}^{\prime}=4 \pi V_{\boldsymbol{t}_{\epsilon} 0}=\mathbf{1}  \tag{5.48}\\
\Longrightarrow \quad V_{\boldsymbol{t}_{\epsilon} 0} & =\frac{1}{4 \pi}
\end{align*}
$$

as $V_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}$ for $\boldsymbol{t}_{\boldsymbol{\epsilon}} \rightarrow \mathbf{0}$ degenerates to a $\delta$-function. That is why the $V_{\boldsymbol{t}_{\boldsymbol{\epsilon}} l}$ are expressed by $V_{\boldsymbol{t}_{\boldsymbol{\epsilon}} 1}$ and the determination of $V_{t_{\epsilon} 1}$ remains to be calculated. We set

$$
\begin{equation*}
\lim _{\epsilon \rightarrow 0} \frac{\left(V_{\boldsymbol{t} \epsilon} 1 \frac{4 \pi}{3}-1\right)}{\epsilon}=\boldsymbol{\zeta} \tag{5.49}
\end{equation*}
$$

Multiplying equation (5.47) with $2 \pi$ leads to

$$
\begin{equation*}
\frac{4 \pi}{2 l+1} V_{\boldsymbol{t}_{\boldsymbol{\epsilon}} l}=4 \pi V_{\boldsymbol{t}_{\boldsymbol{\epsilon}} 0}-4 \pi \frac{l(l+1)}{2} V_{\boldsymbol{t}_{\boldsymbol{\epsilon}} 0}+\frac{4 \pi}{3} \frac{l(l+1)}{2} V_{\boldsymbol{t}_{\boldsymbol{\epsilon}} 1} . \tag{5.50}
\end{equation*}
$$

I.e.

$$
\begin{equation*}
\frac{4 \pi}{2 l+1} V_{\boldsymbol{t}_{\epsilon} l}-\mathbf{1}=\frac{l(l+1)}{2}\left(\frac{4 \pi}{3} V_{\boldsymbol{t}_{\epsilon} 1}-1\right)=-\frac{l(l+1)}{2} \boldsymbol{\zeta}+O\left(\epsilon^{2}\right)=\mathbf{\Upsilon}_{\boldsymbol{l}}+O\left(\epsilon^{2}\right) \tag{5.51}
\end{equation*}
$$

$\Longrightarrow$

$$
\begin{equation*}
\Upsilon_{l}=-\frac{l(l+1)}{2} \zeta \quad \zeta=\text { const } . \tag{5.52}
\end{equation*}
$$

Now the equation of turbulent aerosol transport is written

$$
\begin{equation*}
\frac{\partial \bar{f}}{\partial t}+\bar{v} \overrightarrow{\boldsymbol{\Omega}} \times \overrightarrow{\boldsymbol{\Theta}} \cdot \nabla \bar{f}=-\frac{1}{t_{E}} \sum_{l=0}^{+\infty} \sum_{m=-l}^{+l} \sum_{k=-\infty}^{+\infty} \frac{l(l+1)}{2} \bar{f}_{l m k}(\overrightarrow{\boldsymbol{x}}, t) P_{l m}(\overrightarrow{\boldsymbol{\Omega}}) H_{k}(\overrightarrow{\boldsymbol{\Theta}}) \tag{5.53}
\end{equation*}
$$

the coefficient $\frac{\zeta}{t_{E}}$ replaced by $\frac{1}{t_{E}}$. A more complicated dependency of $t_{E}=t_{E}(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\Omega}}, \overrightarrow{\boldsymbol{\Theta}})$ possibly remains. Maybe, physically justified simplifications lead to practical solutions.

The total derivative with respect to time gives

$$
\begin{equation*}
\frac{d}{d t} \bar{f}(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\Omega}}, \overrightarrow{\boldsymbol{\Theta}})=\sum_{l=0}^{+\infty} \sum_{m=-l}^{m=+l} \sum_{k=-\infty}^{+\infty} \frac{d}{d t} \bar{f}_{l m k}(\overrightarrow{\boldsymbol{x}}, t) P_{l m}(\overrightarrow{\boldsymbol{\Omega}}) H_{k}(\overrightarrow{\boldsymbol{\Theta}})=\frac{1}{t_{E}} \sum_{l=1}^{+\infty} \gamma_{l} \cdot \sum_{m=-l}^{+l} \sum_{k=-\infty}^{+\infty} \bar{f}_{l m k}(\overrightarrow{\boldsymbol{x}}, t) P_{l m}(\overrightarrow{\boldsymbol{\Omega}}) H_{k}(\overrightarrow{\boldsymbol{\Theta}}) . \tag{5.54}
\end{equation*}
$$

The time behavior of the single modes are obtained by

$$
\begin{equation*}
\frac{d}{d t} \bar{f}_{l m k}(t)=\frac{\gamma_{l}}{t_{E}} \bar{f}_{l m k}, \quad \bar{f}_{l m k}(t) \sim \boldsymbol{e x p}\left(\frac{\gamma_{l}}{t_{E}} \cdot t\right) \tag{5.55}
\end{equation*}
$$

$$
\Longrightarrow
$$

The greater the order 1 the more powerful is its temporal decay. The function development can be terminated with the first order, since, as shown, such an approximation approaches asymptotically the exact solution with the distance against assumed sources and the time.

### 5.5 Reconstruction of the Transition Probabilities $\bar{W}_{\boldsymbol{t}_{\epsilon}}$

The transition probability $\widetilde{W}_{\boldsymbol{t}_{\boldsymbol{\epsilon}}, 0 \rightarrow 1 \rightarrow 2}$, an aerosol changing its motion pair of directions $(\overrightarrow{\boldsymbol{\Omega}}, \overrightarrow{\boldsymbol{\Theta}})$ at the times $t_{0}, t_{1}, t_{2}$ from $\left(\overrightarrow{\boldsymbol{\Omega}}_{0}, \overrightarrow{\boldsymbol{\Theta}}_{0}\right)$ via $\left(\overrightarrow{\boldsymbol{\Omega}}_{1}, \overrightarrow{\boldsymbol{\Theta}}_{1}\right)$ to $\left(\overrightarrow{\boldsymbol{\Omega}}_{2}, \overrightarrow{\boldsymbol{\Theta}}_{2}\right)$,

results out of the product of the single probabilities of the pairs of directions (vortex vector and radius vector direction of motion in a circle segment). The graphical presentation is meant symbolically because such a pair of directions does not compose to an overall direction. $\overrightarrow{\boldsymbol{\Omega}}_{i}$ is always orthogonal to $\overrightarrow{\boldsymbol{\Theta}}_{i}$. A vectorial overall direction of $\overrightarrow{\boldsymbol{\Omega}}_{i}$ and $\overrightarrow{\boldsymbol{\Theta}}_{i}$ has no physical meaning in the 3 dimensional space. ${ }^{2}$

$$
\begin{equation*}
\widetilde{W}_{\boldsymbol{t}_{\epsilon}, 0 \rightarrow 1 \rightarrow 2}=\widetilde{W}_{\frac{t_{\boldsymbol{\epsilon}}}{2}}\left(\overrightarrow{\boldsymbol{\Omega}}_{0} \cdot \overrightarrow{\boldsymbol{\Omega}}_{1}, \overrightarrow{\boldsymbol{\Theta}}_{0} \cdot \overrightarrow{\boldsymbol{\Theta}}_{1}\right) \cdot \widetilde{W}_{\frac{\boldsymbol{t}_{\boldsymbol{\epsilon}}}{2}}\left(\overrightarrow{\boldsymbol{\Omega}}_{1} \cdot \overrightarrow{\boldsymbol{\Omega}}_{2}, \overrightarrow{\boldsymbol{\Theta}}_{1} \cdot \overrightarrow{\boldsymbol{\Theta}}_{2}\right) \tag{5.56}
\end{equation*}
$$

[^5]The probability, that a aerosol changes its pair of directions within a time $t_{\varepsilon}=\varepsilon \cdot t_{E}$ from $\left(\overrightarrow{\boldsymbol{\Omega}}_{0}, \overrightarrow{\boldsymbol{\Theta}}_{0}\right)$ to $\left(\overrightarrow{\boldsymbol{\Omega}}_{2}, \overrightarrow{\boldsymbol{\Theta}}_{2}\right)$, is obtained by

$$
\begin{equation*}
\widetilde{W}_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}\left(\overrightarrow{\boldsymbol{\Omega}}_{0} \cdot \overrightarrow{\boldsymbol{\Omega}}_{2}, \overrightarrow{\boldsymbol{\Theta}}_{0} \cdot \overrightarrow{\boldsymbol{\Theta}}_{2}\right)=\int_{2 \pi} \int_{4 \pi} \widetilde{W}_{\frac{\boldsymbol{t}_{\boldsymbol{\epsilon}}^{2}}{}}\left(\overrightarrow{\boldsymbol{\Omega}}_{0} \cdot \overrightarrow{\boldsymbol{\Omega}}_{1}, \overrightarrow{\boldsymbol{\Theta}}_{0} \cdot \overrightarrow{\boldsymbol{\Theta}}_{1}\right) \cdot \widetilde{W}_{\frac{t_{\epsilon}}{2}}\left(\overrightarrow{\boldsymbol{\Omega}}_{1} \cdot \overrightarrow{\boldsymbol{\Omega}}_{2}, \overrightarrow{\boldsymbol{\Theta}}_{1} \cdot \overrightarrow{\boldsymbol{\Theta}}_{2}\right) d \overrightarrow{\boldsymbol{\Omega}}_{1} d \overrightarrow{\boldsymbol{\Theta}}_{1} . \tag{5.57}
\end{equation*}
$$

The evolution coefficients of the transition probability are available for sufficiently small $\epsilon$

$$
\begin{equation*}
\widetilde{W}_{\frac{t_{\epsilon} l}{2}} \approx\left\{1+\frac{\epsilon \cdot \Upsilon_{l}}{2}\right\} \frac{2 l+1}{4 \pi} \cdot \frac{1}{2 \pi} \tag{5.58}
\end{equation*}
$$

and therefore
$\widetilde{W}_{\frac{\boldsymbol{t}_{\boldsymbol{\epsilon}}}{2}}\left(\overrightarrow{\boldsymbol{\Omega}}_{0} \cdot \overrightarrow{\boldsymbol{\Omega}}_{1}, \overrightarrow{\boldsymbol{\Theta}}_{0} \cdot \overrightarrow{\boldsymbol{\Theta}}_{1}\right) \approx \sum_{l=0}^{+\infty}\left\{1+\frac{\epsilon \cdot \Upsilon_{l}}{2}\right\} \frac{2 l+1}{4 \pi} \cdot \frac{1}{2 \pi} \sum_{m=-l}^{+l} P_{l m}\left(\overrightarrow{\boldsymbol{\Omega}}_{1}\right) P_{l m}^{*}\left(\overrightarrow{\boldsymbol{\Omega}}_{0}\right) \cdot \frac{1}{2} \sum_{k=-\infty}^{+\infty}\left[H_{k}\left(\overrightarrow{\boldsymbol{\Theta}}_{1}\right) H_{k}^{*}\left(\overrightarrow{\boldsymbol{\Theta}}_{0}\right)+H_{-k}\left(\overrightarrow{\boldsymbol{\Theta}}_{1}\right) H_{-k}^{*}\left(\overrightarrow{\boldsymbol{\Theta}}_{0}\right)\right]$.
respectively
$\widetilde{W}_{\frac{t_{\boldsymbol{\epsilon}}}{2}}\left(\overrightarrow{\boldsymbol{\Omega}}_{1} \cdot \overrightarrow{\boldsymbol{\Omega}}_{2}, \overrightarrow{\boldsymbol{\Theta}}_{1} \cdot \overrightarrow{\boldsymbol{\Theta}}_{2}\right) \approx \sum_{l=0}^{+\infty}\left\{1+\frac{\epsilon \cdot \Upsilon_{l}}{2}\right\} \frac{2 l+1}{4 \pi} \cdot \frac{1}{2 \pi} \sum_{m=-l}^{+l} P_{l m}\left(\overrightarrow{\boldsymbol{\Omega}}_{2}\right) P_{l m}^{*}\left(\overrightarrow{\boldsymbol{\Omega}}_{1}\right) \cdot \frac{1}{2} \sum_{k=-\infty}^{+\infty}\left[H_{k}\left(\overrightarrow{\boldsymbol{\Theta}}_{2}\right) H_{k}^{*}\left(\overrightarrow{\boldsymbol{\Theta}}_{1}\right)+H_{-k}\left(\overrightarrow{\boldsymbol{\Theta}}_{2}\right) H_{-k}^{*}\left(\overrightarrow{\boldsymbol{\Theta}}_{1}\right)\right]$.
Integrating (5.57) one obtains

$$
\begin{equation*}
\widetilde{W}_{\boldsymbol{t} \boldsymbol{\epsilon}}\left(\overrightarrow{\boldsymbol{\Omega}}_{0} \cdot \overrightarrow{\boldsymbol{\Omega}}_{2}, \overrightarrow{\boldsymbol{\Theta}}_{0} \cdot \overrightarrow{\boldsymbol{\Theta}}_{2}\right) \approx \sum_{l=0}^{+\infty}\left\{1+\frac{\epsilon \cdot \Upsilon_{l}}{2}\right\}^{2} \frac{2 l+1}{4 \pi} \cdot \frac{1}{2 \pi} \sum_{m=-l}^{+l} P_{l m}^{*}\left(\overrightarrow{\boldsymbol{\Omega}}_{0}\right) P_{l m}\left(\overrightarrow{\boldsymbol{\Omega}}_{2}\right) \cdot \frac{1}{2} \sum_{k=-\infty}^{+\infty}\left[H_{k}\left(\overrightarrow{\boldsymbol{\Theta}}_{0}\right) H_{k}^{*}\left(\overrightarrow{\boldsymbol{\Theta}}_{2}\right)+H_{-k}\left(\overrightarrow{\boldsymbol{\Theta}}_{0}\right) H_{-k}^{*}\left(\overrightarrow{\boldsymbol{\Theta}}_{2}\right)\right] . \tag{5.61}
\end{equation*}
$$

Using n intermediate stages $\widetilde{W}_{\boldsymbol{t}_{\epsilon}}$ is expressed by an integral over the product of the single transition probabilities.

$$
\begin{align*}
& \widetilde{W}_{t_{\boldsymbol{\epsilon}}, 0 \rightarrow 1 \ldots \rightarrow n}=\widetilde{W}_{\frac{\boldsymbol{t} \boldsymbol{\epsilon}^{\epsilon}}{}}\left(\overrightarrow{\boldsymbol{\Omega}}_{0} \cdot \overrightarrow{\boldsymbol{\Omega}}_{1}, \overrightarrow{\boldsymbol{\Theta}}_{0} \cdot \overrightarrow{\boldsymbol{\Theta}}_{1}\right) \cdot \widetilde{W}_{\frac{\boldsymbol{t}_{\boldsymbol{\epsilon}}}{n}}\left(\overrightarrow{\boldsymbol{\Omega}}_{1} \cdot \overrightarrow{\boldsymbol{\Omega}}_{2}, \overrightarrow{\boldsymbol{\Theta}}_{1} \cdot \overrightarrow{\boldsymbol{\Theta}}_{2}\right) \ldots \widetilde{W}_{\frac{\boldsymbol{t}_{\boldsymbol{\epsilon}}}{n}}\left(\overrightarrow{\boldsymbol{\Omega}}_{n-1} \cdot \overrightarrow{\boldsymbol{\Omega}}_{n}, \overrightarrow{\boldsymbol{\Theta}}_{n-1} \cdot \overrightarrow{\boldsymbol{\Theta}}_{n}\right)  \tag{5.62}\\
& \widetilde{W}_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}\left(\overrightarrow{\boldsymbol{\Omega}}_{0} \cdot \overrightarrow{\boldsymbol{\Omega}}_{n}, \overrightarrow{\boldsymbol{\Theta}}_{0} \cdot \overrightarrow{\boldsymbol{\Theta}}_{n}\right)=\int_{2 \pi} \int_{2 \pi} \int_{4 \pi} \ldots \int_{2 \pi} \int_{4 \pi} \widetilde{W}_{\frac{\boldsymbol{t}_{\boldsymbol{\epsilon}}}{}} \cdot \widetilde{W}_{\frac{\boldsymbol{t}_{\boldsymbol{\epsilon}}}{n}} \ldots \widetilde{W}_{\frac{\boldsymbol{t}_{\boldsymbol{\epsilon}}}{n}} d \overrightarrow{\boldsymbol{\Omega}}_{1} d \overrightarrow{\boldsymbol{\Theta}}_{1} \ldots d \overrightarrow{\boldsymbol{\Omega}}_{n-1} d \overrightarrow{\boldsymbol{\Theta}}_{n-1} \tag{5.63}
\end{align*}
$$

$\widetilde{W}_{\boldsymbol{t}_{\epsilon}}\left(\overrightarrow{\boldsymbol{\Omega}} \cdot \overrightarrow{\boldsymbol{\Omega}}^{\prime}, \overrightarrow{\boldsymbol{\Theta}} \cdot \overrightarrow{\boldsymbol{\Theta}}^{\prime}\right)=\lim _{n \rightarrow \infty} \sum_{l=0}^{+\infty}\left\{1+\frac{\epsilon \cdot \Upsilon_{l}}{n}\right\}^{n} \frac{2 l+1}{4 \pi} \cdot \frac{1}{2 \pi} \cdot \sum_{m=-l}^{+l} P_{l m}^{*}(\overrightarrow{\boldsymbol{\Omega}}) P_{l m}\left(\overrightarrow{\boldsymbol{\Omega}}^{\prime}\right) \cdot \frac{1}{2} \sum_{k=-\infty}^{+\infty}\left[H_{k}\left(\overrightarrow{\boldsymbol{\Theta}}^{\prime}\right) H_{k}^{*}(\overrightarrow{\boldsymbol{\Theta}})+H_{-k}\left(\overrightarrow{\boldsymbol{\Theta}}^{\prime}\right) H_{-k}^{*}(\overrightarrow{\boldsymbol{\Theta}})\right]$
For $n \rightarrow \infty$ arises

$$
\begin{equation*}
\lim _{n \rightarrow \infty}\left\{1+\frac{\epsilon \cdot \Upsilon_{l}}{n}\right\}^{n}=e^{\Upsilon_{l} \cdot \epsilon} \tag{5.65}
\end{equation*}
$$

and using (5.63)
$\Longrightarrow$

$$
\begin{equation*}
\widetilde{W}_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}\left(\overrightarrow{\boldsymbol{\Omega}} \cdot \overrightarrow{\boldsymbol{\Omega}}^{\prime}, \overrightarrow{\boldsymbol{\Theta}} \cdot \overrightarrow{\boldsymbol{\Theta}}^{\prime}\right)=\sum_{l=0}^{+\infty} e^{\Upsilon_{l} \cdot \epsilon} \frac{2 l+1}{4 \pi} \cdot \frac{1}{2 \pi} \cdot \sum_{m=-l}^{+l} P_{l m}^{*}(\overrightarrow{\boldsymbol{\Omega}}) P_{l m}\left(\overrightarrow{\boldsymbol{\Omega}}^{\prime}\right) \cdot \frac{1}{2} \sum_{k=-\infty}^{+\infty}\left[H_{k}\left(\overrightarrow{\boldsymbol{\Theta}}^{\prime}\right) H_{k}^{*}(\overrightarrow{\boldsymbol{\Theta}})+H_{-k}\left(\overrightarrow{\boldsymbol{\Theta}}^{\prime}\right) H_{-k}^{*}(\overrightarrow{\boldsymbol{\Theta}})\right] \tag{5.66}
\end{equation*}
$$

Choosing $\varepsilon=\frac{t_{\varepsilon}}{\boldsymbol{t}_{\boldsymbol{E}}(\overrightarrow{\boldsymbol{x}}, t, \bar{\Omega})}$ the exchange function $\widetilde{W}_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}$ may be understood in the dependencies

$$
\begin{equation*}
\widetilde{W}_{\boldsymbol{t}_{\epsilon}}=\widetilde{W}_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}\left(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\Omega}} \cdot \overrightarrow{\boldsymbol{\Omega}}^{\prime}, \overrightarrow{\boldsymbol{\Theta}} \cdot \overrightarrow{\boldsymbol{\Theta}}^{\prime}\right) \tag{5.67}
\end{equation*}
$$

and

$$
\begin{equation*}
\bar{W}_{\boldsymbol{t}_{\epsilon}}\left(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\Omega}}, \overrightarrow{\boldsymbol{\Theta}}, \overrightarrow{\boldsymbol{\Omega}}^{\prime}, \overrightarrow{\boldsymbol{\Theta}}^{\prime}\right) \approx \widetilde{W}_{\boldsymbol{t}_{\epsilon}}\left(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\Omega}} \cdot \overrightarrow{\boldsymbol{\Omega}}^{\prime}, \overrightarrow{\boldsymbol{\Theta}} \cdot \overrightarrow{\boldsymbol{\Theta}}^{\prime}\right) \tag{5.68}
\end{equation*}
$$

is given, too. $\Longrightarrow$

$$
\begin{gather*}
\bar{f}_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}(\overrightarrow{\mathbf{x}}, \overrightarrow{\boldsymbol{\Omega}}, \overrightarrow{\boldsymbol{\Theta}}, t)=\int_{2 \pi} \int_{4 \pi} \widetilde{W}_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}\left(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\Omega}}, \overrightarrow{\boldsymbol{\Theta}}, \overrightarrow{\boldsymbol{\Omega}}^{\prime}, \overrightarrow{\boldsymbol{\Theta}}^{\prime}\right) \bar{f}_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}\left(\overrightarrow{\mathbf{x}}-t_{\varepsilon} \cdot \vec{v}^{\prime} \overrightarrow{\boldsymbol{\Omega}}^{\prime} \times \overrightarrow{\boldsymbol{\Theta}}^{\prime}, \overrightarrow{\boldsymbol{\Omega}}^{\prime}, \overrightarrow{\boldsymbol{\Theta}}^{\prime}, t-t_{\boldsymbol{\epsilon}}\right) d \overrightarrow{\boldsymbol{\Omega}}^{\prime} d \overrightarrow{\boldsymbol{\Theta}}^{\prime}  \tag{5.69}\\
\bar{v}^{\prime}=\bar{v}^{\prime}\left(\overrightarrow{\mathbf{x}}, \overrightarrow{\boldsymbol{\Omega}}^{\prime}, \overrightarrow{\boldsymbol{\Theta}}^{\prime}, t\right)=\bar{\omega}^{\prime}\left(\overrightarrow{\mathbf{x}}, \overrightarrow{\boldsymbol{\Omega}}^{\prime}, \overrightarrow{\boldsymbol{\Theta}}^{\prime}, t\right) \cdot \vec{r}^{\prime}\left(\overrightarrow{\mathbf{x}}, \overrightarrow{\boldsymbol{\Omega}}^{\prime}, \overrightarrow{\boldsymbol{\Theta}}^{\prime}, t\right)
\end{gather*}
$$

## Chapter 6

## Verification that the Stochastic Aerosol Motions occur through turbulently moving Continua

### 6.1 The Relationship between Stochastic Aerosol Transport and known Fluid Dynamics.

## Introduction

In this last chapter, it will be shown that the stochastic aerosol transport in terms of an ensemble theory can indeed be assigned to turbulent continua characterized by fluid elements

$$
\overrightarrow{\mathrm{v}}_{t_{\epsilon}}=\vec{\omega}_{t_{\epsilon}} \times \overrightarrow{\mathbf{r}}_{t_{\epsilon}}
$$

. Separation of vectors $\overrightarrow{\boldsymbol{\omega}}$ and $\overrightarrow{\boldsymbol{r}}$ in magnitude and direction corresponding to (5.2) is not needed now. As with aerosol transport, the term

$$
\begin{equation*}
\lim _{t_{\epsilon} \rightarrow 0} \frac{\int_{\overrightarrow{\mathbf{r}}} \int_{\overrightarrow{\boldsymbol{\omega}}} W_{t_{\epsilon}} f_{t_{\epsilon}}^{\prime} d \overrightarrow{\boldsymbol{\omega}}^{\prime} d \overrightarrow{\mathbf{r}}^{\prime}-f_{t_{\epsilon}}}{t_{\epsilon}}=F(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\omega}}, \overrightarrow{\mathbf{r}}) \tag{6.1}
\end{equation*}
$$

turns out to be the key to the problem. F is the exchange term of $\overrightarrow{\mathbf{r}}$ and $\overrightarrow{\boldsymbol{\omega}}$ not integrated out with respect to vector amounts as in (5.22).

Turbulently moved one phase fluids are examined considering statistical deliberations and its deterministic counterparts. That a linking of deterministic and stochastic theory may be available and further more that out of this connection additionally important (sometimes otherwise not known) relations arise for deterministic formulations, is shown in the following.

## The Transition: Stochastic Theory $\longrightarrow$ Deterministic Theory

Every space-time-point $(\overrightarrow{\boldsymbol{x}}, t)$ is assigned a continuously differentiable fluid element distribution over the motion amounts $\vec{\omega}_{t_{\epsilon}}$ and $\overrightarrow{\mathbf{r}}_{t_{\epsilon}}$ according to

$$
\begin{equation*}
f_{\boldsymbol{t}_{\epsilon}}=f_{t_{\epsilon}}(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\omega}}, \overrightarrow{\mathbf{r}}) . \tag{6.2}
\end{equation*}
$$

For indexed functions with $\boldsymbol{t}_{\boldsymbol{\epsilon}}$, it is automatically assumed that the dependent motion quantities $(\overrightarrow{\boldsymbol{\omega}}, \overrightarrow{\mathbf{r}})$ are assigned to a $\boldsymbol{t}_{\boldsymbol{\epsilon}}$-measurement accuracy. The indexing of the motion quantities may be omitted in the functions if the functions are accordingly indexed.

After an execution of a $\lim \boldsymbol{t}_{\boldsymbol{\epsilon}} \rightarrow 0$ process, such as

$$
\begin{equation*}
\lim _{t_{\epsilon} \rightarrow 0} f_{t_{\epsilon}}(\vec{x}, t, \vec{\omega}, \vec{r})=f(\overrightarrow{\boldsymbol{x}}, t, \vec{\omega}, \vec{r}) \tag{6.3}
\end{equation*}
$$

f and $(\overrightarrow{\boldsymbol{\omega}}, \overrightarrow{\mathbf{r}})$ are understood as results of an exact measuring process.
The change of motion quantities in point $(\overrightarrow{\boldsymbol{x}}, t)$

$$
\left(\overrightarrow{\boldsymbol{\omega}}_{t_{\epsilon}}^{\prime}\left(\vec{x}-\Delta \overrightarrow{\boldsymbol{x}}, t-t_{\epsilon}\right), \overrightarrow{\boldsymbol{r}}_{t_{\epsilon}}^{\prime}\left(\overrightarrow{\boldsymbol{x}}-\Delta \overrightarrow{\boldsymbol{x}}, t-t_{\epsilon}\right)\right) \longrightarrow\left(\overrightarrow{\boldsymbol{\omega}}_{t_{\epsilon}}(\overrightarrow{\boldsymbol{x}}, t), \vec{r}_{t_{\epsilon}}(\overrightarrow{\boldsymbol{x}}, t)\right)
$$

is controlled by the transition probability density $W_{t_{\epsilon}}=W_{t_{\epsilon}}\left(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\omega}}, \overrightarrow{\boldsymbol{r}}, \overrightarrow{\boldsymbol{\omega}}^{\prime}, \overrightarrow{\boldsymbol{r}}^{\prime}\right) .{ }^{1}$ with

$$
\begin{align*}
\lim _{t_{\epsilon} \rightarrow 0} W_{t_{\epsilon}} & =\delta\left(\overrightarrow{\boldsymbol{\omega}}, \overrightarrow{\mathbf{r}} ; \overrightarrow{\boldsymbol{\omega}}^{\prime}, \overrightarrow{\mathbf{r}}^{\prime}\right) \\
f_{t_{\epsilon}}(\overrightarrow{\mathbf{x}}, t, \overrightarrow{\boldsymbol{\omega}}, \overrightarrow{\mathbf{r}}) & =\int_{\overrightarrow{\mathbf{r}}} \int_{\vec{\omega}} W_{t_{\epsilon}}\left(\overrightarrow{\mathbf{x}}, t, \overrightarrow{\boldsymbol{\omega}}, \overrightarrow{\mathbf{r}}, \overrightarrow{\boldsymbol{\omega}}^{\prime}, \overrightarrow{\mathbf{r}}^{\prime}\right) \cdot f_{t_{\epsilon}}\left(\overrightarrow{\mathbf{x}}-\Delta \overrightarrow{\mathbf{x}}, t-t_{\epsilon}, \overrightarrow{\boldsymbol{\omega}}^{\prime}, \overrightarrow{\mathbf{r}}^{\prime}\right) d \overrightarrow{\boldsymbol{\omega}}^{\prime} d \overrightarrow{\mathbf{r}}^{\prime}  \tag{6.4}\\
\Delta \overrightarrow{\mathbf{x}} & =t_{\epsilon} \cdot \overrightarrow{\boldsymbol{\omega}}^{\prime} \times \overrightarrow{\mathbf{r}}^{\prime}
\end{align*}
$$

These equations characterize stochastic turbulence of the continuum in the frame of an ensemble theory and represent a Markov Process with natural causality. (This is a definition of the author.)
$f_{t_{\epsilon}}$ is developed in (6.4) until the 1st order around $(\overrightarrow{\mathrm{x}}, t) \Longrightarrow$

$$
\begin{equation*}
f_{t_{\epsilon}}\left(\overrightarrow{\mathbf{x}}-\triangle \overrightarrow{\mathbf{x}}, t-t_{\epsilon}, \overrightarrow{\boldsymbol{\omega}}^{\prime}, \overrightarrow{\mathbf{r}}^{\prime}\right)=f_{t_{\epsilon}}\left(\overrightarrow{\mathbf{x}}, t, \overrightarrow{\boldsymbol{\omega}}^{\prime}, \overrightarrow{\mathbf{r}}^{\prime}\right)-\frac{\partial f_{t_{\epsilon}}^{\prime}}{\partial t} \cdot t_{\epsilon}-\triangle \overrightarrow{\mathbf{x}} \cdot \overrightarrow{\boldsymbol{\nabla}} f_{t_{\epsilon}}\left(\overrightarrow{\mathbf{x}}, t, \overrightarrow{\boldsymbol{\omega}}^{\prime}, \overrightarrow{\mathbf{r}}^{\prime}\right)+\boldsymbol{O}\left(t_{\epsilon}^{2}\right) \tag{6.5}
\end{equation*}
$$

with $f_{t_{\epsilon}}^{\prime}=f_{t_{\epsilon}}\left(\overrightarrow{\mathbf{x}}, t, \overrightarrow{\boldsymbol{\omega}}^{\prime}, \overrightarrow{\mathbf{r}}^{\prime}\right)$ and one obtains

$$
\begin{equation*}
\int_{\overrightarrow{\mathbf{r}}} \int_{\overrightarrow{\boldsymbol{\omega}}} W_{t_{\epsilon}}\left[\frac{\partial f_{t_{\epsilon}}^{\prime}}{\partial t}+\overrightarrow{\boldsymbol{\omega}^{\prime}} \times \overrightarrow{\mathbf{r}}^{\prime} \cdot \overrightarrow{\boldsymbol{\nabla}} f_{t_{\epsilon}}^{\prime}\right] d \overrightarrow{\boldsymbol{\omega}}^{\prime} d \overrightarrow{\mathbf{r}}^{\prime}+\boldsymbol{O}\left(t_{\epsilon}{ }^{2}\right)=\frac{\int_{\overrightarrow{\mathbf{r}}} \int_{\overrightarrow{\boldsymbol{\omega}}} W_{t_{\epsilon}} f_{t_{\epsilon}}^{\prime} d \overrightarrow{\boldsymbol{\omega}}^{\prime} d \overrightarrow{\mathbf{r}}^{\prime}-f_{t_{\epsilon}}}{t_{\epsilon}} \tag{6.6}
\end{equation*}
$$

$\lim \boldsymbol{t}_{\boldsymbol{\epsilon}} \rightarrow 0$ applied to (6.6) leads to

$$
\begin{equation*}
\frac{\partial f}{\partial t}+\overrightarrow{\boldsymbol{\omega}} \times \overrightarrow{\mathbf{r}} \cdot \overrightarrow{\boldsymbol{\nabla}} f=\lim _{t_{\epsilon} \rightarrow 0} \frac{\int_{\overrightarrow{\mathbf{r}}} \int_{\overrightarrow{\boldsymbol{\omega}}} W_{t_{\epsilon}} f_{t_{\epsilon}}^{\prime} d \overrightarrow{\boldsymbol{\omega}}^{\prime} d \overrightarrow{\mathbf{r}}^{\prime}-f_{t_{\epsilon}}}{t_{\epsilon}} \tag{6.7}
\end{equation*}
$$

The right side must contain the characteristics of the turbulent fluid.

$$
\begin{equation*}
\lim _{t_{\epsilon} \rightarrow 0} \frac{\int_{\overrightarrow{\mathbf{r}}} \int_{\overrightarrow{\boldsymbol{\omega}}} W_{t_{\epsilon}} f_{t_{\epsilon}}^{\prime} d \overrightarrow{\boldsymbol{\omega}}^{\prime} d \overrightarrow{\mathbf{r}}^{\prime}-f_{t_{\epsilon}}}{t_{\epsilon}}=F(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\omega}}, \overrightarrow{\mathbf{r}}) \tag{6.8}
\end{equation*}
$$

$F$ has to be chosen such, that the deterministic vortex equations result under the influence of the assumed acceleration field. Thus one obtains

$$
\begin{equation*}
\frac{\partial f}{\partial t}+\overrightarrow{\boldsymbol{\omega}} \times \overrightarrow{\mathbf{r}} \cdot \vec{\nabla} f=F(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\omega}}, \overrightarrow{\mathbf{r}}) \tag{6.9}
\end{equation*}
$$

If we restrict ourselves to a single system of the ensemble with the identifier $\nu$ in a space-time point $(\overrightarrow{\boldsymbol{x}}, t)$, which has exactly these movement sizes $\overrightarrow{\boldsymbol{\omega}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}$ and $\overrightarrow{\mathbf{r}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}$ in this system, then the distribution function f degenerates to a $\delta$-function with respect to these movement sizes. $\overrightarrow{\boldsymbol{\omega}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}$ and $\overrightarrow{\mathbf{r}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}$ are not vector functions but constant vectors in $(\overrightarrow{\boldsymbol{x}}, t)$, whereas $\overrightarrow{\boldsymbol{\omega}}(\overrightarrow{\boldsymbol{x}}, t)$ and $\overrightarrow{\mathbf{r}}(\overrightarrow{\boldsymbol{x}}, t)$ represent spatiotemporal fields in dependence of $(\overrightarrow{\boldsymbol{x}}, t)$.

$$
\begin{equation*}
f(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\omega}}, \overrightarrow{\mathbf{r}}) \rightarrow \delta\left(\overrightarrow{\boldsymbol{\omega}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}, \overrightarrow{\mathbf{r}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)} ; \overrightarrow{\boldsymbol{\omega}}, \overrightarrow{\mathbf{r}}\right) \tag{6.10}
\end{equation*}
$$

and

$$
\begin{equation*}
F(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\omega}}, \overrightarrow{\mathbf{r}}) \rightarrow \frac{1}{2}\left[\frac{\overrightarrow{\boldsymbol{\omega}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}}{\omega_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}^{2}} \cdot(\overrightarrow{\boldsymbol{\nabla}} \times \overrightarrow{\mathbf{q}})_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}\right] \delta\left(\overrightarrow{\boldsymbol{\omega}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}, \overrightarrow{\mathbf{r}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)} ; \overrightarrow{\boldsymbol{\omega}}, \overrightarrow{\mathbf{r}}\right), \tag{6.11}
\end{equation*}
$$

as will be shown in the following.

[^6]The equation for stochastic propagation in terms of an ensemble theory thus degenerates to the following equation, from now on called key equation.

$$
\begin{equation*}
\left(\frac{\partial}{\partial t}+\overrightarrow{\boldsymbol{\omega}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)} \times \overrightarrow{\mathbf{r}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)} \cdot \overrightarrow{\boldsymbol{\nabla}}-\frac{1}{2}\left[\frac{\overrightarrow{\boldsymbol{\omega}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}}{\omega_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}^{2}} \cdot(\overrightarrow{\boldsymbol{\nabla}} \times \overrightarrow{\mathbf{q}})_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}\right]\right) \delta=\mathbf{0} . \tag{6.12}
\end{equation*}
$$

For this $\delta$-function applies

$$
\begin{equation*}
\int_{\overrightarrow{\mathbf{r}}} \int_{\overrightarrow{\boldsymbol{\omega}}} \delta\left(\overrightarrow{\boldsymbol{\omega}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}, \overrightarrow{\mathbf{r}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)} ; \overrightarrow{\boldsymbol{\omega}}, \overrightarrow{\mathbf{r}}\right) d \overrightarrow{\boldsymbol{\omega}} d \overrightarrow{\mathbf{r}}=\mathbf{1} \tag{6.13}
\end{equation*}
$$

Definition of the operator $\boldsymbol{\Xi}[. .$.$] :$
From the vector $\overrightarrow{\mathbf{A}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}$ respectively the scalar function value $\mathbf{f}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}$ which is defined in the space-time point $(\overrightarrow{\boldsymbol{x}}, t)$ of the system $\nu$, a vector function or a scalar function is obtained by the operator $\boldsymbol{\Xi}$, if a corresponding field exists around the point $(\overrightarrow{\boldsymbol{x}}, t)$

$$
\begin{equation*}
\boldsymbol{\Xi}\left[\overrightarrow{\mathbf{A}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}\right]=\overrightarrow{\mathbf{A}}(\overrightarrow{\boldsymbol{x}}, t), \quad \boldsymbol{\Xi}\left[\mathbf{f}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}\right]=\mathbf{f}(\overrightarrow{\boldsymbol{x}}, t) \tag{6.14}
\end{equation*}
$$

The Operator $\boldsymbol{\Xi}[\ldots]$ brings this functionality to "life".
Accordingly the following relationships are noted:

$$
\begin{align*}
& \boldsymbol{\Xi}\left[\overrightarrow{\boldsymbol{\omega}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}\right]=\overrightarrow{\boldsymbol{\omega}}(\overrightarrow{\boldsymbol{x}}, t) \\
& \boldsymbol{\Xi}\left[\overrightarrow{\mathbf{r}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}\right]=\overrightarrow{\mathbf{r}}(\overrightarrow{\boldsymbol{x}}, t)  \tag{6.15}\\
& \boldsymbol{\Xi}\left[\boldsymbol{\omega}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}^{2} \overrightarrow{\mathbf{r}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}\right]=\boldsymbol{\omega}^{2}(\overrightarrow{\mathbf{x}}, t) \overrightarrow{\mathbf{r}}(\overrightarrow{\mathbf{x}}, t) \\
& \boldsymbol{\Xi}\left(\frac{1}{2}\left[\frac{\overrightarrow{\boldsymbol{\omega}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}}{\omega_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}^{2}} \cdot(\overrightarrow{\boldsymbol{\nabla}} \times \overrightarrow{\mathbf{q}})_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}\right]\right)=\frac{1}{2} \frac{\overrightarrow{\boldsymbol{\omega}}(\overrightarrow{\boldsymbol{x}}, t)}{\omega^{2}(\overrightarrow{\boldsymbol{x}}, t)} \cdot \overrightarrow{\boldsymbol{\nabla}} \times \overrightarrow{\mathbf{q}}(\overrightarrow{\boldsymbol{x}}, t)
\end{align*}
$$

## Deterministic Equations of Turbulence

From the general momentum equation

$$
\begin{equation*}
\frac{\partial \overrightarrow{\mathbf{v}}}{\partial t}+(\overrightarrow{\mathbf{v}} \cdot \overrightarrow{\boldsymbol{\nabla}}) \overrightarrow{\mathbf{v}}=\overrightarrow{\boldsymbol{q}} \tag{6.16}
\end{equation*}
$$

the special momentum equations, Navier-Stokes equations, are questioned by the author for the description of turbulence phenomena for several reasons,
the vortex equation ${ }^{2}$ may be developed using the $\vec{\nabla} \times$-operator

$$
\begin{equation*}
\frac{\partial}{\partial t} \overrightarrow{\boldsymbol{\omega}}-\overrightarrow{\boldsymbol{\nabla}} \times(\overrightarrow{\mathbf{v}} \times \overrightarrow{\boldsymbol{\omega}})-\frac{1}{2} \overrightarrow{\boldsymbol{\nabla}} \times \overrightarrow{\boldsymbol{q}}=0 \tag{6.17}
\end{equation*}
$$

The relations between deterministic and stochastic description is established when the known deterministic vortex equation can be reconstructed from an associated stochastic equation of the ensemble theory. In the following the method is presented developing the dual pair of deterministic vector equations from the key equation (6.12).

$$
\left(\frac{\partial}{\partial t}+\overrightarrow{\boldsymbol{\omega}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)} \times \overrightarrow{\mathbf{r}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)} \cdot \overrightarrow{\boldsymbol{\nabla}}-\frac{1}{2}\left[\frac{\overrightarrow{\boldsymbol{\omega}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}}{\omega_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}^{2}} \cdot(\overrightarrow{\boldsymbol{\nabla}} \times \overrightarrow{\mathbf{q}})_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}\right]\right) \delta=\mathbf{0}
$$

In this situation the vectors may be pushed before and after the differential operators. The Term

$$
\begin{equation*}
\frac{1}{2}\left[\frac{\overrightarrow{\boldsymbol{\omega}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}}{\omega_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}^{2}} \cdot(\overrightarrow{\boldsymbol{\nabla}} \times \overrightarrow{\mathbf{q}})_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}\right] \delta \tag{6.18}
\end{equation*}
$$

[^7]guarantees the finding of equation (6.17) and its dual one. It is
\[

$$
\begin{equation*}
\overrightarrow{\mathbf{v}} \perp \vec{\omega} \perp \overrightarrow{\mathbf{r}} . \tag{6.19}
\end{equation*}
$$

\]

and setting

$$
\begin{equation*}
\overrightarrow{\mathbf{a}}=\overrightarrow{\mathbf{v}} \times \vec{\omega} \tag{6.20}
\end{equation*}
$$

this results in

$$
\begin{equation*}
\overrightarrow{\mathbf{r}} \| \overrightarrow{\mathbf{a}} \tag{6.21}
\end{equation*}
$$

Such $\overrightarrow{\mathbf{a}}$ and $\overrightarrow{\mathbf{r}}$ are linked as follows ${ }^{3}$

$$
\begin{equation*}
\overrightarrow{\mathbf{r}}=\frac{\overrightarrow{\mathbf{a}}}{\omega^{2}} \tag{6.22}
\end{equation*}
$$

$\Longrightarrow$

$$
\text { with } \quad \delta=\delta\left(\overrightarrow{\boldsymbol{\omega}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}, \overrightarrow{\mathbf{r}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)} ; \overrightarrow{\boldsymbol{\omega}}, \overrightarrow{\mathbf{r}}\right)
$$

$$
\overrightarrow{\boldsymbol{\omega}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)} \times \overrightarrow{\mathbf{r}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)} \cdot \overrightarrow{\boldsymbol{\nabla}} \delta=-\overrightarrow{\mathbf{r}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)} \times \overrightarrow{\boldsymbol{\omega}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)} \cdot \overrightarrow{\boldsymbol{\nabla}} \delta=-\overrightarrow{\boldsymbol{\omega}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)} \cdot \overrightarrow{\boldsymbol{\nabla}} \times \overrightarrow{\mathbf{r}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)} \delta=-\frac{\overrightarrow{\boldsymbol{\omega}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}}{\omega_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}^{2}} \cdot \overrightarrow{\boldsymbol{\nabla}} \times \overrightarrow{\mathbf{a}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)} \delta
$$

Inserting in (6.12) gives

$$
\begin{align*}
& \frac{\partial}{\partial t}\left(\frac{\overrightarrow{\boldsymbol{\omega}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)} \cdot \overrightarrow{\boldsymbol{\omega}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}}{\omega_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}^{2}} \delta\right)-\frac{\overrightarrow{\boldsymbol{\omega}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}}{\omega_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}^{2}} \cdot \overrightarrow{\boldsymbol{\nabla}} \times\left(\overrightarrow{\mathbf{a}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)} \delta\right)-\frac{1}{2}\left[\frac{\overrightarrow{\boldsymbol{\omega}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}}{\omega_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}^{2}} \cdot(\overrightarrow{\boldsymbol{\nabla}} \times \overrightarrow{\mathbf{q}})_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}\right] \delta=0 \\
& \Longrightarrow \frac{\overrightarrow{\boldsymbol{\omega}}_{(\vec{x}}}{\omega_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}^{2}} \cdot\left(\frac{\partial}{\partial t}\left(\overrightarrow{\boldsymbol{\omega}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)} \delta\right)-\overrightarrow{\boldsymbol{\nabla}} \times\left(\overrightarrow{\left.\left.\mathbf{a}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)} \delta\right)-\frac{1}{2}\left[\cdot(\overrightarrow{\boldsymbol{\nabla}} \times \overrightarrow{\mathbf{q}})_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}\right] \delta\right)=0}\right.\right.  \tag{6.23}\\
& \Longrightarrow \frac{\partial}{\partial t}\left(\overrightarrow{\boldsymbol{\omega}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)} \delta\right)-\overrightarrow{\boldsymbol{\nabla}} \times\left(\overrightarrow{\mathbf{a}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)} \delta\right)-\frac{1}{2}\left[\cdot(\overrightarrow{\boldsymbol{\nabla}} \times \overrightarrow{\mathbf{q}})_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}\right] \delta=0
\end{align*}
$$

and

$$
\begin{equation*}
\boldsymbol{\Xi}\left[\int_{\overrightarrow{\mathbf{r}}} \int_{\overrightarrow{\boldsymbol{\omega}}}\left[\frac{\partial}{\partial t}\left(\overrightarrow{\boldsymbol{\omega}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)} \delta\right)-\overrightarrow{\boldsymbol{\nabla}} \times\left(\overrightarrow{\mathbf{a}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)} \delta\right)-\frac{1}{2}\left[(\overrightarrow{\boldsymbol{\nabla}} \times \overrightarrow{\mathbf{q}})_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}\right] \delta\right] d \overrightarrow{\boldsymbol{\omega}} d \overrightarrow{\mathbf{r}}=\mathbf{0}\right. \tag{6.24}
\end{equation*}
$$

is obtained and since integration and differentiation are interchangeable in order, it follows that

$$
\begin{equation*}
\left[\frac{\partial}{\partial t} \boldsymbol{\Xi}\left[\overrightarrow{\boldsymbol{\omega}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}\right]-\vec{\nabla} \times \boldsymbol{\Xi}\left[\overrightarrow{\mathrm{a}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}\right]-\frac{1}{2} \overrightarrow{\boldsymbol{\nabla}} \times \boldsymbol{\Xi}\left[\overrightarrow{\mathbf{q}}_{(\overrightarrow{\boldsymbol{x}}, t, \nu)}\right]=0\right. \tag{6.25}
\end{equation*}
$$

Now we have the first of the dual turbulence equations

$$
\begin{equation*}
\frac{\partial}{\partial t} \overrightarrow{\boldsymbol{\omega}}-\overrightarrow{\boldsymbol{\nabla}} \times \overrightarrow{\mathbf{a}}-\frac{1}{2} \overrightarrow{\boldsymbol{\nabla}} \times \overrightarrow{\mathbf{q}}=0 \tag{6.26}
\end{equation*}
$$

, accordingly

$$
\frac{\partial}{\partial t} \overrightarrow{\boldsymbol{\omega}}-\overrightarrow{\boldsymbol{\nabla}} \times(\overrightarrow{\mathbf{v}} \times \overrightarrow{\boldsymbol{\omega}})-\frac{1}{2} \overrightarrow{\boldsymbol{\nabla}} \times \overrightarrow{\boldsymbol{q}}=0
$$

Hereby the connection of stochastics and deterministics is achieved. From the key-equation above a second equation, the dual one, may be derived.
Back to the initial equation (6.12)

$$
\left(\frac{\partial}{\partial t}+\overrightarrow{\boldsymbol{\omega}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)} \times \overrightarrow{\mathbf{r}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)} \cdot \overrightarrow{\boldsymbol{\nabla}}-\frac{1}{2}\left[\frac{\overrightarrow{\boldsymbol{\omega}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}}{\omega_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}} \cdot(\overrightarrow{\boldsymbol{\nabla}} \times \overrightarrow{\mathbf{q}})_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}\right]\right) \delta=\mathbf{0}
$$

Simple conversions give

$$
\begin{array}{r}
\frac{\partial}{\partial t}\left(\overrightarrow{\mathbf{r}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)} \cdot \frac{\overrightarrow{\mathbf{r}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}}{r_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}^{2}} \delta\right)+\overrightarrow{\mathbf{r}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)} \cdot \overrightarrow{\boldsymbol{\nabla}} \times\left(\overrightarrow{\boldsymbol{\omega}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)} \delta\right)-\frac{\overrightarrow{\mathbf{r}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)} \cdot \overrightarrow{\mathbf{r}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}}{r_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}^{2}} \frac{1}{2}\left[\frac{\overrightarrow{\boldsymbol{\omega}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}}{\omega_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}^{2}} \cdot(\overrightarrow{\boldsymbol{\nabla}} \times \overrightarrow{\mathbf{q}})_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}\right] \delta=0  \tag{6.27}\\
\quad \longrightarrow \overrightarrow{\boldsymbol{r}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}\left[\frac{\partial}{\partial t} \frac{\overrightarrow{\mathbf{r}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}}{r_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}^{2}} \delta+\overrightarrow{\boldsymbol{\nabla}} \times\left(\overrightarrow{\boldsymbol{\omega}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)} \delta\right)-\frac{\overrightarrow{\mathbf{r}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}}{r_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}^{2}} \frac{1}{2}\left[\frac{\overrightarrow{\boldsymbol{\omega}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}}{\omega_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}^{2}} \cdot(\overrightarrow{\boldsymbol{\nabla}} \times \overrightarrow{\mathbf{q}})_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}\right] \delta\right]=0
\end{array}
$$

[^8]Using the curvature vector field of the fluid trajectories $\overrightarrow{\mathbf{b}}=\frac{\overrightarrow{\mathbf{r}}}{r^{2}}$ the equation is written

$$
\begin{equation*}
\frac{\partial}{\partial t}\left(\overrightarrow{\mathbf{b}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)} \delta\right)+\overrightarrow{\boldsymbol{\nabla}} \times\left(\overrightarrow{\boldsymbol{\omega}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)} \delta\right)-\frac{1}{2} \overrightarrow{\mathbf{b}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)} \frac{\overrightarrow{\boldsymbol{\omega}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}}{\omega_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}^{2}} \cdot(\overrightarrow{\boldsymbol{\nabla}} \times \overrightarrow{\mathbf{q}})_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)} \delta=0 \tag{6.28}
\end{equation*}
$$

and applying the operators $\boldsymbol{\Xi}$ arises

$$
\begin{equation*}
\boldsymbol{\Xi}\left[\int_{\overrightarrow{\mathbf{r}}} \int_{\overrightarrow{\boldsymbol{\omega}}}\left[\frac{\partial}{\partial t}\left(\overrightarrow{\mathbf{b}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)} \delta\right)+\overrightarrow{\boldsymbol{\nabla}} \times\left(\overrightarrow{\boldsymbol{\omega}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)} \delta\right)-\frac{1}{2} \overrightarrow{\mathbf{b}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)} \frac{\overrightarrow{\boldsymbol{\omega}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}}{\omega_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}^{2}} \cdot(\overrightarrow{\boldsymbol{\nabla}} \times \overrightarrow{\mathbf{q}})_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)} \delta\right] d \overrightarrow{\boldsymbol{\omega}} d \overrightarrow{\mathbf{r}}=\mathbf{0}\right. \tag{6.29}
\end{equation*}
$$

respectively

$$
\begin{equation*}
\frac{\partial}{\partial t} \boldsymbol{\Xi}\left[\overrightarrow{\mathbf{b}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}\right]+\overrightarrow{\boldsymbol{\nabla}} \times \boldsymbol{\Xi}\left[\overrightarrow{\boldsymbol{\omega}}_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}\right]-\frac{1}{2} \boldsymbol{\Xi}\left[\overrightarrow{\mathbf{b}} \cdot\left(\frac{\overrightarrow{\boldsymbol{\omega}}}{\omega^{2}} \cdot(\overrightarrow{\boldsymbol{\nabla}} \times \overrightarrow{\mathbf{q}})_{(\overrightarrow{\boldsymbol{x}}, t ; \nu)}\right)\right]=0 \tag{6.30}
\end{equation*}
$$

Thus, the second of the dual turbulence equations is obtained

$$
\begin{equation*}
\frac{\partial}{\partial t} \overrightarrow{\mathbf{b}}+\overrightarrow{\boldsymbol{\nabla}} \times \overrightarrow{\boldsymbol{\omega}}-\frac{1}{2} \overrightarrow{\mathbf{b}}\left[\frac{\overrightarrow{\boldsymbol{\omega}}}{\omega^{2}} \cdot \overrightarrow{\boldsymbol{\nabla}} \times \overrightarrow{\mathbf{q}}\right]=0 \tag{6.31}
\end{equation*}
$$

Overall, this results in the dual system of equations

$$
\begin{align*}
& \frac{\partial}{\partial t} \overrightarrow{\boldsymbol{\omega}}-\overrightarrow{\boldsymbol{\nabla}} \times \overrightarrow{\mathbf{a}}-\frac{1}{2} \overrightarrow{\boldsymbol{\nabla}} \times \overrightarrow{\mathbf{q}}=0  \tag{6.32}\\
& \frac{\partial}{\partial t} \overrightarrow{\mathbf{b}}+\overrightarrow{\boldsymbol{\nabla}} \times \overrightarrow{\boldsymbol{\omega}}-\frac{1}{2} \overrightarrow{\mathbf{b}}\left[\frac{\overrightarrow{\boldsymbol{\omega}}}{\omega^{2}} \cdot \overrightarrow{\boldsymbol{\nabla}} \times \overrightarrow{\mathbf{q}}\right]=0 \\
& \overrightarrow{\mathbf{v}}=\overrightarrow{\boldsymbol{\omega}} \times \frac{\overrightarrow{\mathbf{b}}}{b^{2}}, \overrightarrow{\mathbf{a}}=\overrightarrow{\mathbf{v}} \times \overrightarrow{\boldsymbol{\omega}}
\end{align*}
$$

The term

$$
-\frac{1}{2} \overrightarrow{\mathbf{b}}\left[\frac{\overrightarrow{\boldsymbol{\omega}}}{\omega^{2}} \cdot \overrightarrow{\boldsymbol{\nabla}} \times \overrightarrow{\mathbf{q}}\right]
$$

leads to removable singularities in space-time-points $(\overrightarrow{\mathbf{x}}, t)$ if $\overrightarrow{\boldsymbol{\omega}}=0$ and $\overrightarrow{\mathbf{b}}=\mathbf{0}$ occur in the fluid-element trajectories. In this case the whole term is calculated from its surroundings. The same shall apply for the calculation of the velocity $\overrightarrow{\mathrm{v}}$.
These matters are to be discussed in connection with a statement of a complete system of equations of deterministic turbulence, which will be done in another paper.

## Chapter 7

## Summary and Outlook

Aerosol motions in a turbulently moving continuum are studied assuming that they accurately describe the trajectories of individual fluid elements due to their size and weight. These movements, which are actually deterministic, were considered stochastically in the sense of an ensemble theory. After the consequent derivation of the aerosol transport equations, two coefficients $\bar{v}(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\Omega}}, \overrightarrow{\boldsymbol{\Theta}})$ and $t_{E}(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\Omega}}, \overrightarrow{\boldsymbol{\Theta}})$, which are still very complicated in their dependencies, remain. Simplifying model assumptions can lead to correspondingly simplified coefficients. Furthermore, the function development can be terminated with the first order, since, as shown, such an approximation approaches asymptotically the exact solution with the distance against assumed sources and the time.

$$
\begin{gathered}
\frac{\partial \bar{f}}{\partial t}+\bar{v} \overrightarrow{\boldsymbol{\Omega}} \times \overrightarrow{\boldsymbol{\Theta}} \cdot \nabla \bar{f}=\frac{-1}{t_{E}} \sum_{l=0}^{+\infty} \sum_{m=-l}^{+l} \sum_{k=-\infty}^{+\infty} \frac{l(l+1)}{2} \bar{f}_{l m k}(\overrightarrow{\mathbf{x}}, t) P_{l m}(\overrightarrow{\boldsymbol{\Omega}}) H_{k}(\overrightarrow{\boldsymbol{\Theta}}) \\
\bar{v}=\bar{v}(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\Omega}}, \overrightarrow{\boldsymbol{\Theta}}), t_{E}=t_{E}(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\Omega}}, \overrightarrow{\boldsymbol{\Theta}}) \\
\bar{f}_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}(\overrightarrow{\mathbf{x}}, \overrightarrow{\boldsymbol{\Omega}}, \overrightarrow{\boldsymbol{\Theta}}, t)=\int_{2 \pi} \int_{4 \pi} \widetilde{W}_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}\left(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\Omega}}, \overrightarrow{\boldsymbol{\Theta}}, \overrightarrow{\boldsymbol{\Omega}}^{\prime}, \overrightarrow{\boldsymbol{\Theta}}^{\prime}\right) \bar{f}_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}\left(\overrightarrow{\mathbf{x}}-t_{\varepsilon} \cdot \bar{v}^{\prime} \overrightarrow{\boldsymbol{\Omega}}^{\prime} \times \overrightarrow{\boldsymbol{\Theta}}^{\prime}, \overrightarrow{\boldsymbol{\Omega}}^{\prime}, \overrightarrow{\boldsymbol{\Theta}}^{\prime}, t-t_{\boldsymbol{\epsilon}}\right) d \overrightarrow{\boldsymbol{\Omega}}^{\prime} d \overrightarrow{\boldsymbol{\Theta}}^{\prime} \\
\widetilde{W}_{\boldsymbol{t}_{\boldsymbol{\epsilon}}}\left(\overrightarrow{\boldsymbol{\Omega}} \cdot \overrightarrow{\boldsymbol{\Omega}}^{\prime}, \overrightarrow{\boldsymbol{\Theta}} \cdot \overrightarrow{\boldsymbol{\Theta}}^{\prime}\right)=\sum_{l=0}^{+\infty} e^{\Upsilon_{l} \cdot \epsilon} \frac{2 l+1}{4 \pi} \cdot \frac{1}{2 \pi} \cdot \sum_{m=-l}^{+l} P_{l m}^{*}(\overrightarrow{\boldsymbol{\Omega}}) P_{l m}\left(\overrightarrow{\boldsymbol{\Omega}}^{\prime}\right) \cdot \frac{1}{2} \sum_{k=-\infty}^{+\infty}\left[H_{k}\left(\overrightarrow{\boldsymbol{\Theta}}^{\prime}\right) H_{k}^{*}(\overrightarrow{\boldsymbol{\Theta}})+H_{-k}\left(\overrightarrow{\boldsymbol{\Theta}}^{\prime}\right) H_{-k}^{*}(\overrightarrow{\boldsymbol{\Theta}})\right] \\
\Upsilon_{l}=-\frac{l(l+1)}{2} \zeta, \quad \zeta=\mathrm{const}
\end{gathered}
$$

The integral form of the transport equation with its explicitly formulated transition probability indicates the possibility of using Monte Carlo methods for its numerical evaluation.

It is probably relatively difficult to experimentally confirm the relationships presented. Therefore, the connection between such a stochastics in the sense of an ensemble theory and a deterministic fluid dynamics is established.

$$
\begin{gather*}
f_{\boldsymbol{t}_{\epsilon}}(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\omega}}, \overrightarrow{\boldsymbol{r}})=\iint W_{\boldsymbol{t}_{\epsilon}}\left(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\omega}}, \overrightarrow{\boldsymbol{r}}, \overrightarrow{\boldsymbol{\omega}}^{\prime}, \overrightarrow{\boldsymbol{r}}^{\prime}\right) f_{t_{\epsilon}}\left(\overrightarrow{\boldsymbol{x}}-\overrightarrow{\boldsymbol{\omega}}^{\prime} \times \overrightarrow{\boldsymbol{r}}^{\prime} \cdot t_{\epsilon}, \overrightarrow{\boldsymbol{\omega}}^{\prime}, \overrightarrow{\boldsymbol{r}}^{\prime}, t-t_{\epsilon}\right) d \overrightarrow{\boldsymbol{\omega}}^{\prime} d \overrightarrow{\boldsymbol{r}}^{\prime} \\
\Downarrow \\
\Downarrow  \tag{7.1}\\
\frac{\partial}{\partial t} \overrightarrow{\boldsymbol{\omega}}-\overrightarrow{\boldsymbol{\nabla}} \times \overrightarrow{\mathbf{a}}-\frac{1}{2} \overrightarrow{\boldsymbol{\nabla}} \times \overrightarrow{\mathbf{q}}=0 \\
\frac{\partial}{\partial t} \overrightarrow{\mathbf{b}}+\overrightarrow{\boldsymbol{\nabla}} \times \overrightarrow{\boldsymbol{\omega}}-\frac{1}{2} \overrightarrow{\mathbf{b}}\left[\frac{\overrightarrow{\boldsymbol{\omega}}}{\omega^{2}} \cdot \overrightarrow{\boldsymbol{\nabla}} \times \overrightarrow{\mathbf{q}}\right]=0 \\
\overrightarrow{\mathbf{v}}=\overrightarrow{\boldsymbol{\omega}} \times \frac{\overrightarrow{\mathbf{b}}}{b^{2}}, \overrightarrow{\mathbf{a}}=\overrightarrow{\mathbf{v}} \times \overrightarrow{\boldsymbol{\omega}}, \quad \overrightarrow{\mathbf{v}} \perp \overrightarrow{\boldsymbol{\omega}} \perp \overrightarrow{\mathbf{r}} \\
\hline
\end{gather*}
$$

The result is a dual pair of deterministic equations of turbulence. In this respect, the desired goal is achieved. However, this pair of equations is not yet complete. The completion happens in a further paper, where then the whole system of equations represents a geometrodynamics of turbulence. I.e. the whole system of equations consists of vector fields of velocities, vortex rotations, their curvature vector fields and non-conservative accelerations.

## Chapter 8

## Appendix

### 8.1 Legendre-Polynomials

The Legendre-polynomials are defined within the interval $[-1,+1]$ by

$$
\begin{equation*}
P_{n}=\frac{1}{2^{n} n!} \frac{d^{n}}{d x^{n}}\left(x^{2}-1\right)^{n}, n \in N \tag{8.1}
\end{equation*}
$$

They represent a complete orthogonal function system with

$$
\int_{-1}^{+1} P_{n}(x) P_{m}(x) d x= \begin{cases}\frac{2}{2 m+1} & \text { for } \mathrm{m}=n  \tag{8.2}\\ 0 & \text { else } .\end{cases}
$$

Every continuously differentiable function $f(x)$ defined within $[-1,+1]$ can be developed by Legendre-polynomials according to

$$
\begin{equation*}
f(x)=\sum_{l=0}^{\infty} f_{l} P_{l}(x) \tag{8.3}
\end{equation*}
$$

The $f_{l}$ are the evolution coefficients. A presentation of the $\delta-f u n c t i o n$ by Legendre-polynomials is obtained by

$$
\begin{equation*}
\delta\left(x, x^{\prime}\right)=\sum_{l=0}^{\infty} \frac{2 m+1}{2} P_{l}(x) P_{l}\left(x^{\prime}\right) . \tag{8.4}
\end{equation*}
$$

Important recurrence equations are

$$
\begin{align*}
(n+1) P_{n+1} & =(2 n+1) x P_{n}(x)-n P_{n-1}(x) \\
P_{n+1}^{\prime}(x)-x P_{n}^{\prime}(x) & =(n+1) P_{n}(x), n=0,1,2, \ldots  \tag{8.5}\\
\left(1-x^{2}\right) P_{n}^{\prime}(x) & =n P_{n-1}(x)-n x P_{n}(x) .
\end{align*}
$$

An integral representation of the Legendre-polynomials is obtained by

$$
\begin{equation*}
P_{n}(x)=\frac{1}{\pi} \int_{0}^{\pi}\left(x+\sqrt{x^{2}-1} \cos (\varphi)\right)^{n} d \varphi \tag{8.6}
\end{equation*}
$$

Owing to $\left|x+\sqrt{x^{2}-1} \cos (\theta)\right|=|\cos (\theta)+i \sin (\theta) \cos (\theta)| \leqslant 1$

$$
\begin{equation*}
\left|P_{n}(x)\right| \leqslant 1 \tag{8.7}
\end{equation*}
$$

follows. These polynomials have their maximum for $x=1$, particularly

$$
\begin{equation*}
P_{n}(1)=1 . \tag{8.8}
\end{equation*}
$$

$$
\begin{equation*}
\left.\frac{d P_{l}(x)}{d x}\right|_{1}=\frac{l(l+1)}{2} \tag{8.9}
\end{equation*}
$$

is proved by complete induction.

## Proof:

1. $P_{0}^{\prime}(1)=0$

Assumption:
$2 . P_{n}^{\prime}(1)=\frac{n(n+1)}{2}$
$\Longrightarrow$
3. $P_{n+1}^{\prime}(1)=\frac{(n+2)(n+1)}{2} \quad$ wegen $\quad(8.5) \quad P_{n+1}^{\prime}(1)-P_{n}^{\prime}(1)=(n+1) P_{n}(1) \quad$ q.e.d.

### 8.2 Spherical Harmonics

The Spherical harmonics [[7] page 224] represent a complete orthogonal, complex function system on the spherical surface

$$
\begin{equation*}
P_{l m}(\overrightarrow{\boldsymbol{\Omega}})=e^{i m \varphi} \frac{(-\sin (\vartheta))^{m}}{l!2^{l}} \cdot\left(\frac{(l-m)!}{(l+m)!}\right)^{\frac{1}{2}} \frac{d^{l+m}\left(\cos ^{2} \vartheta-1\right)^{l}}{(d \cos \vartheta)^{l+m}}=e^{i m \varphi} \frac{(\sin (\vartheta))^{-m}}{l!2^{l}} \cdot\left(\frac{(l+m)!}{(l-m)!}\right)^{\frac{1}{2}} \frac{d^{l-m}\left(\cos ^{2} \vartheta-1\right)^{l}}{(d \cos \vartheta)^{l-m}} \tag{8.10}
\end{equation*}
$$

with

$$
\begin{equation*}
P_{l,-m}(\overrightarrow{\boldsymbol{\Omega}})=(-)^{m} P_{l m}^{*}(\overrightarrow{\boldsymbol{\Omega}}) \tag{8.11}
\end{equation*}
$$

and

$$
\begin{equation*}
\int_{4 \pi} d \overrightarrow{\boldsymbol{\Omega}} P_{l^{\prime} m^{\prime}}(\overrightarrow{\boldsymbol{\Omega}}) P_{l m}^{*}(\overrightarrow{\boldsymbol{\Omega}})=\delta_{l^{\prime} l} \delta_{m^{\prime} m} \frac{4 \pi}{2 l+1} \tag{8.12}
\end{equation*}
$$

All continuously differentiable functions on the spherical surface $f(\Omega)=f(\theta, \phi)$ can be developed according to

$$
\begin{equation*}
f(\overrightarrow{\boldsymbol{\Omega}})=\sum_{l=0}^{\infty} \sum_{m=-l}^{m=+l} f_{l m} P_{l m}(\overrightarrow{\boldsymbol{\Omega}})=\sum_{l=0}^{\infty} \sum_{m=-l}^{m=+l} f_{l m} P_{l m}^{*}(\overrightarrow{\boldsymbol{\Omega}}) \tag{8.13}
\end{equation*}
$$

the $f_{l m}$ representing the evolution coefficients. The $P_{l m}^{*}(\overrightarrow{\boldsymbol{\Omega}})$ being complex to $P_{l m}(\overrightarrow{\boldsymbol{\Omega}}) f(\overrightarrow{\boldsymbol{\Omega}})$ can be alternatively considered

The spherical harmonics for $l=0,1$ are

$$
\begin{align*}
P_{00} & =P_{00}^{*}=1 \\
P_{1,-1}(\overrightarrow{\boldsymbol{\Omega}}) & =2^{-\frac{1}{2}} e^{-i \varphi} \sin (\vartheta), P_{1,-1}^{*}=2^{-\frac{1}{2}} e^{i \varphi} \sin (\vartheta) \\
P_{1,0}(\overrightarrow{\boldsymbol{\Omega}}) & =P_{1,0}^{*}(\overrightarrow{\boldsymbol{\Omega}})=\cos (\vartheta)=P_{1}(\overrightarrow{\boldsymbol{\Omega}})  \tag{8.14}\\
P_{1,1}(\overrightarrow{\boldsymbol{\Omega}}) & =-2^{-\frac{1}{2}} e^{i \varphi} \sin (\vartheta), P_{1,1}^{*}(\overrightarrow{\boldsymbol{\Omega}})=-2^{-\frac{1}{2}} e^{-i \varphi} \sin (\vartheta) .
\end{align*}
$$

The connection of spherical harmonics and Legendre-polynomials is obtained by

$$
\begin{equation*}
P_{l 0}=P_{l 0}^{*}=P_{l} \tag{8.15}
\end{equation*}
$$

Furthermore the addition theorem

$$
\begin{equation*}
P_{l}(\cos (\vartheta))=\sum_{m=-l}^{m=+l} P_{l m}\left(\overrightarrow{\boldsymbol{\Omega}}^{\prime}\right) P_{l m}^{*}(\overrightarrow{\boldsymbol{\Omega}}), \quad \cos (\vartheta)=\overrightarrow{\boldsymbol{\Omega}}^{\prime} \cdot \overrightarrow{\boldsymbol{\Omega}} \tag{8.16}
\end{equation*}
$$

The $\delta$-function depending on the spherical harmonics may be stated by

$$
\begin{equation*}
\delta\left(\overrightarrow{\boldsymbol{\Omega}}, \overrightarrow{\boldsymbol{\Omega}}^{\prime}\right)=\sum_{l=0}^{\infty} \frac{2 l+1}{4 \pi} \sum_{m=-l}^{m=+l} P_{l m}(\overrightarrow{\boldsymbol{\Omega}}) P_{l m}^{*}\left(\overrightarrow{\boldsymbol{\Omega}}^{\prime}\right)=\sum_{l=0}^{\infty} \frac{2 l+1}{4 \pi} P_{l}\left(\overrightarrow{\boldsymbol{\Omega}} \cdot \overrightarrow{\boldsymbol{\Omega}}^{\prime}\right) \tag{8.17}
\end{equation*}
$$

### 8.3 Turbulence-Functions

Functions of the unit direction vectors $\overrightarrow{\boldsymbol{\Omega}} \perp \overrightarrow{\boldsymbol{\Theta}}$ are represented by a complete orthogonal function system meaning an extension of the spherical harmonics. We call them turbulence functions.

$$
\begin{align*}
& Q_{l m k}(\overrightarrow{\boldsymbol{\Omega}}, \overrightarrow{\boldsymbol{\Theta}})= P_{l m}(\overrightarrow{\boldsymbol{\Omega}}) H_{k}(\overrightarrow{\boldsymbol{\Theta}}) \\
& P_{l m}(\overrightarrow{\boldsymbol{\Omega}}) \quad \text { spherical harmonics } \\
& \int_{2 \pi} H_{k^{\prime}}(\overrightarrow{\boldsymbol{\Theta}}) H_{k}^{*}(\overrightarrow{\boldsymbol{\Theta}}) d \overrightarrow{\boldsymbol{\Theta}}= \begin{cases}2 \pi & \text { for } \mathrm{k}^{\prime}=\mathrm{k} \\
0 & \text { else }\end{cases}  \tag{8.18}\\
& H_{k}(\overrightarrow{\boldsymbol{\Theta}})=e^{i k \theta} \\
& \cos (\vartheta)=\overrightarrow{\boldsymbol{\Omega}}^{\prime} \cdot \overrightarrow{\boldsymbol{\Omega}} . \tag{8.19}
\end{align*}
$$

with

$$
\int_{2 \pi} \int_{4 \pi} Q_{l m k}(\overrightarrow{\boldsymbol{\Omega}}, \overrightarrow{\boldsymbol{\Theta}}) Q_{l m k}^{*}\left(\overrightarrow{\boldsymbol{\Omega}}^{\prime}, \overrightarrow{\boldsymbol{\Theta}}^{\prime}\right) d \overrightarrow{\boldsymbol{\Omega}}^{\prime} d \overrightarrow{\boldsymbol{\Theta}}^{\prime}= \begin{cases}\frac{8 \pi^{2}}{2 l+1} & \text { for } l=l^{\prime} ; m=m^{\prime} ; k=k^{\prime}  \tag{8.20}\\ 0 & \text { else }\end{cases}
$$

Such, suitable distribution functions are described by

$$
\begin{align*}
f_{t_{\epsilon}}(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\Omega}}, \overrightarrow{\boldsymbol{\Theta}}) & =\sum_{l=0}^{+\infty} \sum_{m=-l}^{+l} \sum_{k=-\infty}^{+\infty} f_{l m k}(\overrightarrow{\boldsymbol{x}}, t) Q_{l m k}(\overrightarrow{\boldsymbol{\Omega}}, \overrightarrow{\boldsymbol{\Theta}}) \\
f(\overrightarrow{\boldsymbol{x}}, t, \overrightarrow{\boldsymbol{\Omega}}, \boldsymbol{\Theta}) & =\sum_{l=0}^{+\infty} \sum_{m=-l}^{+l} P_{l m}(\overrightarrow{\boldsymbol{\Omega}}) \sum_{k=-\infty}^{+\infty} f_{l m k}(\overrightarrow{\boldsymbol{x}}, t) H_{k}(\overrightarrow{\boldsymbol{\Theta}}) . \tag{8.21}
\end{align*}
$$

Die $\delta$-function depending on the turbulence functions is expressed

$$
\begin{equation*}
\delta\left(\overrightarrow{\boldsymbol{\Omega}}, \overrightarrow{\boldsymbol{\Omega}}^{\prime} ; \overrightarrow{\boldsymbol{\Theta}}, \overrightarrow{\boldsymbol{\Theta}}^{\prime}\right)=\sum_{l=0}^{\infty} \frac{2 l+1}{4 \pi} \sum_{m=-l}^{m=+l} P_{l m}(\overrightarrow{\boldsymbol{\Omega}}) P_{l m}^{*}\left(\overrightarrow{\boldsymbol{\Omega}}^{\prime}\right) \sum_{k=-\infty}^{+\infty} \frac{1}{2 \pi} H_{k}(\overrightarrow{\boldsymbol{\Theta}}) H_{k}^{*}\left(\overrightarrow{\boldsymbol{\Theta}}^{\prime}\right) \tag{8.22}
\end{equation*}
$$

and such

$$
\begin{equation*}
\delta\left(\overrightarrow{\boldsymbol{\Omega}}, \overrightarrow{\boldsymbol{\Omega}}^{\prime} ; \overrightarrow{\boldsymbol{\Theta}}, \overrightarrow{\boldsymbol{\Theta}}^{\prime}\right)=\frac{1}{8 \pi^{2}} \sum_{l=0}^{\infty}(2 l+1) P_{l}\left(\overrightarrow{\boldsymbol{\Omega}} \cdot \overrightarrow{\boldsymbol{\Omega}}^{\prime}\right) \sum_{k=-\infty}^{+\infty} \exp \left(i k\left(\Theta-\Theta^{\prime}\right)\right) \tag{8.23}
\end{equation*}
$$

## Bibliography

[1] Book
Chandrasekhar S.: Stochastic Problems in Physics and Astronomy Reviews of Modern Physics 15, no. 1 (1943): 1-89
[2] Book
Feynman Richard P.: The Feynman LECTURES ON PHYSICS Volume I Addison-Wesley Publishing Company, Inc 1964 page 3-9
[3] Book
Greiner Walter/Reinhardt Joachim: Theoretische Physik Band 7: Quantenelektrodynamik Verlag Harri Deutsch . Thun und Frankfurt am Main 1984 page 18
[4] Book
Grosche C. Steiner: Handbook of Feynman Path Integrals 1998 Springer page 2
[5] Online Document
Khanh Tuoc,Trinh: A THEORY OF TURBULENCE-K.T.Trinh@massey.ac.nz page 43
[6] Book
Sen R. N.: Causality, Measurement Theory and the Differentiable Structure of Space-Time CAMBRIDGE MONOGRAPHS ON MATHEMATICAL PHYSICS 2010 page 11
[7] Book
Weinberg Alwin M., Wigner Eugine P.: THE PHYSICAl THEORY of NEUTRON CHAIN REACTORS The University of Chicago Press 1958
[8] Online-Document
Wilczek M.: Theory for the single-point velocity statistics of fully developed turbulence- A. Daitche and R. Friedrich February 2011


[^0]:    ${ }^{1}$ in english literature $\operatorname{curl}(\overrightarrow{\mathbf{v}}) \neq \mathbf{0}$ is used but in turbulence the name $\boldsymbol{r o t}$ is more adapted as will be seen

[^1]:    ${ }^{2}$ This relationship is not to be found in literature, although it is obvious and mathematically not very demanding.
    ${ }^{3}$ This is one reason why the known millenium prize question does not lead to a solution of the turbulence problem. However the validity problem of the Navier-Stokes-equations is more fatal.So this is not a question for mathematics at first, but for physics.

[^2]:    ${ }^{1}$ That is why turbulence can not be uniquely identified by experiments of local velocity statistics.
    ${ }^{2}$ This statement contradicts that of Wilczek[8].
    ${ }^{3}$ The temporal and spatial neighborhood of a turning point does not have such singular properties.

[^3]:    ${ }^{1}$ This statement applies to the Fokker-Planck and Langevin equation. See, for example, Chandrasekhar[1]

[^4]:    ${ }^{1}$ Applying the deterministic theory this problem must be treated numerically.

[^5]:    ${ }^{2} \overrightarrow{\boldsymbol{\Omega}}, \overrightarrow{\boldsymbol{\Theta}}$ would make a single direction vector in a 4-dimensional space. The longitudinal fluctuations in the 4-dimensional space should accord to turbulence in the 3 -dimensional space.

[^6]:    ${ }^{1}$ The test functions otherwise used in distribution theory have an immediate physical meaning in this context with the formulation of the transition probability density.

[^7]:    ${ }^{2}$ by vortex is in this paper allways the swirl $\omega=\frac{1}{2} \vec{\nabla} \times \overrightarrow{\boldsymbol{v}}$ meant

[^8]:    ${ }^{3}$ Symbols as $\omega, r, a, v$ etc. always mean amounts of the corresponding vectors.

