# Energy Stored in the Gravitational Field 

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#### Abstract

We evaluate the energy of the gravitational field.


Key Words: Newtonian Gravity.

## 1 Introduction

The attractive force between two charges is given by:

$$
\begin{equation*}
F=\frac{1}{4 \pi \epsilon} \frac{q_{1} q_{2}}{d^{2}} \tag{1}
\end{equation*}
$$

By definition the electric field $e$ is given by:

$$
\begin{equation*}
e=\frac{F}{q} \tag{2}
\end{equation*}
$$

where $F$ is the force experienced by a probe charge $q$. $F$ and $e$ are vectors. The energy density stored in the electric field is given by:

$$
\begin{equation*}
\hat{E}=\epsilon|e|^{2} \tag{3}
\end{equation*}
$$

We turn now our attention to the Newtonian gravitational field. The attractive force between two masses is given by:

$$
\begin{equation*}
F=G \frac{m_{1} m_{2}}{d^{2}} \tag{4}
\end{equation*}
$$

By definition the gravitational field $g$ is given by:

$$
\begin{equation*}
g=\frac{F}{m} \tag{5}
\end{equation*}
$$

where $F$ is the force experienced by a probe mass $m$. Once again $F$ and $g$ are vectors. The field $g$ has units of acceleration and it is in fact the acceleration of the probe mass $m$ if free to move.

By analogy with the electric field we feel safe to say that the energy density stored in the gravitational field is given by:

$$
\begin{equation*}
\hat{E}=\frac{1}{2}\left(\frac{1}{4 \pi G}\right)|g|^{2} \tag{6}
\end{equation*}
$$

[^0]We should be satisfied by the above equation at this point! However, in a boring Saturday afternoon and just to kill some time, we want to prove the above equation by a method of brute force.

To do that, we will assume that the energy density of the gravitational field is given by:

$$
\begin{equation*}
\hat{E}=\frac{1}{2} \Omega|g|^{2} \tag{7}
\end{equation*}
$$

In the next paragraph we will evaluate $\Omega$.

## 2 Evaluation of $\Omega$

Given Fig. 1, the potential of $U$ the gravitational field of two point masses $m_{1}$ and $m_{2}$ is given by:

$$
\begin{equation*}
U=-G\left(\frac{m_{2}}{r_{1}}+\frac{m_{1}}{r_{2}}\right) \tag{8}
\end{equation*}
$$



Figure 1: Definition of $r_{1}$ and $r_{2}$
with $d=2 \delta$ and:

$$
\begin{align*}
& r_{1}=\sqrt{(x-\delta)^{2}+y^{2}}  \tag{9}\\
& r_{2}=\sqrt{(x+\delta)^{2}+y^{2}} \tag{10}
\end{align*}
$$

we have:

$$
\begin{equation*}
U=-G\left(\frac{m_{2}}{\sqrt{(x+\delta)^{2}+y^{2}}}+\frac{m_{1}}{\sqrt{(x-\delta)^{2}+y^{2}}}\right) \tag{11}
\end{equation*}
$$

The gravitational field is given by $g=-\nabla U$. The components of $g$ are therefore:

$$
\begin{equation*}
g_{x}=-\frac{\partial U}{\partial x}=-G\left(\frac{m_{2}(x+\delta)}{\left((x+\delta)^{2}+y^{2}\right)^{\frac{3}{2}}}+\frac{m_{1}(x-\delta)}{\left((x-\delta)^{2}+y^{2}\right)^{\frac{3}{2}}}\right) \tag{12}
\end{equation*}
$$

and:

$$
\begin{equation*}
g_{y}=-\frac{\partial U}{\partial y}=-G\left(\frac{m_{2} y}{\left(y^{2}+(x+\delta)^{2}\right)^{\frac{3}{2}}}+\frac{m_{1} y}{\left(y^{2}+(x-\delta)^{2}\right)^{\frac{3}{2}}}\right) \tag{13}
\end{equation*}
$$

and we have:

$$
\begin{align*}
|g|^{2}=G^{2} & {\left[\left(\frac{m_{2} y}{\left(y^{2}+(x+\delta)^{2}\right)^{\frac{3}{2}}}+\frac{m_{1} y}{\left(y^{2}+(x-\delta)^{2}\right)^{\frac{3}{2}}}\right)^{2}\right.} \\
& \left.+\left(\frac{m_{2}(x+\delta)}{\left(y^{2}+(x+\delta)^{2}\right)^{\frac{3}{2}}}+\frac{m_{1}(x-\delta)}{\left(y^{2}+(x-\delta)^{2}\right)^{\frac{3}{2}}}\right)^{2}\right] \tag{14}
\end{align*}
$$

Let $\Gamma$ to be the semi-plane of the $(x, y)$ plane with $y>0$. We have that:

$$
\begin{gather*}
E=\frac{1}{2} \Omega \int_{V}|g|^{2} d V=\frac{1}{2} \Omega \int_{V}|g(x, y)|^{2} d x d y y d \theta \\
=\frac{G^{2}}{2} \Omega \int_{0}^{2 \pi} d \theta \int_{\Gamma}\left[\frac{y \mid g\left(x,\left.y\right|^{2}\right.}{G^{2}}\right] d x d y=\pi \Omega G^{2} \int_{\Gamma} \Lambda(\delta) d x d y \tag{15}
\end{gather*}
$$

where:

$$
\begin{equation*}
\Lambda(x, y, \delta)=\frac{y \mid g\left(x,\left.y\right|^{2}\right.}{G^{2}} d x d y \tag{16}
\end{equation*}
$$

With some manipulation we have:

$$
\begin{align*}
\Lambda=\frac{m_{1}^{2} y}{\left(y^{2}+(x-\delta)^{2}\right)^{2}}+ & \frac{2 m_{1} m_{2} y\left(y^{2}+x^{2}-\delta^{2}\right)}{\left(y^{2}+(x-\delta)^{2}\right)^{\frac{3}{2}}\left(y^{2}+(x+\delta)^{2}\right)^{\frac{3}{2}}}+\frac{m_{2}^{2} y}{\left(y^{2}+(x+\delta)^{2}\right)^{2}} \\
& =m_{1}^{2} \lambda_{1}+2 m_{1} m_{2} \lambda_{12}+m_{2}^{2} \lambda_{2} \tag{17}
\end{align*}
$$

with obvious meaning of the $\lambda$ symbols.
This result was expected because the energy given by the field $\left|u_{1}+u_{2}\right|$ generated by the two masses has two components depending by the two masses and a cross-component. The two components depending by the masses should not depend on $d=2 \delta$. This is obvious by the fact that with a simple change of coordinates which has no effect on $d x$ we can make the parameter $\delta$ to disappear. If we try to evaluate the integrals relevant to $\lambda_{1}$ and $\lambda_{2}$ they diverge:

$$
\begin{align*}
& \int_{\Gamma} \lambda_{1,2} d x d y=\int_{\Gamma} \frac{m_{1}^{2} y}{\left(y^{2}+(x \pm \delta)^{2}\right)^{2}} d x d y \\
&=\int_{-\infty}^{\infty} \frac{1}{2 x^{2}} d x=\infty \tag{18}
\end{align*}
$$

This is also expected since the energy associated with a point mass is infinite. However they do not depend on $d=2 \delta$ and therefore since we are interested by the derivative of the energy with respect to $d$, as it will be clear later, we
can ignore them for the purpose of our calculations. We are more interested in the mixed term. To evaluate the integral of $\lambda_{1,2}$ it helps to assume that $x>0$. This is not a problem since we know from the problem at hand that the energy is an even function with respect to $x$ and therefore we can evaluate it in one quadrant and double the result. We have:

$$
\begin{gather*}
\int_{\Gamma} \lambda_{12} d x d y=2 \int_{0}^{\infty} d x \int_{0}^{\infty} \lambda_{12} d x d y \\
2 \int_{0}^{\infty} d x \int_{0}^{\infty} \frac{y\left(y^{2}+x^{2}-\delta^{2}\right)}{\left(y^{2}+(x-\delta)^{2}\right)^{\frac{3}{2}}\left(y^{2}+(x+\delta)^{2}\right)^{\frac{3}{2}}} d y \\
=2 \int_{0}^{\infty} \frac{|x-\delta|+x-\delta}{4 x^{2}(x-\delta)} d x \\
=2 \int_{0}^{\delta} 0 d x+2 \int_{\delta}^{\infty} \frac{1}{2 x^{2}} d x=\frac{1}{\delta}=\frac{2}{d} \tag{19}
\end{gather*}
$$

If we do not take into account the terms that go to infinite we have therefore:

$$
\begin{equation*}
E=\pi \Omega G^{2} \int_{\Gamma} \Lambda(\delta) d x d y=\pi \Omega G^{2} 2 m_{1} m_{2} \int_{\Gamma} \lambda_{12} d x d y \tag{20}
\end{equation*}
$$

which is:

$$
\begin{equation*}
E=\frac{4 \pi \Omega G^{2} m_{1} m_{2}}{d} \tag{21}
\end{equation*}
$$

Moreover:

$$
\begin{equation*}
F=-\frac{\partial E}{\partial d}=\frac{4 \pi \Omega G^{2} m_{1} m_{2}}{d^{2}} \tag{22}
\end{equation*}
$$

By equating the above with the attractive force between two masses:

$$
\begin{equation*}
F=G \frac{m_{1} m_{2}}{d^{2}}=\frac{4 \pi \Omega G^{2} m_{1} m_{2}}{d^{2}} \tag{23}
\end{equation*}
$$

we find:

$$
\begin{equation*}
\Omega=\frac{1}{4 \pi G} \tag{24}
\end{equation*}
$$

As expected. Note that we get a $8 \pi G$ factor that is the same present in Einstein field equations and that relate gravitational fields and sources.


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