

Fast Radio Bursts from Terraformation

A. Yalinewich,¹★ M. Rahman,², A. Obertas^{3,1} and P. C. Breysse¹

¹Canadian Institute for Theoretical Astrophysics, 60 St. George St., Toronto, ON M5S 3H8, Canada

²Dunlap Institute for Astronomy and Astrophysics, University of Toronto, 50 St. George Street, Toronto, ON, M5S 3H4, Canada

³Department of Astronomy and Astrophysics, University of Toronto, 50 St. George Street, Toronto, ON, M5S 3H4, Canada

Accepted XXX. Received YYY; in original form ZZZ

ABSTRACT

Fast radio bursts (FRBs) are, as the name implies, short and intense pulses of radiation at wavelengths of roughly one metre. FRBs have extremely high brightness temperatures, which points to a coherent source of radiation. The energy of a single burst ranges from 10^{36} to 10^{39} erg. At the high end of the energy range, FRBs have enough energy to unbind an earth-sized planet, and even at the low end, there is enough energy to vaporise and unbind the atmosphere and the oceans. We therefore propose that FRBs are signatures of an artificial terraformer, capable of eradicating life on another planet, or even destroy the planet entirely. The necessary energy can be harvested from Wolf-Rayet stars with a Dyson sphere ($\sim 10^{38}$ erg s⁻¹), and the radiation can be readily produced by astrophysical masers. We refer to this mechanism as Volatile Amplification of a Destructive Emission of Radiation (VADER). We use the observational information to constrain the properties of the apparatus. We speculate that the non-repeating FRBs are low-energy pulses used to exterminate life on a single planet, but leaving it otherwise intact, and that the stronger repeating FRB is part of an effort to destroy multiple objects in the same solar system, perhaps as a preventative measure against panspermia. In this picture, the persistent synchrotron source associated with the first repeating FRB arises from the energy harvesting process. Finally we propose that Oumuamua might have resulted from a destruction of a planet in this manner.

Key words: radio continuum: transients – astrobology – masers

1 INTRODUCTION

Fast radio bursts (FRBs) are intense, and short flashes of radiation with a wavelength of around one metre whose origins are among the final frontier of astrophysics. Their large dispersion measure (hundreds to about two thousand pc cm⁻³) implies the source is of extragalactic origin. Their intrinsic duration (after de-dispersion) is of the order of microseconds, and their typical flux is a few Jansky. Their inferred brightness temperature greatly exceeds the Kellerman limit (Tsang & Kirk 2006) and implies a coherent emission mechanism. For a comprehensive discussion of the observed properties of FRBs we refer the reader to the FRB catalogue (Petroff et al. 2016) and references therein. For a list of proposed mechanisms for FRBs we refer the reader to the FRB theory wiki (Platts et al. 2018).

The vast majority of scenarios involve “natural” sources. One example for a model involving an “artificial” source is the light sail (Lingam & Loeb 2017). In this work we consider a more nefarious artificial source for FRBs. We pro-

pose that FRBs are signatures of an alien weapon of mass destruction, capable of vaporising an earth size planet. The required radiant energy at the right frequency range can be produced using astrophysical maser, composed primarily of volatile compounds. We refer to this apparatus as the Volatile Amplification of a Destructive Emission of Radiation (or VADER).

The plan of the paper is as follows. In section 2 we discuss the theoretical constraints on the VADER system from the observations. In section 4 we discuss the results and their implications.

2 THEORETICAL CONSTRAINTS

2.1 Energy Budget

So far, two FRBs have been localised. The first one is the first repeating FRB 121102 (Spitler et al. 2016). The isotropic equivalent energy for each burst is about 10^{39} erg. The other is FRB 171020 (Mahony et al. 2018). In this case the isotropic equivalent energy is considerably lower - about 10^{36} erg.

★ E-mail: almog.yalin@gmail.com

The binding energy of a terrestrial planet of mass M_p and radius R_p is roughly given by

$$U_b \approx 10^{39} \left(\frac{M_p}{M_\oplus} \right)^2 \left(\frac{R_p}{R_\oplus} \right)^{-1} \text{ erg.} \quad (1)$$

Therefore, a repeating burst has enough energy to entirely unbind a terrestrial planet. The energy in the non repeating burst would suffice to unbind the atmosphere and the oceans on the surface of the planet.

The minimum mass of the emitter can be estimated by assuming that each molecule emits a single photon

$$M_e > 0.1 \frac{\mu}{m_p} \frac{E}{10^{39} \text{ erg}} \left(\frac{\nu}{1 \text{ GHz}} \right)^{-1} M_\odot \quad (2)$$

where E is the energy of the burst, μ is the mass of a single molecule, m_p is the proton mass and ν is the frequency of the radiation. In principle, it is possible to increase the efficiency of the emitter by exciting multiple degrees of freedom. This increase in energy is bounded by the number degrees of freedom, and therefore cannot reduce the minimum mass by more than about an order of magnitude.

We note that this method of destroying a planet usually requires less energy than diverting the planet in the habitable zone toward the host star. This is because usually the orbital Keplerian velocity is larger than the escape velocity from the planet. Moreover, the biggest challenge with this approach is to get rid of the planet's orbital angular momentum.

2.2 Maser Emission

Masers are a well known source of coherent radiation in astrophysics (Gray 1999). Astrophysical masers are primarily produced by molecules comprising volatile elements (with the exception of Silicon). The lowest frequency ever recorded for an astrophysical maser is about 700 MHz for a CH maser (Ziurys & Turner 1985), and the highest frequency is about 3.4 THz, from a CO maser (Storey et al. 1981). This is consistent with the non detection of FRBs below 200 MHz (Sokolowski et al. 2018), while most detections are at or above 800 MHz. However, account for lower apparent frequencies, redshift can.

Each of the maser lines is extremely narrow, but if multiple lines are emitted simultaneously, then when observed with coarse enough frequency resolution, the spectrum may seem continuous.

We note that since many of the compounds found in molecular clouds are also present in planets' atmosphere and mantles, then the emitted energy will be readily absorbed rather than reflected from the target.

2.3 Duration

One of the properties of coherent emission is that it can produce short, intense and polarised pulses. As the density of excited molecules increase, the intensity of the radiation increases and the duration decreases. This effect is often referred to a Dicke's superradiance (Dicke 1954). It has been shown that Dicke's superradiance in astrophysical masers can account for the observed duration and energy of FRBs (Houde et al. 2017).

Even if the pulse is shorter than what is observed, the signal will be broadened due to reflection of the radiation from a curved target. The typical light crossing time for an earth sized planet is:

$$t_{lc} \approx 20 \frac{R}{R_\oplus} \text{ ms} \quad (3)$$

If the radius of the maser beam is smaller than the size of the planet, then the duration of the observed pulse will be shorter. If the size of the beam is a factor of 5 smaller than the radius of the planet, then the spread in arrival times of photons to the surface of the planet will be similar to the observed duration of FRBs (a few microseconds).

2.4 Pump

We propose that the persistent synchrotron source associated with the first repeating FRB (Chatterjee et al. 2017) is related to the energy source used for population inversion in the maser. The radio luminosity of the persistent source is of the order of 10^{38} erg/s, which is comparable with the bolometric luminosity of some Wolf Rayet stars (Hainich et al. 2014). We therefore propose that this energy is harvested by a Dyson sphere (Semiz & Oğur 2015; Osmanov 2016).

The most straightforward way to transport the energy from the Wolf Rayet star is to accelerate its stellar wind to relativistic particles in a magnetically collimated beam. It has been previously estimated that the magnetic field is of the order of $B \approx 10$ mG and the Lorentz factor of the electrons is of the order of $\gamma \approx 100$ (Waxman 2017). The bolometric luminosity of each electron is given by $L_1 \approx c\sigma_T B^2 \gamma^2$ (where σ_T is the Thompson cross section) and the number of electrons is roughly given by $\dot{M}d/cm_p$ where \dot{M} is the mass loss rate, d is the distance between the WR star and the molecular cloud and m_p is the mass of the proton. The total synchrotron luminosity of the beam is therefore

$$L_b \approx 6 \cdot 10^{38} \frac{\dot{M}}{10^{-3} M_\odot/\text{y}} \left(\frac{\gamma}{100} \right)^2 \frac{d}{\text{pc}} \text{ erg/s.} \quad (4)$$

Hence we get that the synchrotron emission from the beam is comparable with the observed value. This means that the majority of energy harvested from the WR star is spent in transporting this energy to the molecular cloud. This could explain why the time between consecutive bursts is considerably longer the ideal charging time (i.e. the ratio between the energy of an individual burst and the luminosity of the synchrotron source, about one minute). In other words, the energy needed to destroy a planet may be insignificant next to the power of the source.

One property that sets the first repeater from other FRBs is the exceptionally high rotation measure, roughly 10^5 rad/meter², which requires a high magnetic field. This magnetic field could be the same magnetic field that confines the beam.

A persistent synchrotron source of a similar power has not been detected around another localised, but non repeating, FRB (Mahony et al. 2018). Moreover, rotation measure of other FRBs (e.g. Petroff et al. 2017) are significantly lower than that of the repeater. For this reason we assume that the pump is turned off prior to triggering maser emission.

The pump remains active in the case of the repeater because the maser has to fire multiple times. One reason to do so is in order to destroy not just a single planet, but multiple objects in the same solar system. This may be especially necessary, as the technology level necessary to create the system we describe here would also allow the creation of planet- or moon-sized objects which are in fact artificial space stations, increasing the number of targets in a single system.

3 INTERSTELLAR DEBRIS

Recently, a first interstellar object, dubbed Oumuamua, has been detected passing through our solar system (Jewitt et al. 2017). Oumuamua was detected by the telescope PAN-STARRS, which have been observing for about ten years before the detection. If this detection is typical, and an interstellar object the size of Oumuamua enters our solar system once per decade, then this would require every solar system to eject debris with a total mass of about $4M_{\oplus}$, which is problematic with current models of planet formation (Do et al. 2018).

In this section we explore the prospect that Oumuamua is a part of the debris from a planet destroyed by the mechanism described in the previous section. A planet destroyed in this manner is expected to produce a wide debris field, as is illustrated by the numerical simulations presented in figure 1. If the escape velocity from the planet is greater than the Keplerian orbital velocity, then the debris are guaranteed to leave the solar system. Stronger explosions can also expel debris out of the solar system from closer or less massive planets.

One of the peculiar features of Oumuamua, an object some already consider to be unnatural (Bialy & Loeb 2018), is the large variations in its light curve, which indicates a large aspect ratio (Fraser et al. 2017). Such a large aspect ratio cannot be readily produced in natural environments. Explosions, however, are known to produce irregularly shaped debris (Baker et al. 1981).

In order for a fragment to have reached us, the explosion had to have been relatively close, and therefore recent (in comparison to the age of the Galaxy). Travelling at roughly 20 km/s, an object would take about a million years to get to us if it were travelling at a straight line. Oumuamua came roughly from the direction of the constellation Lyra, which sports a number of relatively close exoplanets, like Kepler 37b (Barclay et al. 2013), Kepler 444 (Campante et al. 2015) and possibly Vega (Harper et al. 1984).

If VADER mechanisms have been active in the Milky Way in the recent past, there exists a troubling possibility that such a system could be aimed at Earth. We do not calculate the probability of this, as the authors prefer not to be told the odds. The reader may get a bad feeling about this, as with current levels of technology, resistance to such a weapon would of course be futile.

4 DISCUSSION

In this paper we discuss the possibility that fast radio bursts are signatures of an artificial device capable of destroying terrestrial planets. We propose that the device is based on

maser emission in multiple spectral lines. Since such a molecular cloud is primarily composed of volatile elements, we refer to this mechanism as a Volatile Amplification of a Destructive Emission of Radiation (VADER). We show that the frequency range of FRBs is compatible with cosmologically redshifted maser lines, the energy of FRBs is comparable to the binding energy of terrestrial planets and that the de-dispersed FRB duration is comparable to the delay time from the reflection of a planar wave from a planet surface.

In this model, non-repeating FRBs are incidents where just a single planet in a particular solar system is destroyed, while the repeating FRBs are cases where multiple objects in the same system are destroyed. We note dynamical instabilities restrict the number of planets in the habitable zone to 5 or below (Obertas et al. 2017), whereas the number of repeated pulses from FRB121102 is close to 100 (Zhang et al. 2018). This means that not only planets, but also moons and asteroids were destroyed. One reason to do so is to prevent panspermia (Melosh 1988; Hornebeck et al. 1994). The persistent synchrotron source and high rotation measure detected for the first repeating FRB are signatures of the pump energy source. We postulate that this energy source is a Wolf-Rayet star, harvested by a Dyson sphere. This energy is transferred to the molecular cloud by a relativistic particle beam.

As FRBs have been observed across a wide area of sky at extragalactic distances, this model implies that VADER mechanisms are active in numerous, widely separated galaxies. From this, we can imply one of three things: either (a) many civilisations across the universe have independently developed this technology, possibly as a result of some kind of interstellar wars, (b) a single civilisation has existed for the multi-megayear timescales necessary to make the long trek between distant stars, or (c) some kind of “hyperdrive” technology exists allowing for many-parsec journeys to be undertaken faster than relativity would allow (Solo, H., private communication).

To achieve higher efficiency, in terms of conversion of radiative energy to heat, the maser beam should be directed to a part of the planet surface that is rich in volatile elements and metals. On earth, such locations would be where there is fertile soil. For this reason, it could very well be that crop circles are target marks for such a weapon. Therefore, it could be that earth has been marked for destruction multiple times, perhaps to facilitate an intergalactic highway. We note, however, that this may not be the solution that you are looking for.

ACKNOWLEDGEMENTS

We would like to thank George Lucas, Douglas Adams, Isaac Asimov, Gene Roddenberry, and Ridley Scott for their inspirational works.

REFERENCES

- Baker W. E., Kulesz J. J., Westine P. S., Cox P. A., Wilbeck J. S., 1981, Technical report, A manual for the prediction of blast and fragment loadings on structures. SOUTHWEST RESEARCH INST SAN ANTONIO TX
- Barclay T., et al., 2013, *Nature*, 494, 452

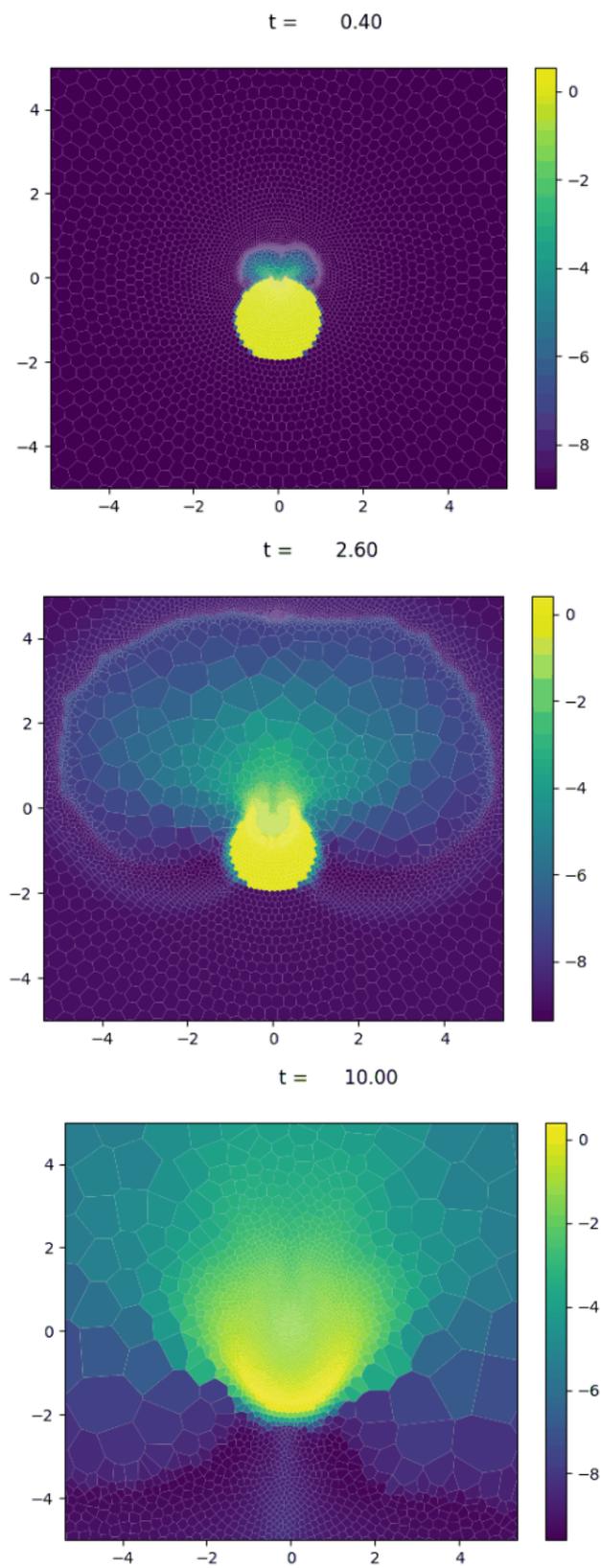


Figure 1. Log density snapshots from an numerical simulation of the passage of a shock wave in a terrestrial planet as a result of a deposition of radiative energy in the top part. The shock wave creates a substantial disturbance in the force of gravity holding the planet together and produces a wide debris field.

- Bialy S., Loeb A., 2018, *The Astrophysical Journal*, 868, L1
- Campante T. L., et al., 2015, *The Astrophysical Journal*, 799, 170
- Chatterjee S., et al., 2017, *Nature*
- Dicke R. H., 1954, *Physical Review*, 93, 99
- Do A., Tucker M. A., Tonry J., 2018, *The Astrophysical Journal*, 855, L10
- Fraser W. C., Pravec P., Fitzsimmons A., Lacerda P., Bannister M. T., Snodgrass C., Smolić I., 2017, *Nature Astronomy*, Volume 2, p. 383-386, 2, 383
- Gray M., 1999, *Philosophical Transactions of the Royal Society A: Mathematical, Physical and Engineering Sciences*, 357, 3277
- Hainich R., et al., 2014, *Astronomy & Astrophysics*, 565, A27
- Harper D. A., Loewenstein R. F., Davidson J. A., 1984, *The Astrophysical Journal*, 285, 808
- Horneck G., Bücker H., Reitz G., 1994, *Advances in Space Research*, 14, 41
- Houde M., Mathews A., Rajabi F., 2017, *Monthly Notices of the Royal Astronomical Society*, Volume 475, Issue 1, p.514-522, 475, 514
- Jewitt D., Luu J., Rajagopal J., Kotulla R., Ridgway S., Liu W., Augusteijn T., 2017, *The Astrophysical Journal*, 850, L36
- Lingam M., Loeb A., 2017, *The Astrophysical Journal Letters*, Volume 837, Issue 2, article id. L23, 5 pp. (2017)., 837
- Mahony E. K., et al., 2018, *ApJL*, 867, 10
- Melosh J., 1988, *Nature*, 332, 687
- Obertas A., Van Laerhoven C., Tamayo D., 2017, *Icarus*, Volume 293, p. 52-58., 293, 52
- Osmanov Z., 2016, *International Journal of Astrobiology*, 15, 127
- Petroff E., et al., 2016, *Publications of the Astronomical Society of Australia*, Volume 33, id.e045 7 pp., 33
- Petroff E., et al., 2017, *MNRAS*, 469, 4465
- Platts E., Weltman A., Walters A., Tendulkar S. P., Gordin J. E. B., Kandhai S., 2018, eprint arXiv:1810.05836
- Semiz A., Oğur S., 2015, arxiv
- Sokolowski M., et al., 2018, *The Astrophysical Journal Letters*, Volume 867, Issue 1, article id. L12, 6 pp. (2018)., 867
- Spitler L. G., et al., 2016, *Nature*, Volume 531, Issue 7593, pp. 202-205 (2016)., 531, 202
- Storey J. W. V., Watson D. M., Townes C. H., Haller E. E., Hansen W. L., 1981, *The Astrophysical Journal*, 247, 136
- Tsang O., Kirk J. G., 2006, *Astronomy and Astrophysics*, Volume 463, Issue 1, February III 2007, pp.145-152, 463, 145
- Waxman E., 2017, *The Astrophysical Journal*, 842, 34
- Zhang Y. G., Gajjar V., Foster G., Siemion A., Cordes J., Law C., Wang Y., 2018,] 10.3847/1538-4357/aadf31
- Ziurys L. M., Turner B. E., 1985, in , *Molecular Astrophysics*. Springer Netherlands, Dordrecht, pp 591–601, doi:10.1007/978-94-009-5432-8_26, http://www.springerlink.com/index/10.1007/978-94-009-5432-8_26

This paper has been typeset from a $\text{T}_{\text{E}}\text{X}/\text{L}^{\text{A}}\text{T}_{\text{E}}\text{X}$ file prepared by the author.