# Neutrino Oscillations Hint at a New Fundamental Interaction

Mario Everaldo de Souza

Departamento de Física, Universidade Federal de Sergipe, São Cristóvão, Sergipe, Brazil

#### Abstract

It is suggested that neutrino oscillations do not exist and, in the light of the latest results of July 2016 of the Daya Bay Collaboration and the recent discovery of a new fundamental interaction, it is proposed that the neutrino production at the Sun's core in the pp and CNO cycles are much larger than what is predicted by the Standard Solar Model, exactly because of this new interaction. Therefore, the claimed neutrino deficits on Earth are completely misleading.

# 1. Introduction

Electron neutrinos are copiously produced in the interiors of stars. In stars like the sun they are mainly produced in the pp cycle which dominates the fusion process in cool stars

$$\begin{bmatrix} pp \rightarrow de^+ v_e \\ ppe^- \rightarrow dv_e \\ {}^7Bee^- \rightarrow {}^7Liv_e \\ {}^8B \rightarrow {}^7Be^*e^+ v_e \end{bmatrix}$$

This process generates a certain electron neutrino flux  $\Phi_E$  which would be expected to be detected at the Earth. The Super-Kamiokande data [1,2] show that only about 45% of the predicted flux for <sup>8</sup>B electron neutrinos is actually detected. Other experiments with atmospheric electron neutrinos have also reported electron neutrino fluxes with deficits compared to the predicted fluxes by the standard solar model [3-8]. There are also reports of electron neutrino deficits from reactors [9,10]. The neutrino oscillation proposal explains the above problem by supposing that the electron neutrinos are changed into the other neutrino flavors, muon and tau neutrinos, especially to tau neutrinos, through the Mikheyev–Smirnov–Wolfenstein mechanism [11-13].

## 2. A General Critique for the neutrino Oscillation Proposal

The neutrino oscillation proposal has many flaws. I present below only 3 very relevant ones.

### 2.1 Lack of detection of tau neutrino fluxes

The enhancement in tau neutrino fluxes due to neutrino oscillations has not been measured. Actually only the DONUT collaboration [14,15] from Fermilab detected tau neutrinos from the annihilation of  $e^-e^+$  pairs and the OPERA collaboration [16] found one tau neutrino in its quest for neutrino oscillations. **The lack of**  $v_\tau$  **fluxes invalidates the claim for the neutrino oscillations**  $v_e \rightarrow v_\tau$  **and**  $v_\mu \rightarrow v_\tau$ .

#### **2.2 Lepton Number Violation Decays**

According to the neutrino oscillation proposal the three neutrino flavors would be resonant states of the same particle. If this were true we would observe, for example, the weak decays

$$n \rightarrow p + e^{-} + \overline{v}_{\tau}$$

$$n \rightarrow p + e^{-} + v_{\mu}$$

$$\mu^{-} \rightarrow e^{-} + \overline{v}_{\tau} + v_{\tau}$$

$$\mu^{-} \rightarrow e^{-} + \overline{v}_{\mu} + v_{\tau}$$

but these decays have not been observed at all.

#### 2.3 Why the strange preference for the tau neutrino?

If the disappearance of electron solar neutrinos were caused by oscillations we would detect a very high flux of muon neutrinos, but what has been reported is exactly the opposite, that is, there is also a muon neutrino deficit [17]. Why would both neutrinos  $v_e$  and  $v_{\mu}$  prefer to change to  $v_{\tau}$  which, according to the neutrino oscillation proposal, would be the most massive neutrino? Why not the opposite, that is, the oscillation of  $v_{\mu}$  and  $v_{\tau}$  to  $v_e$ ? In such a case we observe higher fluxes of  $v_e$ .

### 3. Report of a New fundamental Interaction

On the other hand there have been theoretical proposals for a fifth force of nature [18] and a recent report on the observation of a new fundamental interaction of matter [19,20] mediated by a light boson. The data of reference 19 was carefully analyzed in detail by Feng et al. [20]. They reinforce the idea of a new fundamental force of nature mediated by means of a light vector boson X with a mass of about 17 MeV which, although produced through hadronic couplings, only decays to  $e^-e^+$ ,  $v\bar{v}$  and  $\gamma\gamma\gamma$ . Therefore, it is expected that neutrinos should interact with matter by means of this interaction. It is also expected that there should be an interplay between this new interaction and the weak interaction because both interactions involve leptons. It is important at this point to comment on the Dava Bay results. This collaboration had reported in 2012 [10] a disappearance of about 6% in the electron antineutrino flux along a distance of 1648 m with respect to the produced flux according to the current models of nuclear theory, but in a paper that has just been published [21] the collaboration corrects the previous results because they have now found that the actual production is 6% larger than that one expected from the theoretical nuclear models, and thus, this invalidates their previous disappearance claims.

### 4. Tentative Solution to the Solar Neutrino Puzzle

Although we do not know much yet on this new interaction above discussed, which will be the subject of many experiments throughout the world in the near future, we can propose that the electron

neutrinos interact with matter in the Sun's core and that this is the reason for the disappearance of part of their flux produced in the pp and CNO cycles, and their disappearance in experiments with neutrino fluxes on Earth. Because of the interplay above mentioned between the new interaction and the weak interaction, we should also expect disappearance of a certain neutrino flavor due to the interactions in the medium and appearance of another flavor due to interaction in the same medium, as has been observed by T2K [22] and OPERA collaborations [16]. It is important to emphasize this important conclusion by OPERA: "We therefore claim the observation of a first candidate  $v_{\star}$  CC interaction. Its significance, based on our best conservative knowledge of the background, exceeds two  $\sigma$ . This does not allow yet claiming the observation of  $v_{\mu} \rightarrow v_{\tau}$ -oscillation. Given its sensitivity, the OPERA experiment will require the detection of a few more candidate events in order to firmly establish neutrino oscillations in direct appearance mode through the identification of the final charged lepton."

We make use below of the concept of particle transparency in a medium and use it for the electron neutrinos in the Sun's core. Extending the concept of transparency developed by Shapiro and Teukolsky [23] in their analysis of neutrino transparency in stars with core temperatures smaller than  $10^9 K$ , we can say that the effective mean free path of a neutrino in the Sun's core is given by

$$\lambda_{eff} \sim \left(\lambda_n \lambda_e \lambda_X\right)^{1/3}$$

in which  $\lambda_n$  and  $\lambda_e$  are the neutrino mean free paths due to elastic scattering off neutrons and the  $v_e - e^-$  weak inelastic scattering, respectively, and  $\lambda_x$  is the neutrino mean free path due to the new interaction.

Let us try to find an estimate for the new interaction cross section by making the following considerations. In order to diminish the electron neutrino flux to half in the Sun's core we should have  $\lambda_{eff} \sim R_c/-\ln 0.45 \approx 0.25R_{\odot}/-\ln 0.45 \approx 2.13 \times 10^8 m$  where  $R_c$  is the core radius of the Sun and  $R_{\odot}$  is the Sun's radius, assuming that the neutrino flux diminishes exponentially along the travelled distance. The product  $\lambda_e \lambda_n$  is given by the following relation [23]

$$\left(\lambda_e \lambda_n\right)^{1/2} = 2 \times 10^5 \, km \left(\frac{\rho_{nucl}}{\rho}\right)^{7/6} \left(\frac{10^5}{E_v}\right)^{5/2}$$

where  $E_{\nu}$  is the neutrino energy in electron-volt,  $\rho_{nucl} = 2.8 \times 10^{14} \text{ g/cm}^3$  is the nuclear density and  $\rho \approx 100 \text{ g/cm}^3$  is the Sun's core density. For <sup>8</sup>B neutrinos the energy spectrum [24] goes from zero up to 14 MeV, and the neutrino flux peaks around 7 MeV [25]. Thus, making  $E_{\nu} \approx 7$  MeV we obtain  $\lambda_e \lambda_n \approx 2.6 \times 10^{36} m^2$ , and from the relation  $\lambda_{eff} \sim (\lambda_n \lambda_e \lambda_x)^{1/3} = 2.13 \times 10^8 m$  we have then  $\lambda_x \approx 2.6 \times 10^{-12}$  m. The mean free path  $\lambda_x$  is related to the inelastic cross section  $\sigma_x$  due to the new interaction by the equation

$$\lambda_{X} = \frac{1}{\sigma_{X} n_{N}}$$

where  $n_N$  is the number of nucleons per unit volume. For the Sun's core the density of 100  $g/cm^3$  means  $6 \times 10^{25}$  nucleons/cm<sup>3</sup> and, hence, we obtain  $\sigma_X \approx 6.4 \times 10^{-19}$  cm<sup>2</sup>.

As  $\lambda_e = 1/(\sigma_e n_e)$  and  $\lambda_n = 1/(\sigma_n n_n)$  depend on the electron and neutron densities, we should take the above calculation with caution because the electron and neutron densities depend strongly on the distance from the core center [26], and thus,  $\lambda_x$  is an average number. A very important fact raised by Lopes and Turck-Chièze [26] is that different electron neutrinos are generated in different shells in the Sun's core, and thus, we cannot compare the different values for  $\lambda_x$  and  $\sigma_x$  from the neutrino data of different collaborations when they refer to different neutrinos, but because of the conclusion below we do not need to make such comparisons.

The problem with the above calculation is that  $\sigma_x \approx 6.4 \times 10^{-19} \text{ cm}^2$ is too large to be true. And as we saw above, although the new interaction is produced through hadronic couplings, its boson only decays to  $e^-e^+$ ,  $v\bar{v}$  and  $\gamma\gamma\gamma$ . This means that not only the neutrino, but also the electron would interact strongly by means of this new interaction, but this has not been seen for the electron. All this means that  $\sigma_x$  has to be much smaller than the above number. Reinforcing this reasoning it is worth recalling that the strong interaction between two protons has a cross section of about 40 mbarn, and thus,  $\sigma_x$ would be about 10000 times larger than the strong interaction cross section. Therefore, the solar neutrino problem may have only the other solution presented below.

# 4. A Possible Line of Solution to the Solar Neutrino Puzzle

In the light of the recent results of the Daya Bay Collaboration (21) and of the discovery of the new interaction above mentioned, we can propose that the actual production of electron neutrinos in the Sun's core is much larger than what has been predicted by the Standard Solar Model, exactly because of this new interaction. For example, if the production of <sup>8</sup>B electron neutrinos in the pp cycle were 55% larger than what is predicted by the Standard Solar Model, the solar neutrino puzzle would be solved. Therefore, it is urgent to investigate further this new interaction and find out the order of magnitude of its cross section for the different leptons. And this new interaction may also have an important role in the beginning of the Universe.

# 5. Conclusion

Taking into account the discovery of a new fundamental interaction of matter which involves neutrinos and the latest results of the Daya Bay Collaboration [21], I propose a completely different solution to the solar neutrino deficits which is more logical than the neutrino oscillation solution. My proposal is that more neutrinos are generated in the pp and CNO cycles in the Sun's core in such proportions that there are cancellations with the claimed neutrino deficits. We should investigate further the existence of this new fundamental interaction and the role that it plays in the pp and CNO cycles in the Sun's core for modifying the Solar Model accordingly. The above simple calculations provide a rationale for the definitive solution to the solar neutrino puzzle.

### REFERENCES

- S. Fukuda *et al.* (The Super-Kamiokande Collaboration), Solar 8B and hep Neutrino Measurements from 1258 Days of Super-Kamiokande Data, Phys. Rev. Lett. 86, 5651, 2001.
- 2. S. Fukuda *et al.* (The Super-Kamiokande Collaboration), Constraints on Neutrino Oscillations Using 1258 Days of Super-Kamiokande Solar Neutrino Data, Phys. Rev. Lett. **86**, 5656 (2001).

- **3.** Q. R. Ahmad *et al.* (SNO Collaboration), Direct Evidence for Neutrino Flavor Transformation from Neutral-Current Interactions in the Sudbury Neutrino Observatory, Phys. Rev. Lett. **89**, 011301 (2002).
- **4.** Q. R. Ahmad *et al.* (SNO Collaboration), Measurement of the Total Active 8B Solar Neutrino Flux at the Sudbury Neutrino Observatory with Enhanced Neutral Current Sensitivity, Phys. Rev. Lett. **92**, 181301 (2004).
- **5.** M. Altmann *et al.* (GNO Collaboratin), Complete results for five years of GNO solar neutrino observations, Phys. Lett. B, **616**, 174, (2005).
- 6. W. Hampel *et al.* (GALLEX Collaboratin), GALLEX solar neutrino observations: Results for GALLEX IV, Phys. Lett. B, **447**, 127 (1999).
- 7. C. Arpesella *et al.* (Borexino Collaboration), First real time detection of Be7 solar neutrinos by Borexino, Phys. Lett. B **658**,101 (2008).
- **8.** M. G. Aartsen *et al.* (IceCube Collaboration), Measurement of Atmospheric Neutrino Oscillations with IceCube, Phys. Rev. Lett. **111**, 081801 (2013).
- **9.** Y. Abe *et al.* (Double Chooz Collaboration), Reactor  $\overline{v_e}$  disappearance in the Double Chooz experiment, Phys. Rev. D **86**, 052008 (2012)
- **10.** F. P. An *et al.* (Daya Bay Collaboration), Observation of electron-antineutrino disappearance at Daya Bay, Phys. Rev. Lett. **108**, 171803 (2012).
- L. Wolfenstein, Neutrino oscillations in matter, Phys. Rev. D 17(9), 2369 (1978).
- **12.** L. Wolfenstein, Neutrino oscillations and stellar collapse, Phys. Rev. D **20** (10), 2634 (1979).
- 13. S. P. Mikheev and A. Yu. Smirnov, Resonance enhancement of oscillations in matter and solar neutrino spectroscopy, Soviet Journal of Nuclear Physics 42 (6): 913 (1985).
- **14.** Physicists Find First Direct Evidence for Tau neutrino at Fermilab (Press release), Fermilab 20 July 2000.
- 15. K. Kodama *et al.* (DONUT Collaboration), Observation of tau neutrino interactions, Phys. Lett. B, **504** (3), 218, 2001.
- 16. N. Agafonova *et al.* (OPERA Collaboration), Observation of a first tau neutrino candidate in the OPERA experiment in the CNGS beam, Phys. Lett. B 691, 138, 2010.
- **17.** D. G. Michael *et al.* (MINOS Collaboration), Observation of muon neutrino disappearance with the MINOS detectors in the NuMI neutrino beam, Phys. Rev.Lett. **97**, 19180, 2006.
- **18.** A. Franklin, *The Rise and Fall of the Fifth Force: Discovery, Pursuit, and Justification in Modern Physics*, American Institute of Physics, New York, 1993.
- **19.** A. Krasznahorkay, Observation of Anomalous Internal Pair Creation in <sup>8</sup>Be: A Possible Indication of a Light, Neutral Boson, Phys. Rev. Lett. **116**, 042501 (2016).
- 20. Jonathan L. Feng, Bartosz Fornal, Iftah Galon, Susan Gardner, Jordan Smolinsky, Tim M. P. Tait, and Philip Tanedo, Evidence for a Protophobic Fifth Force from 8Be Nuclear Transitions, arXiv:1604.07411 [hep-ph].

- 21. F. P. An *et al.* (Daya Bay Collaboration), Measurement of the Reactor Antineutrino Flux and Spectrum at Daya Bay, Phys. Rev. Lett. 116, 061801 (2016).
- **22.** K. Abe *et al.* (T2K Collaboration), Evidence of electron neutrino appearance in a muon neutrino beam, Phys. Rev. D **88**, 032002 (2013)
- 23. S. L. Shapiro and S. A. Teukolsky, Black Holes, White Dwarfs and Neutron Stars, John Wiley & Sons, New York (1983), pp. 26, 27.
- 24. J. Bahcall, E. Lisi, D. E. Alberger, L. De Braeckleer, S. J. Freedman, and J. Napolitano, Standard Neutrino Spectrum from <sup>8</sup>B Decay, Phys. Rev. C 54, 411 (1996).
- 25. J. N. Bahcall, Where We Are, Where We Are Going, Ap. J. 467, 475, 1996.
- **26.** I. Lopes, S. Turck-Chièze, Solar neutrino physics oscillations: Sensitivity to the electronic density in the Sun's core, Ap. J. **765**(1), 14, 2013.