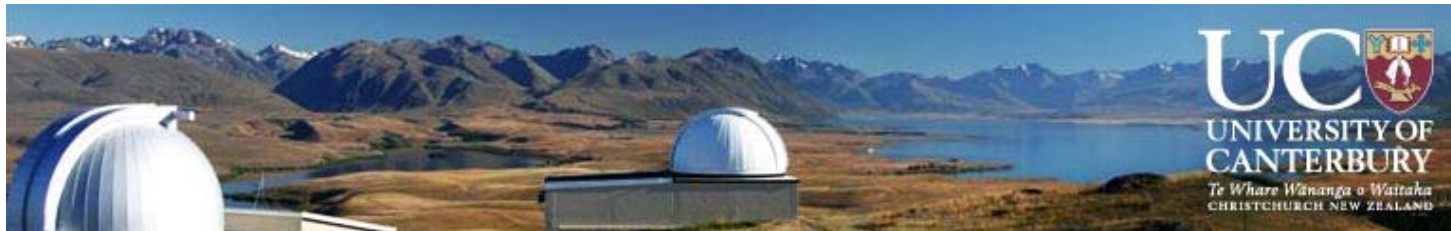


**Applications of Euclidian Snyder
geometry to space time physics
& Deceleration parameter (DE replacement?)
Analysis of linkage between 1st, 2nd inflation?**

**For ACGRG5 & Beyond Standard Model, 2010, Capetown, S.A.
Xian, PIERS 2010, PRC**



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Abstract

A thought experiment: LQG or string theory as an initial space-time template for emergent gravity?

- Applications of deformed Euclidian space to questions about the role of string theory and/or LQG
- To what degree are the fundamental constants of nature preserved between different cosmological cycles?
- To what degree is gravity an emergent field that is partly/largely classical with extreme nonlinearity, or a QM/quantum field theory phenomenon?

Math – Physics representation of core issues of higher dimensional contribution

- **Start off with a basic statement of strength of matter - graviton interaction, assuming KK graviton**

$$\mathfrak{J} = -\frac{K}{2} \cdot \sum_n \int d^4x h_{uv}^n \cdot T^{uv} \sim 1 / M_{PL}^2$$

The stress energy tensor comes from the standard model, and the h term is from using a KK graviton interactions model , up to the n th mode.

Has some similarity with graviton-neutrino interaction issues talked about in this PPT

Does the last slide hold if we make the following modification of a KK tower of gravitons ?

- Modification put in, as seen in later to mimic DE
- Suggestion to look at, here, is to consider what if
- $m_n(\textit{Graviton}) = \frac{n}{L} + 10^{-65}$ grams?

Issues to raise

- 1. Is there a link between a 1st inflation, ending in 10^{-35} seconds after big bang, and 2nd inflation commencing before 1 billion years ago ?
- 2. Commonality between the two ?
- 3. Do gravitons, with tiny mass play a role in the 1st and 2nd inflationary episodes ?

Look at A. Yurov; arXiv : hep-th/028129 v1, 19 Aug, 2002

- Claim: Exist one emergent complex scalar field . Accounts for both 1st and 2nd inflation

-

Potential in both cases chaotic inflation of the type

$$V = \vec{m}^2 \Phi^* \Phi$$

Mass m is for inflaton, and 2nd \vec{m} expression has links to 5th dim length L

$$\vec{m} \approx \sqrt{\frac{3}{8}} \cdot \left[\sqrt{\frac{3H^2}{4\pi G}} \Big|_{\text{time} \sim 10^{-35} \text{ sec}} + \sqrt{\frac{3H^2}{4\pi G}} \Big|_{\text{time} \sim 10^{-44} \text{ sec}} \right]$$

$$|\vec{m}| \leq \left[\frac{l^2}{4} \right]$$

Important simplification used

- From beginning of inflation, assume V potential energy is much smaller than H contribution

$$\frac{3H^2}{4\pi G} \gg V(t) \Big|_{\text{time} \sim 10^{-44} \text{ sec}}$$

The 5th dimensional length L is for
a brane theory “arc-length”

- From Roy Maarten’s Brane theory work

$$dS^2 \Big|_{5\text{-dim}} = \frac{l^2}{z^2} \cdot \left[\eta_{uv} dx^\mu dx^\nu + dz^2 \right]$$

Find equivalence between

- 1st , 2nd equations for Friedman relations

$$H^2 = \frac{1}{6} \cdot \left[\dot{\phi}^2 + \tilde{m}^2 \phi^2 + \frac{M^2}{\phi^2} \right] \leftrightarrow \left(\frac{\tilde{\kappa}^2}{3} \left[\rho + \frac{\rho^2}{2\lambda} \right] \right) + \frac{m}{a^4}$$

$$\dot{H}^2 \cong \left[-2 \frac{m}{a^4} \right] \leftrightarrow \dot{H} = V - 3H^2$$

Beckwith asserts there may be some reason to expect linkage between 1st and 2nd inflation

- Look at if the scalar field arises in 2nd inflation due to

$$dS^2 \Big|_{5\text{-dim}} = \frac{l^2}{z^2} \cdot \left[\eta_{uv} dx^\mu dx^\nu + dz^2 \right]$$

$$|\vec{m}| \leq \left[\frac{l^2}{4} \right]$$

$$\phi_{0,-} = \sqrt{2/3} \cdot \vec{m} \cdot \left[t_{1st-EXIT} \sim 10^{-35} \text{ sec} \right]$$

$$\phi_+ = \left[\phi_{0,+}^3 - \sqrt{3/2} \cdot \frac{3M^2 t}{\vec{m}} \right]^{1/3}$$

Snyder formulation of HUP

1st Basic relation

$$[q, p] = i \cdot \sqrt{1 - \alpha \cdot p^2} \Leftrightarrow \Delta q \Delta p \geq \frac{1}{2} \cdot \left| \left\langle \sqrt{1 - \alpha \cdot p^2} \right\rangle \right|$$

2nd Basic relation

$$\Delta q \geq \left[(1/\Delta p) + l_s^2 \cdot \Delta p \right] \equiv (1/\Delta p) - \alpha \cdot \Delta p$$

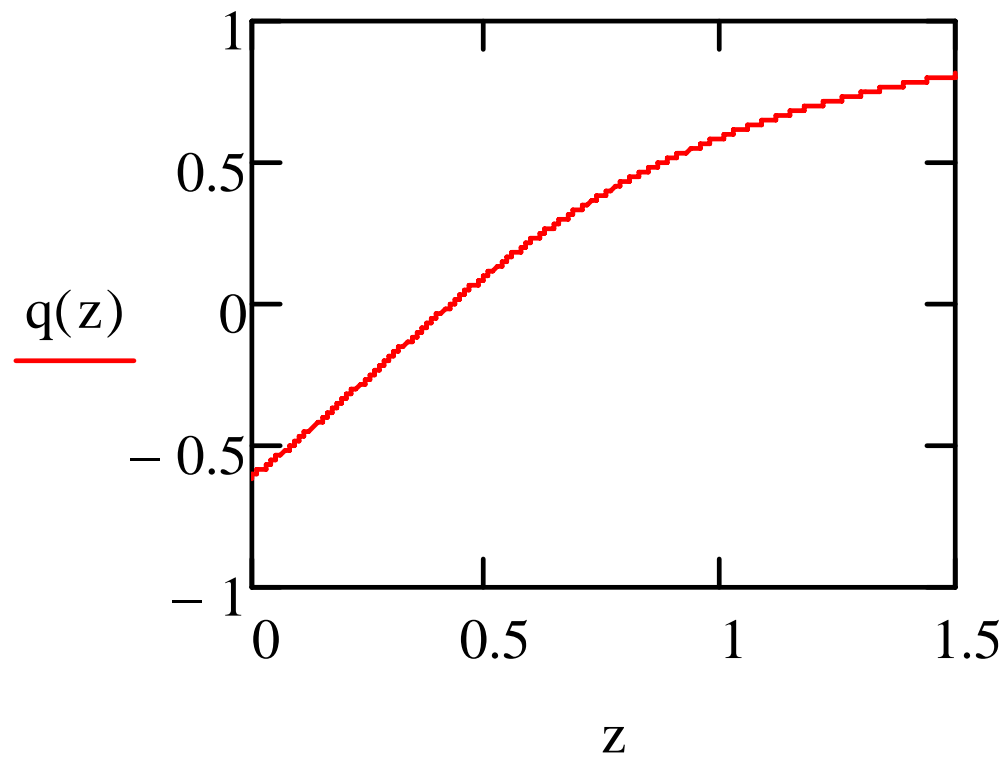
3rd Basic relation

LQG has $\alpha > 0$

Braneworld $\alpha < 0$

Jerk calculation leads to

- If we define the jerk $q = -\frac{\ddot{a}a}{\dot{a}^2}$



Assuming a brane world

Z (red shift value). Change in sign for Z $\sim .42$ is almost one billion years ago, corresponding to reacceleration of the universe, i.e

Basic results of [Alves](#), et al. (2009), using their parameter values, with an additional term of "dark flow" added, corresponding to one KK additional dimensions.

For brane world, the following modification of Roy Maarten's

- KK tower assumed to have a small non-zero mass added, i.e. no zero order value for the graviton

4 – D graviton $\sim 10^{-65}$ grams

$$m_n(\text{Graviton}) = \frac{n}{L} + 10^{-65}$$

For brane world, use these
evolution equations

Friedman equation, subsequently
modified

$$\dot{a}^2 = \left[\left(\frac{\rho}{3M_4^2} + \frac{\Lambda_4}{3} + \frac{\rho^2}{36M_{Planck}^2} \right) a^2 - \kappa + \frac{C}{a^2} \right]$$

Density equation, with non-
zero graviton mass

$$\rho \equiv \rho_0 \cdot \left(\frac{a_0}{a} \right)^3 - \left[\frac{m_g c^6}{8\pi G \hbar^2} \right] \cdot \left(\frac{a^4}{14} + \frac{2a^2}{5} - \frac{1}{2} \right)$$

For LQG, use these evolution equations

Friedman equations, assuming 'constant' momentum

$$\left(\frac{\dot{a}}{a}\right)^2 = \frac{\kappa}{3} \cdot \rho \quad \left(\frac{\dot{a}}{a}\right)^2 \equiv \frac{\kappa}{6} \cdot \frac{p_\phi^2}{a^6} \quad \left(\frac{\ddot{a}}{a}\right) = -\frac{2 \cdot \kappa}{3} \cdot \rho$$

Density equation

$$\rho \equiv \rho_0 \cdot \left(\frac{a_0}{a}\right)^3 - \left[\frac{m_g c^6}{8\pi G \hbar^2} \right] \cdot \left(\frac{a^4}{14} + \frac{2a^2}{5} - \frac{1}{2} \right)$$

Can neutrinos interact with Gravitons? Part 1

Bashinsky states that the density of gravitons interacting with neutrinos causes an alteration of overall GR density via

$$\left[1 - 5 \cdot \left(\rho_{neutrino} / \rho \right) + \mathcal{G} \left(\left[\rho_{neutrino} / \rho \right]^2 \right) \right]$$

Can neutrinos interact with Gravitons? Part 2

- **George Fuller and Chad Kishimoto's PRL stretched neutrino hypothesis: a neutrino could be stretched 'across the universe' leading to (if there is an interaction with gravitons):**

A few select gravitons, coupled to almost infinite wavelength stretched neutrinos would lead to at least the following stretched graviton wave

$$\lambda_{graviton} \equiv \frac{\hbar}{m_{graviton} \cdot c} < 10^4 \text{ meters}$$

Semiclassical interpretation of giant graviton waves?

Brought up as a way to interpret the existence of a small graviton mass, which appears to violate the QM correspondence principle (shown later)

Main motivation: a field theory limit demo that shows problems with **massive graviton field theories, and the limit**

$$m_{\text{graviton}} \rightarrow 0$$

How to measure a graviton/ GW ?

- Look at the normalized gravitational wave density function

$$\Omega_{gw} \equiv \frac{\rho_{gw}}{\rho_c} \equiv \int_{f=0}^{f=\infty} d(\log f) \cdot \Omega_{gw}(f) \Rightarrow h_0^2 \Omega_{gw}(f) \cong 3.6 \cdot \left[\frac{n_f}{10^{37}} \right] \cdot \left(\frac{f}{1\text{kHz}} \right)^4$$

- Note that n_f depends upon frequency and is stated to be part of the unit phase space

Infinite Quantum statistics. From the work presented in the Paris observatory, July 2009

Start with

$$Z_N \sim \left(\frac{1}{N!}\right) \cdot \left(\frac{V}{\lambda^3}\right)^N \quad S \approx N \cdot (\log[V/N\lambda^3] + 5/2)$$

$$S \approx N \cdot (\log[V/\lambda^3] + 5/2) \quad V \approx R_H^3 \approx \lambda^3$$

V stands for volume of nucleation regime space.
“particles” nucleate from ‘vacuum’ in QM

For DM. V for nucleation is HUGE. Graviton space
V for nucleation is tiny , well inside inflation/
Therefore, the log factor drops OUT of entropy S
if V chosen properly. For small V, then

$$\Delta S \approx \Delta N_{gravitons}$$

Some considerations about the partition function

Glinka (2007): if we identify $\Omega = \frac{1}{2|u|^2 - 1}$

- as a partition function (with u part of a Bogoliubov transformation) due to a graviton-quintessence gas, to get information theory-based entropy $S \equiv \ln \Omega$

1. Derivation by Glinka explicitly uses the Wheeler De Witt equation
2. 2. Is there in any sense a linkage of Wheeler De Witt equation with String theory results ?

PROBLEM TO CONSIDER:

Ng's result quantum counting algorithm is a **STRING theory** result. Glinka is **Wheeler De Witt equation. Equivalent ?**

Questions to raise.

Can we make a linkage between Glinka's quantum gas argument, and a small space version/ application of Ng's Quantum infinite statistics ?

In addition, if the quantum graviton gas is correct, can we model emergent structure of gravity via linkage between Ng particle count, and Q.G.G argument?

**LQG , while using WdW up to a point, does not admit higher dimensions above 4 dimensions .
String-Brane theory does**

- Why is this relevant to a discussion of the LQG vs Brane theory discussion ?

Breakdown of field theory with respect to massive gravitons in limit

$$m_{\text{graviton}} \rightarrow 0$$

The massless equation of the graviton evolution equation takes the form

$$\partial_{\mu} \partial^{\mu} h_{\mu\nu} = \sqrt{32\pi G} \cdot \left(T_{\mu\nu} - \frac{1}{2} \eta_{\mu\nu} T^{\mu}_{\mu} \right)$$

Consider what happens with a graviton mass

$$m_{\text{graviton}} \neq 0$$

From Maggiore (2008):

$$\left(\partial_{\mu}\partial^{\mu} - m_{\text{graviton}}\right) \cdot h_{\mu\nu} = \left[\sqrt{32\pi G} + \delta^{+}\right] \cdot \left(T_{\mu\nu} - \frac{1}{3}\eta_{\mu\nu}T^{\mu}_{\mu} + \frac{\partial_{\mu}\partial_{\nu}T^{\mu}_{\mu}}{3m_{\text{graviton}}}\right)$$

The mismatch between these two equations when

$$m_{\text{graviton}} \rightarrow 0$$

Is largely due to, even if graviton mass goes to zero

$$m_{\text{graviton}} h_{\mu}^{\mu} \neq 0$$

$$m_{\text{graviton}} \cdot h_{\mu}^{\mu} = -\left[\sqrt{32\pi G} + \delta^+ \right] \cdot T_{\mu}^{\mu}$$

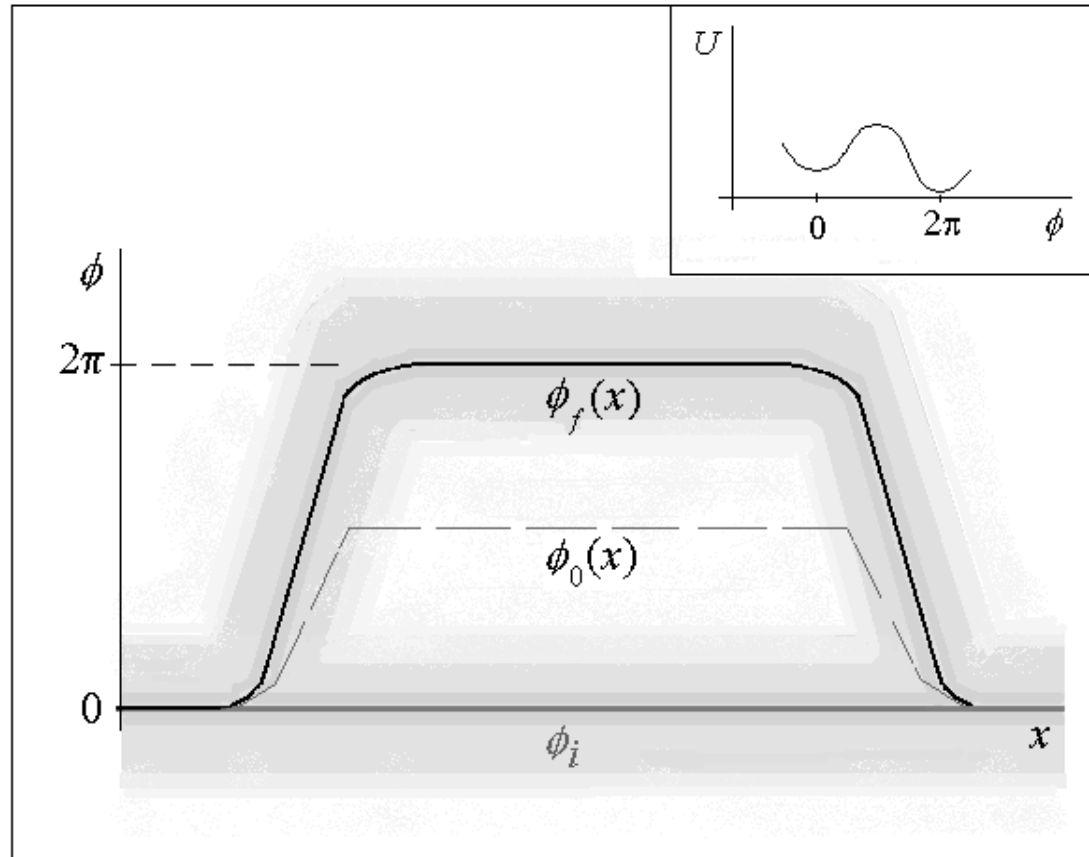
Try semiclassical model of graviton, as kink-anti kink pair

- How does this fit in with t'Hooft's deterministic QM?
- From a 1+ dimensional kink-antikink

$$\Psi_{i,f} [\phi(\mathbf{x})]_{\phi=\phi_{ci,cf}} = c_{i,f} \cdot \exp \left\{ - \int d\mathbf{x} \alpha \left[\phi_{Ci,f}(\mathbf{x}) - \phi_0(\mathbf{x}) \right]^2 \right\},$$

From density wave physics, 1+ dimensions

Kink-antikinks lead to a vacuum wave function. The LHS is a kink; the RHS is an antikink.



The wave functional should have t'Hooft equivalence class structure added, in 4 to 5 dimensions

- T'Hooft used in 2006 an equivalence class argument as an embedding space for simple harmonic oscillators, as given in his Figure 2, on page 8 of his 2006 article.
- “Beneath Quantum Mechanics, there may be a deterministic theory with (local) information loss. This may lead to a sufficiently complex vacuum state.” - t'Hooft
- The author submits, that a kink-anti kink formulation of the graviton, when sufficiently refined, may indeed create such a vacuum state, as a generalization of Fig 2.

One to four-five dimensions in instanton, anti-instanton construction

For one dimension, the semiclassical treatment has (CDW) a kink given by Beckwith(2001) as

$$\phi_+(z, \tau) = 4 \cdot \arctan \left(\exp \left\{ \frac{z + \beta \cdot \tau}{\sqrt{1 - \beta^2}} \right\} \right)$$

$$\frac{\partial^2 \phi(z, \tau)}{\partial \tau^2} - \frac{\partial^2 \phi(z, \tau)}{\partial z^2} + \sin \phi(z, \tau) = 0$$

**In five dimensions, M. Giovannini
(2006) has constructed**

For a five dimensional line element,

$$dS^2 = a(w) \cdot \left[\eta_{uv} dx^u dx^v - dw^2 \right]$$

$$\phi = \tilde{v} + \arctan\left((bw)^v\right)$$

Supposition to get about the singularity in 4 dimensions, in early universe models

- Dropping in of 'information' to form an instanton-anti-instanton pair, and avoiding the cosmological singularity via the 5th dimension?
- This lead to the author presenting in Chongqing, 11/15/2009 the region about the GR singularity is definable via a ring of space-time about the origin, but not overlapping it, with a time dimension defined

$$\Delta t \equiv 10^{\beta} \cdot t_{Planck}$$

The small mass of the graviton
would be for energy in

$$\Delta E \Delta t \geq \hbar$$

- Having said this, the author is fully aware of the String theory HUP variant

$$\Delta x \geq \frac{\hbar}{\Delta p} + \frac{l_s^2}{\hbar} \Delta p$$

- The idea would be to possibly obtain a way to look at counting for GW detectors

$$h_0^2 \Omega_{gw}(f) \cong \frac{3.6}{2} \cdot \left[\frac{n_f [\textit{graviton}] + n_f [\textit{neutrino}]}{10^{37}} \right] \cdot \left(\frac{\langle f \rangle}{1\text{kHz}} \right)^4$$

The following is claimed:

If n (graviton) is obtained, then higher dimensional geometry may be relevant to transmitting information via gravitons from prior to present universes

- How much information can be carried by an individual graviton?
- Assume $\Delta S \approx \Delta N_{\text{gravitons}}$
- Use Seth Lloyd's

$$I = S_{\text{total}} / k_B \ln 2 = [\# \text{operations}]^{3/4} = [\rho \cdot c^5 \cdot t^4 / \hbar]^{3/4}$$

10^{20} relic gravitons yields almost
 10^{27} operations!

This value implies that per graviton, as nucleated at least 4 dimensions, there is at least **one unit** of information associated with the graviton (assuming there is at least **some relationship** between an operation and information)

$\Delta S \approx \Delta N_{\text{gravitons}} \approx 10^{20} \Leftrightarrow 10^{20}$ or higher
amounts of prior universe information transmitted
to our cosmos?

Cosmological parameters and information from prior to present cosmos ?

- The fine structure constant would probably be a place to start, in terms of information

$$\tilde{\alpha} \equiv e^2 / \hbar \cdot c \equiv \frac{e^2}{d} \times \frac{\lambda}{hc}$$

What the author thinks, is that higher dimensional models of gravity need to be developed, investigated, which may allow for such a counting algorithm.

Resolutions of questions about cosmological constants ?

- **1st Conclusion**, one needs a reliable information packing algorithm! I.e. for a wave length , as input into the fine structure constant, we need spatial / information limits defined for geometry
- $\Delta S \approx \Delta N_{gravitons} \approx 10^{20}$ is only a beginning
- **2nd Conclusion**, assumed GW detector sensitivity limits need a comprehensive look over, re do

Important consideration for review - Is there a linkage between neutrinos and gravitons? What about gravitons and E & M (photons)?

- From a 3 page article submitted to the 12th Marcel Grossman conference proceedings
- **STRETCHED NEUTRINOS, AND THE SUPPOSED LINK TO Gravity / Gravitational waves data SETS**
- ANDREW WALCOTT BECKWITH

- The issue of whether or not a correlation exists between neutrino physics and gravitational wave data sets/gravitons is raised anew. Particular emphasis is placed on analysis of the Fuller and Kishimoto scenario, suggesting that the wave function of a relic neutrino may span up to billions of light years across galaxies because of its low energy and particles traveling at different speeds.
- **QUESTION ASKED :**
- If there is an initial close relationship between gravitational waves /gravitons and relic neutrinos in early-universe nucleation, is there a corresponding "stretch-out" of gravitons? If so, what would this imply for improved graviton/gravity wave detectors?

IF so, then what can we say about the following energy density ?

- We first start off with

$$\Omega_{gw} \equiv \frac{\rho_{gw}}{\rho_c} \equiv \int_{f=0}^{f=\infty} d(\log f) \cdot \Omega_{gw}(f)$$

$$\Rightarrow h_0^2 \Omega_{gw}(f) \cong 3.6 \cdot \left[\frac{n_f}{10^{37}} \right] \cdot \left(\frac{f}{1\text{kHz}} \right)^4$$

If neutrino-graviton coupling is possible,
what also about photons coupled with
gravitons? etc ?

- How reasonable then are the following ?

$$n_f \propto n_f [\textit{graviton}] + n_f [\textit{neutrinos}]$$

$$n_f \propto n_f [\textit{graviton}] + n_f [\textit{neutrinos}] + n_f [\textit{photons} - E \ \& \ M]$$

Final inquiry, making sense of the supposed
“radius of the Universe” calculation

- Matt Roos, has put in a foundational way of testing, via experiment, how to calculate a supposed ‘radius of the universe’

$$r_U \equiv \frac{1}{H \cdot \sqrt{|\Omega - 1|}}$$

Tweaking parameters of H , and

$\Omega \equiv \rho(t) / \rho_{critical}$ ***from our inquiry***

- **The choice of H , and of density , as in the**
- **equation below will allow the dynamics of**
- **how the universe expands mesh with a fuller**
- **understanding of structure formation.**

$$\Omega \equiv \rho(t) / \rho_{critical}$$

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- **(Please look at most recent version ! –**
- **for eventual PRD evaluation once cleaned up)**
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Finally a good book summary with up to date summaries

- From “ Series in High Energy physics, Cosmology and Gravitation” - Taylor and Francis (publishers)
- - Particle and Astroparticle Physics (2008)
- By Utpal Sarkar